

High temperature limit of the Standard Model due to gauge groups contraction

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The aim of my talk is to **construct high temperature limit** of the Standard Model and on this base **describe evolution of the particles** in the early Universe.

The Standard Model

- ▶ The modern knowledge of the particle world is concentrated in Standard Model (SM). This theory consist of two parts: Electroweak Model (EWM), which unified electromagnetic and weak interactions, and Quantum Chromodynamics (QCD), describing their strong interactions.
- ▶ The Electroweak Model is a gauge theory with the gauge group $SU(2) \times U(1)$, which act in the boson, lepton and quark sectors.
- ▶ QCD is $SU(3)$ gauge theory based on the local color degrees of freedom of quarks.

Elementary particles content of SM

Gauge bosons:

γ (photon),

W^\pm (charged weak bosons), Z^0 (neutral weak boson),

A^k , $k = 1, \dots, 8$ (gluons).

Special particle:

χ (Higgs boson).

Leptons:

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix}, \quad \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}, \quad \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix} \in \mathbb{C}_2.$$

Quarks:

$$\begin{pmatrix} u \\ d \end{pmatrix}, \quad \begin{pmatrix} c \\ s \end{pmatrix}, \quad \begin{pmatrix} t \\ b \end{pmatrix} \in \mathbb{C}_2.$$

Electroweak Model

The Lagrangian of the model is given by the sum of the boson L_B , of the lepton L_L and of the quark L_Q Lagrangians:

$$L = L_B + L_L + L_Q.$$

It is invariant under the action of the gauge group $SU(2) \times U(1)$ in the 2-dim. complex space \mathbb{C}_2 :

$$SU(2) : \vec{z}' = G\vec{z},$$

$$\begin{pmatrix} z'_1 \\ z'_2 \end{pmatrix} = \begin{pmatrix} \alpha & \beta \\ -\bar{\beta} & \bar{\alpha} \end{pmatrix} \begin{pmatrix} z_1 \\ z_2 \end{pmatrix}, \quad |\alpha|^2 + |\beta|^2 = 1,$$

$$U(1) : \vec{z}' = e^{i\omega} \vec{z}, \quad \omega \in \mathbf{R}.$$

Boson sector $L_B = L_A + L_\phi$ involve two parts:
the gauge field Lagrangian

$$L_A = -\frac{1}{4}[(F_{\mu\nu}^1)^2 + (F_{\mu\nu}^2)^2 + (F_{\mu\nu}^3)^2] - \frac{1}{4}(B_{\mu\nu})^2,$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu + [A_\mu, A_\nu], \quad B_{\mu\nu} = \partial_\mu B_\nu - \partial_\nu B_\mu$$

and **the matter field Lagrangian**

$$L_\phi = \frac{1}{2}(D_\mu\phi)^\dagger D_\mu\phi - \frac{\lambda}{4}(\phi^\dagger\phi - v^2)^2,$$

where $\phi = \begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix} \in \mathbb{C}_2$ are the matter fields.

Fermion sector is represented by **the lepton** L_L and **quark** L_Q Lagrangians.

For the first generation **the lepton** Lagrangian is taken in the form:

$$L_{L,e} = L_l^\dagger i \tilde{\tau}_\mu D_\mu L_l + e_r^\dagger i \tau_\mu D_\mu e_r - h_e [e_r^\dagger (\phi^\dagger L_l) + (L_l^\dagger \phi) e_r],$$

where $L_l = \begin{pmatrix} \nu_l \\ e_l \end{pmatrix} \in \mathbb{C}_2$ is the $SU(2)$ -doublet, e_r is the $SU(2)$ -singlet, h_e is constant, $\tau_0 = \tilde{\tau}_0 = \mathbf{1}$, $\tilde{\tau}_k = -\tau_k$, τ_μ are Pauli matrices and e_r, e_l, ν_l are the two component Lorentz spinors. First and second terms in $L_{L,e}$ describe free movement of left and right fermions and their interactions with gauge fields. Last term corresponds to the electron mass.

The quark Lagrangian is constructed in a similar way

$$L_Q = Q_l^\dagger i \tilde{\tau}_\mu D_\mu Q_l + u_r^\dagger i \tau_\mu D_\mu u_r + d_r^\dagger i \tau_\mu D_\mu d_r - \\ - h_d [d_r^\dagger (\phi^\dagger Q_l) + (Q_l^\dagger \phi) d_r] - h_u [u_r^\dagger (\tilde{\phi}^\dagger Q_l) + (Q_l^\dagger \tilde{\phi}) u_r],$$

where left quark fields form the $SU(2)$ -doublet

$Q_l = \begin{pmatrix} u_l \\ d_l \end{pmatrix} \in \mathbb{C}_2$, **right quark fields u_r, d_r are the**

$SU(2)$ -singlets, $\tilde{\phi}_i = \epsilon_{ik} \bar{\phi}_k$, $\epsilon_{00} = 1$, $\epsilon_{ii} = -1$ is the conjugate representation of $SU(2)$ group and h_u, h_d are constants. All fields u_l, d_l, u_r, d_r are two component Lorentz spinors.

The new gauge fields are introduced

$$W_{\mu}^{\pm} = \frac{1}{\sqrt{2}} (A_{\mu}^1 \mp iA_{\mu}^2), \quad Z_{\mu} = \frac{1}{\sqrt{g^2 + g'^2}} (gA_{\mu}^3 - g'B_{\mu}),$$

$$A_{\mu} = \frac{1}{\sqrt{g^2 + g'^2}} (g'A_{\mu}^3 + gB_{\mu}).$$

They are expressed through the old fields

$$A_{\mu}(x) = -ig \sum_{k=1}^3 T_k A_{\mu}^k(x), \quad B_{\mu}(x) = -ig' B_{\mu}(x).$$

Chromodynamics

- ▶ The QCD gauge group is $SU(3)$, acting in 3-dim. complex space \mathbb{C}_3 of color quark states

$$q = \begin{pmatrix} q_1 \\ q_2 \\ q_3 \end{pmatrix} \equiv \begin{pmatrix} q_R \\ q_G \\ q_B \end{pmatrix} \in \mathbb{C}_3,$$

where $q(x)$ are quark fields $q = u, d, s, c, b, t$ and R (red), G (green), B (blue) are color degrees of freedom.

- ▶ The $SU(3)$ gauge bosons are called **gluons**.

- ▶ QCD Lagrangian is taken in the form

$$\mathcal{L} = \sum_q \bar{q}^i (i\gamma^\mu) (D_\mu)_{ij} q^j - \frac{1}{4} \sum_{\alpha=1}^8 F_{\mu\nu}^\alpha F^{\mu\nu\alpha},$$

- ▶ where $D_\mu q$ are covariant derivatives of quark fields

$$D_\mu q = \left(\partial_\mu - ig_s \left(\frac{\lambda^a}{2} \right) A_\mu^a \right) q,$$

g_s is the strong coupling constant, $t^a = \lambda^a/2$ are generators of $SU(3)$, λ^a — Gell-Mann matrices

- ▶ and gluon stress tensor has the form

$$F_{\mu\nu}^\alpha = \partial_\mu A_\nu^\alpha - \partial_\nu A_\mu^\alpha + g_s f^{\alpha\beta\gamma} A_\mu^\beta A_\nu^\gamma.$$

Contraction of the Electroweak Model

- ▶ The **infinite energy limit** of EWM corresponds to the **consistent rescaling** of $SU(2)$ and \mathbb{C}_2

$$\begin{pmatrix} z'_1 \\ \epsilon z'_2 \end{pmatrix} = \begin{pmatrix} \alpha & \epsilon\beta \\ -\epsilon\bar{\beta} & \bar{\alpha} \end{pmatrix} \begin{pmatrix} z_1 \\ \epsilon z_2 \end{pmatrix}, \quad |\alpha|^2 + \epsilon^2|\beta|^2 = 1,$$

when contraction parameter tends to zero $\epsilon \rightarrow 0$.

- ▶ Substitution $\beta \rightarrow \epsilon\beta$ induces another ones for the gauge fields:

$$W_\mu^\pm \rightarrow \epsilon W_\mu^\pm, \quad Z_\mu \rightarrow Z_\mu, \quad A_\mu \rightarrow A_\mu.$$

and substitution $z_2 \rightarrow \epsilon z_2$ induces the following transformation of lepton and quark fields:

$$e_l \rightarrow \epsilon e_l, \quad d_l \rightarrow \epsilon d_l, \quad \nu_l \rightarrow \nu_l, \quad u_l \rightarrow u_l.$$

- ▶ The next reason is the special mechanism of spontaneous symmetry breaking, which is used to generate mass for the vector bosons. One of L_B ground states

$$\phi^{vac} = \begin{pmatrix} 0 \\ v \end{pmatrix}, \quad A_\mu^k = B_\mu = 0$$

is taken as vacuum and then small field excitations $v + \chi(x)$ with respect to this vacuum are regarded.

- ▶ So, Higgs boson field χ , constant v and particle masses m_p , which depend on v , are multiplied by contraction parameter:

$$\chi \rightarrow \epsilon\chi, \quad v \rightarrow \epsilon v, \quad m_p \rightarrow \epsilon m_p, \quad p = \chi, W, Z, e, u, d.$$

- ▶ After these transformations the Lagrangian can be represented in the form

$$L(\epsilon) = L_\infty - \epsilon m_u (u_r^\dagger u_l + u_l^\dagger u_r) + \\ + \epsilon^2 L_2 + \epsilon^3 L_3 + \epsilon^4 L_4.$$

For $\epsilon \rightarrow 0$ the terms with the higher powers of ϵ contribute less than the terms with lower powers.

- ▶ So the Electroweak Model demonstrates **five stages of behaviour** in the infinite energy limit, which are distinguished by the powers of the contraction parameter.

QCD with contracted gauge group

The contracted group $SU(3; \epsilon)$ is defined by the action

$$q'(\epsilon) = \begin{pmatrix} q'_1 \\ \epsilon q'_2 \\ \epsilon^2 q'_3 \end{pmatrix} = \begin{pmatrix} u_{11} & \epsilon u_{12} & \epsilon^2 u_{13} \\ \epsilon u_{21} & u_{22} & \epsilon u_{23} \\ \epsilon^2 u_{31} & \epsilon u_{32} & u_{33} \end{pmatrix} \begin{pmatrix} q_1 \\ \epsilon q_2 \\ \epsilon^2 q_3 \end{pmatrix} = U(\epsilon)q(\epsilon)$$

on the color space $\mathbb{C}_3(\epsilon)$.

The quark and gluon fields are transformed as follows:

$$q_1 \rightarrow q_1, \quad q_2 \rightarrow \epsilon q_2, \quad q_3 \rightarrow \epsilon^2 q_3,$$

$$A_\mu^{GR} \rightarrow \epsilon A_\mu^{GR}, \quad A_\mu^{BG} \rightarrow \epsilon A_\mu^{BG}, \quad A_\mu^{BR} \rightarrow \epsilon^2 A_\mu^{BR},$$

and diagonal gauge fields are not changed

$$A_\mu^{RR} \rightarrow A_\mu^{RR}, \quad A_\mu^{GG} \rightarrow A_\mu^{GG}, \quad A_\mu^{BB} \rightarrow A_\mu^{BB}.$$

- ▶ With this substitution, we obtain the quark part of QCD Lagrangian in the form

$$\begin{aligned}
 \mathcal{L}_q(\kappa) = & \sum_q i\bar{q}_1\gamma^\mu\partial_\mu q_1 + \frac{g_s}{2}|q_1|^2\gamma^\mu A_\mu^{RR} + \\
 & + \epsilon^2 \left\{ i\bar{q}_2\gamma^\mu\partial_\mu q_2 + \frac{g_s}{2} \left(|q_2|^2\gamma^\mu A_\mu^{GG} + q_1\bar{q}_2\gamma^\mu A_\mu^{GR} + \bar{q}_1q_2\gamma^\mu \bar{A}_\mu^{GR} \right) \right\} + \\
 & + \epsilon^4 \left[i\bar{q}_3\gamma^\mu\partial_\mu q_3 + \frac{g_s}{2} \left(|q_3|^2\gamma^\mu A_\mu^{BB} + q_1\bar{q}_3\gamma^\mu A_\mu^{BR} + \bar{q}_1q_3\gamma^\mu \bar{A}_\mu^{BR} + \right. \right. \\
 & \left. \left. + q_2\bar{q}_3\gamma^\mu A_\mu^{BG} + \bar{q}_2q_3\gamma^\mu \bar{A}_\mu^{BG} \right) \right] = L_q^\infty + \epsilon^2 L_q^{(2)} + \epsilon^4 L_q^{(4)}.
 \end{aligned}$$

- ▶ Gluon part $L_{gl} = -\frac{1}{4}F_{\mu\nu}^\alpha F^{\mu\nu\alpha}$ of Lagrangian is very cumbersome, therefore we omit its general form.

- ▶ The contracted QCD Lagrangian has the form

$$\mathcal{L}(\epsilon) = L^\infty + \epsilon^2 L^{(2)} + \epsilon^4 L^{(4)} + \epsilon^6 L^{(6)} + \epsilon^8 L^{(8)},$$

with the explicit expressions for each $L^{(k)}$.

- ▶ The contraction parameter is monotonous function of the average energy E (or temperature T) with the property $\epsilon(E) \rightarrow 0$ for $E \rightarrow \infty$.

Very higher energies (temperatures) can exist in the early Universe after inflation and reheating on the first stages of the Hot Big Bang ($1eV = 10^4 K$).

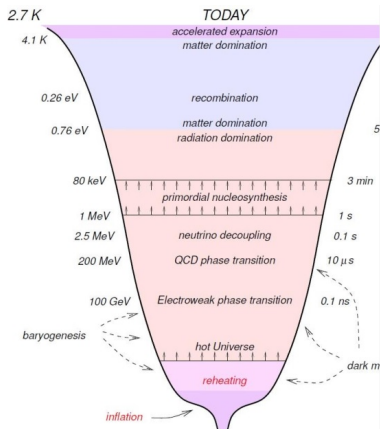


Рис.: History of the Universe (V. Rubakov, D. Gorbunov, INR RAS)

Estimation of boundary values

- ▶ To estimate the absolute dates of different stages of SM development we use the fact that the electroweak epoch starts at the temperature $T_4 = 100 \text{ GeV}$ ($1 \text{ GeV} = 10^{13} \text{ K}$) and the QCD epoch begins at $T_8 = 0,2 \text{ GeV}$,
- ▶ i.e. we assume that complete reconstruction of the EWM, whose Lagrangian has minimal terms $\approx \epsilon^4$, and QCD with minimal terms $\approx \epsilon^8$, take place at these temperatures.

- ▶ Let us denote by Δ cutoff level for ϵ^k , $k = 1, 2, 4, 6, 8$, i.e. for $\epsilon^k < \Delta$ all the terms proportionate to ϵ^k are negligible quantities in Lagrangian.
- ▶ At last we suppose that the contraction parameter inversely depends on temperature

$$\epsilon(T) = \frac{A}{T}, \quad A = \text{const.} \quad (1)$$

- ▶ From the equation for QCD $\epsilon^8(T_8) = A^8 T_8^{-8} = \Delta$ we obtain $A = T_8 \Delta^{1/8} = 0,2 \Delta^{1/8} \text{ GeV}$.

From the similar equation for EWM we obtain the cutoff level $\Delta = (T_8 E_4^{-1})^8 = (0,2 \cdot 10^{-2})^8 \approx 10^{-22}$.

- ▶ From the equation for k -th power $\epsilon^k(T_k) = A^k T_k^{-k} = \Delta$ we have

$$T_k = T_8 \Delta^{\frac{k-8}{8k}} \approx 10^{\frac{88-15k}{4k}} \text{ GeV}$$

and easily find the boundary values (GeV):

$$T_1 = 10^{18}, T_2 = 10^7, T_3 = 10^3, T_4 = 10^2, T_6 = 1, T_8 = 2 \cdot 10^{-1}$$

- ▶ **The estimation for "infinity" temperature $T_1 \approx 10^{18} \text{ GeV}$ is comparable with Planck energy $\approx 10^{19} \text{ GeV}$, where the gravitation effects are important.**
- ▶ So the developed evolution of the elementary particles does not exceed the range of the problems described by electroweak and strong interactions.

Evolution of elementary particles

$$(T > 10^{18} \text{ GeV}).$$

- ▶ At the infinite temperature ($\epsilon = 0$) the EWM Lagrangian is as follows

$$L_\infty = -\frac{1}{4}\mathcal{Z}_{\mu\nu}^2 - \frac{1}{4}\mathcal{F}_{\mu\nu}^2 + \nu_l^\dagger i\tilde{\tau}_\mu \partial_\mu \nu_l + u_l^\dagger i\tilde{\tau}_\mu \partial_\mu u_l + \\ + e_r^\dagger i\tau_\mu \partial_\mu e_r + d_r^\dagger i\tau_\mu \partial_\mu d_r + u_r^\dagger i\tau_\mu \partial_\mu u_r + L_\infty^{\text{int}}(A_\mu, Z_\mu).$$

- ▶ So the Electroweak Model includes only **massless particles**: photons A_μ and neutral bosons Z_μ , left quarks u_l and neutrinos ν_l , right electrons e_r and quarks u_r, d_r .
- ▶ **The electroweak interactions become long-range** because they are mediated by the massless Z -bosons and photons.

- ▶ From the explicit form of the interaction part

$$\begin{aligned}
 L_{\infty}^{int}(A_{\mu}, Z_{\mu}) = & \frac{g}{2 \cos \theta_w} \nu_l^{\dagger} \tilde{\tau}_{\mu} Z_{\mu} \nu_l + \frac{2e}{3} u_l^{\dagger} \tilde{\tau}_{\mu} A_{\mu} u_l + \\
 & + \frac{g}{\cos \theta_w} \left(\frac{1}{2} - \frac{2}{3} \sin^2 \theta_w \right) u_l^{\dagger} \tilde{\tau}_{\mu} Z_{\mu} u_l + g' \sin \theta_w e_r^{\dagger} \tau_{\mu} Z_{\mu} e_r - \\
 & - g' \cos \theta_w e_r^{\dagger} \tau_{\mu} A_{\mu} e_r - \frac{1}{3} g' \cos \theta_w d_r^{\dagger} \tau_{\mu} A_{\mu} d_r + \frac{1}{3} g' \sin \theta_w d_r^{\dagger} \tau_{\mu} Z_{\mu} d_r + \\
 & + \frac{2}{3} g' \cos \theta_w u_r^{\dagger} \tau_{\mu} A_{\mu} u_r - \frac{2}{3} g' \sin \theta_w u_r^{\dagger} \tau_{\mu} Z_{\mu} u_r
 \end{aligned}$$

it follows that **there are no interactions between particles of different kind**, for example neutrinos interact only with each other by neutral currents.

- ▶ It looks like some **stratification of the Electroweak Model** with only one sort of particles in each stratum.

- ▶ From the limit QCD Lagrangian

$$\mathcal{L}_\infty = L_q^\infty + L_{gl}^\infty = \sum_q i\bar{q}_1 \gamma^\mu \partial_\mu q_1 + \frac{g_s}{2} |q_1|^2 \gamma^\mu A_\mu^{RR} - \frac{1}{4} (F_{\mu\nu}^{RR})^2 - \frac{1}{4} (F_{\mu\nu}^{GG})^2 - \frac{1}{4} F_{\mu\nu}^{RR} F_{\mu\nu}^{GG}.$$

it follows that only dynamic terms for the first color components of massless quarks survive under infinite temperature, which means that **quarks are monochromatic**.

- ▶ The terms also survive, which describe the interactions of these components with R -gluons. So **the stratification is conserved in the QCD sector**.

Evolution of elementary particles

$$(\approx \epsilon, \quad 10^{18} \text{ GeV} \geq T > 10^7 \text{ GeV}).$$

- ▶ The mass term of u -quark in the complete Lagrangian $L(\epsilon)$ is proportional to ϵ

$$\epsilon m_u (u_r^\dagger u_l + u_l^\dagger u_r).$$

The same is held for c - and t -quark. So u -, c - and t -quark first restores its mass in the evolution of the Universe.

Evolution of elementary particles

$$(\approx \epsilon^2, \quad 10^7 \text{ GeV} \geq T > 10^3 \text{ GeV}).$$

- ▶ The mass terms of electron and d -quark are multiplied by ϵ^2

$$\epsilon^2 \left[m_e (e_r^\dagger e_l + e_l^\dagger e_r) + m_d (d_r^\dagger d_l + d_l^\dagger d_r) \right],$$

The same is true for μ^- and τ -lepton, for s - and b -quark.
These particles become massive in the second stage.

- ▶ The quarks obtain the second color degree of freedom.
- ▶ The main part of electroweak and color interactions are restored in this epoch.

Evolution of elementary particles

$$(\approx \epsilon^3, \quad 10^3 \text{ GeV} \geq T > 10^2 \text{ GeV}).$$

There is one term in Lagrangian $L_3 = gW_\mu^+ W_\mu^- \chi$ proportionate to ϵ^3 , which describe Higgs boson interaction with charged W -bosons.

$$(\approx \epsilon^4, \quad 10^2 \text{ GeV} \geq T > 1 \text{ GeV}).$$

- ▶ Higgs boson χ and charged W -boson last restore their masses after all other particles of SM.
- ▶ The final reconstruction of the EWM takes place in this epoch.
- ▶ The quarks obtain the third color degree of freedom.

Evolution of elementary particles

$$(\approx \epsilon^6, \quad 1 \text{ GeV} \geq T > 2 \cdot 10^{-1} \text{ GeV}).$$

Next part of color interactions is restored.

$$(\approx \epsilon^8, \quad T \leq 2 \cdot 10^{-1} \text{ GeV}).$$

- ▶ The color interactions are completely valid.
- ▶ **The QCD is fully reconstructed** .
- ▶ **Start the time of the Standard Model.**

Conclusion

On the base of the modern knowledge of the particle world, which is concentrated in SM, **the hypothesis of the particles evolution in the early Universe is offered,**

The exact Lagrangians for any stage of evolution **takes away the hierarchy problem of the SM.**

- ▶ more details:

N.G., Elementary particles in the early Universe,
**Journal of Cosmology and Astroparticle
Physics 03(2016) 053.**

- ▶ Thank you for attention.