

# **Search for physics beyond the SM at KamLAND**

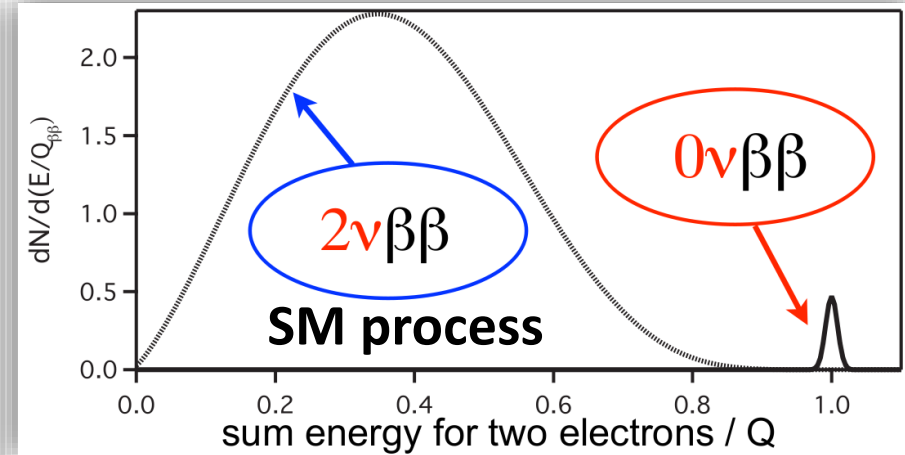
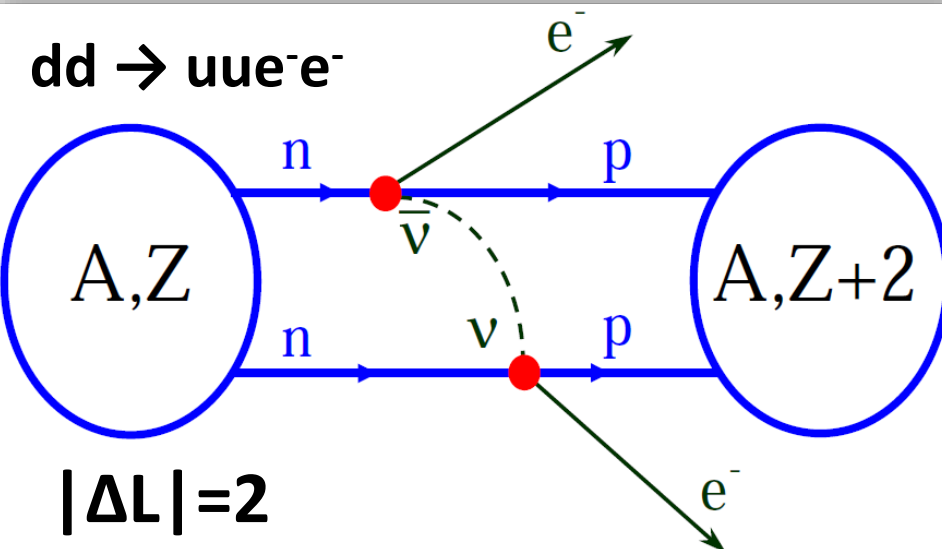
**The University of Tokyo  
Alexandre Kozlov**

**The 3<sup>rd</sup> international conference on particle physics and astrophysics  
Moscow, October 2017**

# Outline

- ❑ Neutrinoless double beta decay search using  $^{136}\text{Xe}$ -loaded liquid scintillator located at the ultra-low background central region of the KamLAND.
- ❑ Development of ultra-low background **NaI(Tl)** detectors and related infrastructure for rare event searches.

# The $0\nu\beta\beta$ test of seesaw mechanism



Test of the **Leptogenesis** (Fukugita & Yanagida) as explanation for **baryon asymmetry of the Universe**

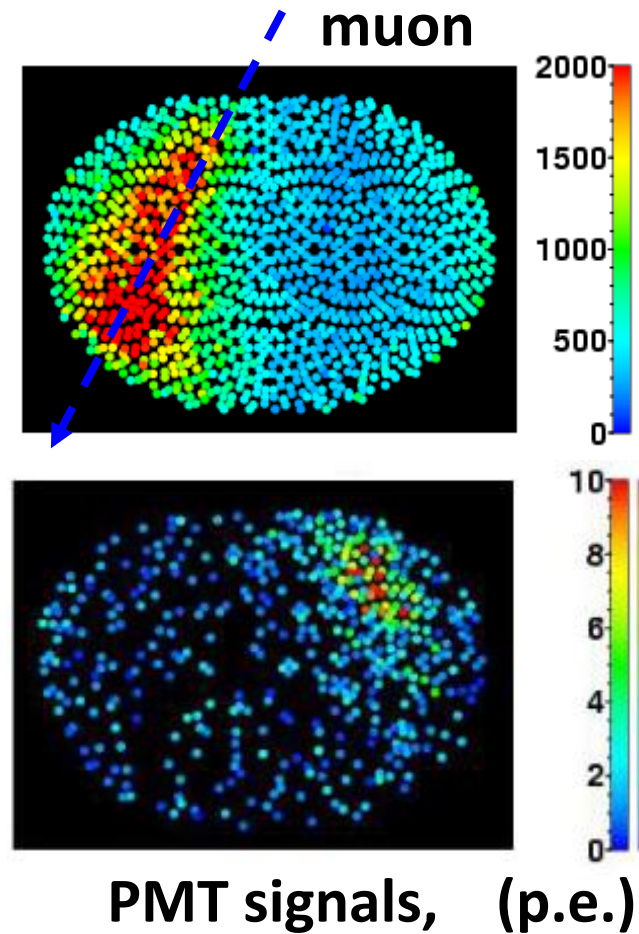
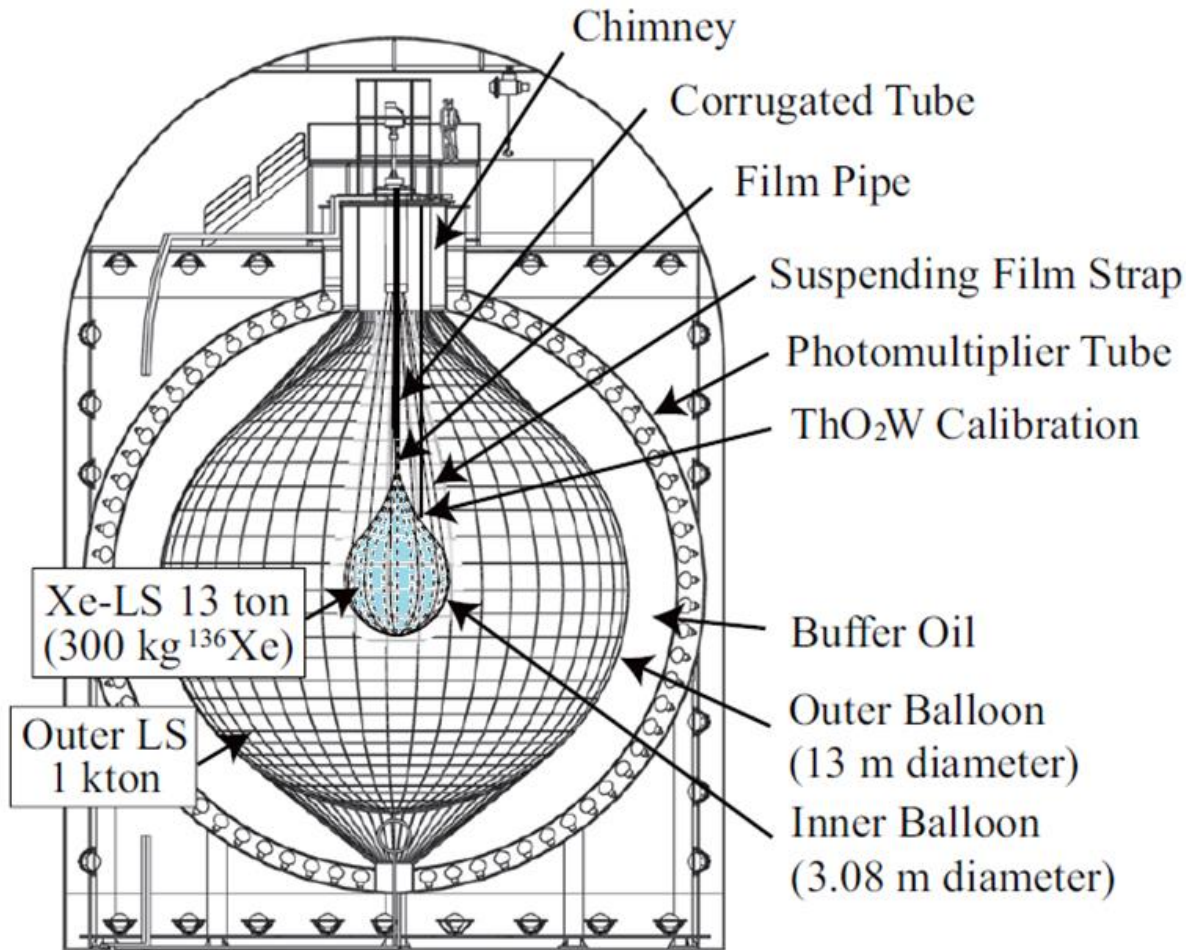
$$(T_{\frac{1}{2}}^{0\nu})^{-1} = G^{0\nu} \cdot g_A^4 \cdot |M^{0\nu}|^2 \cdot m_{\beta\beta}^2$$

- $G^{0\nu}(Q_{\beta\beta}, Z)$  – phase space factor
- $|M^{0\nu}|$  – nuclear matrix elements
- $m_{\beta\beta}$  – effective mass of neutrino

In calorimeters, as KamLAND, **sum of kinetic energies** of two electrons is measured

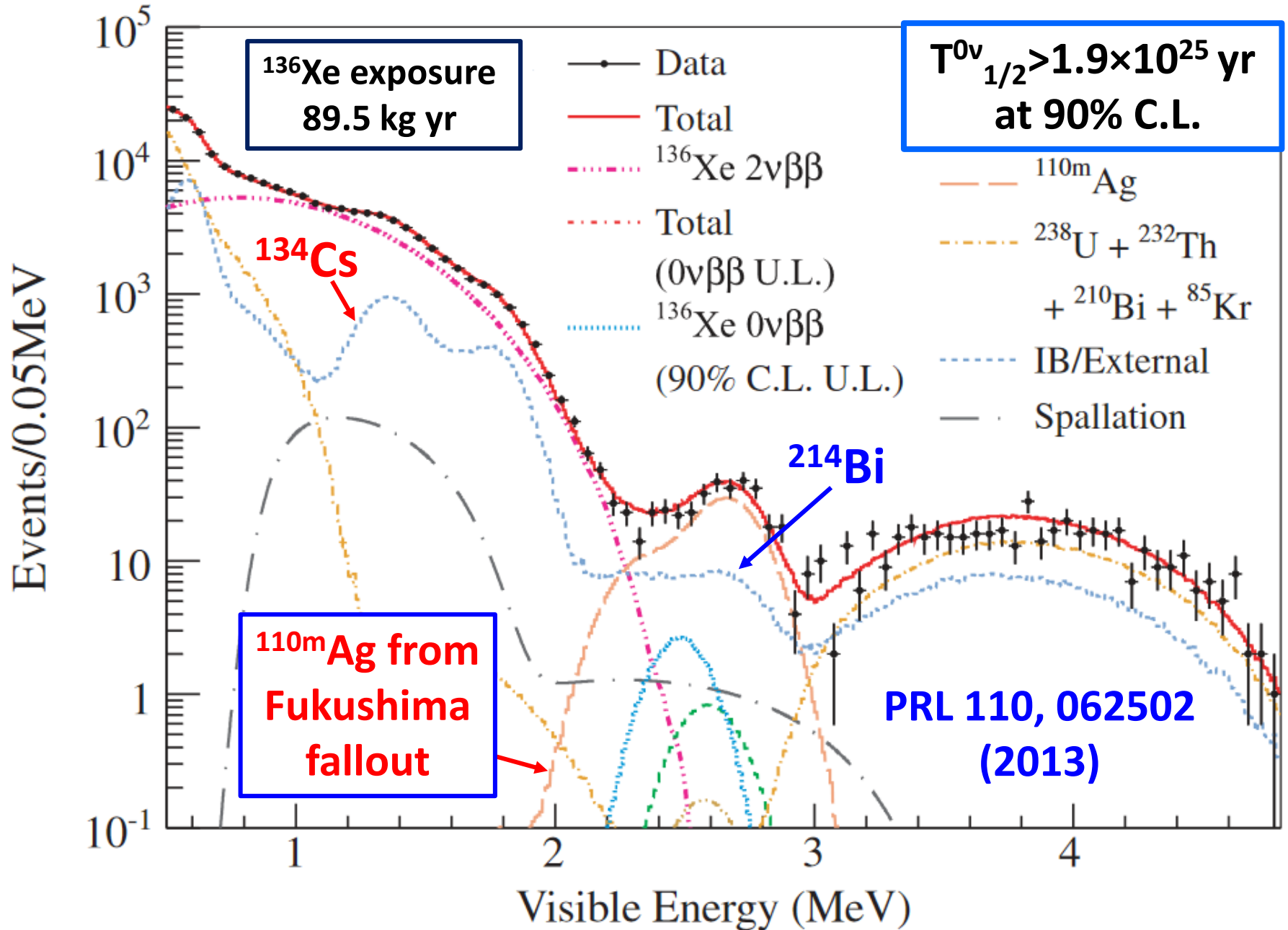
- ❑ The **only** method to measure the **absolute neutrino mass** below quasi-degenerate region.
- ❑ **Neutrino sector** is the **only place** where **physics beyond SM** was observed.

# Structure of the KamLAND-Zen (year 2011)

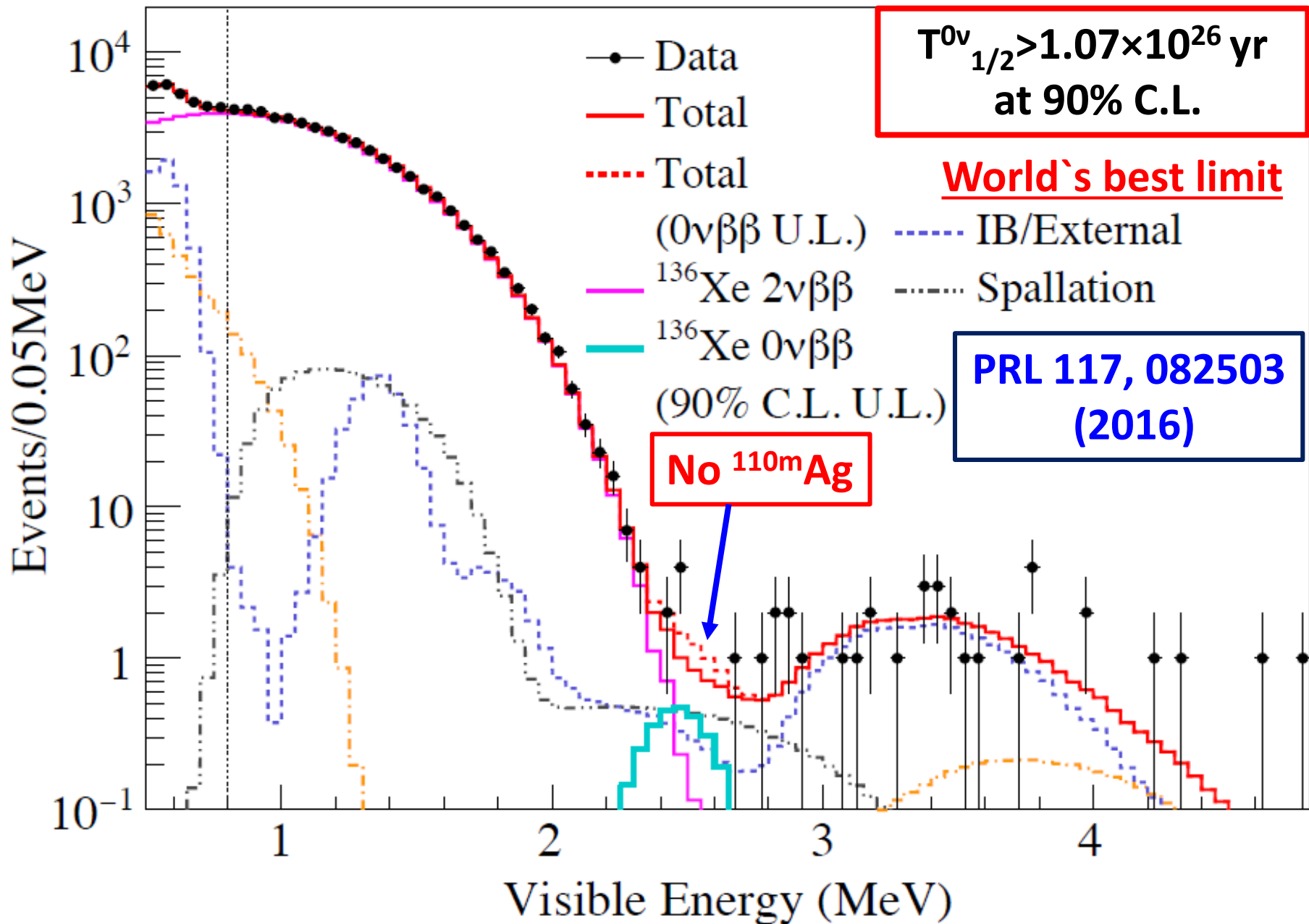


- In 2011, enriched <sup>136</sup>Xe(2.5-3wt%) + Liquid Scintillator in a ø3.08m mini-balloon made of a 25µm-thick Nylon film was deployed at KamLAND.
- It exploits the KamLAND detector **radio-purity, light sensors** (1879 PMTs 17&20-inch) and **data acquisition system**.

# KamLAND-Zen 400: Phase I (year 2012)



# KamLAND-Zen 400: final result (year 2016)



# Joint work with F. Simkovic to determine $g_A$ using shape of the measured $2\nu\beta\beta$ spectrum

$$\left[T_{1/2}^{2\nu\beta\beta}\right]^{-1} \simeq \left(g_A^{\text{eff}}\right)^4 \left|M_{GT-3}^{2\nu}\right|^2 \frac{1}{|\xi_{13}^{2\nu}|^2} \left(G_0^{2\nu} + \xi_{13}^{2\nu} G_2^{2\nu}\right)$$

$$\left(g_A^{\text{eff}}\right)^2 = \frac{1}{\left|M_{GT-3}^{2\nu}\right|} \frac{|\xi_{13}^{2\nu}|}{\sqrt{T_{1/2}^{2\nu\text{-exp}} \left(G_0^{2\nu} + \xi_{13}^{2\nu} G_2^{2\nu}\right)}}$$

Can be calculated exactly

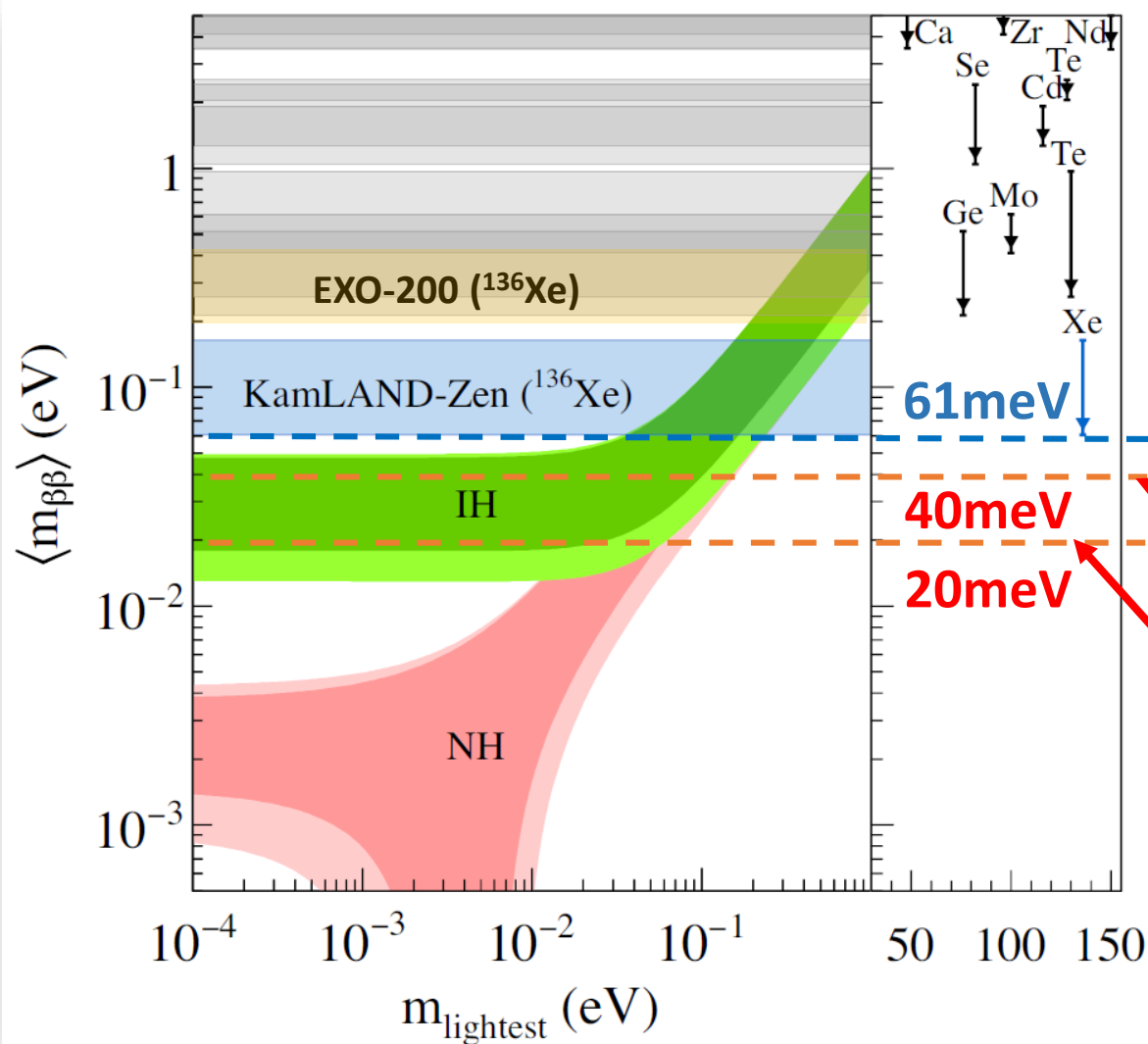
Extracted from the  $2\nu\beta\beta$  spectrum shape

$$M_{GT-1}^{2\nu} \equiv M_{GT}^{2\nu}$$

$$M_{GT-3}^{2\nu} = \sum_n M_n \frac{4 m_e^3}{(E_n - (E_i + E_f)/2)^3} \quad \xi_{13}^{2\nu} = \frac{M_{GT-3}^{2\nu}}{M_{GT-1}^{2\nu}}$$

The  $g_A^{\text{eff}}$  can be determined with the measured half-life, ratio of NMEs and calculated NME dominated by transitions through low lying states of the intermediate nucleus.

# First test of the IH mass region with KL-Zen 800



**KL-Zen 400 final**

**KL-Zen 800 (expected)**  
First result in 1.5-2 yrs

**KL2-Zen (expected)**  
Upgrade & data taking start  
time depend on funding

Prediction for the  $m_{\beta\beta} = 47 \pm 1 \text{ meV}$ : K. Harigaya, M. Ibe, and T. Yanagida  
"Seesaw mechanism with Occam's razor" PRD 86, 013002 (2012)



# The KamLAND-Zen summary

## Accomplished:

- ❑ **The KamLAND-Zen 400** was completed in Oct 2015. Enriched xenon was extracted from liquid scintillator, purified by distillation, and returned back to a storage facility.
- ❑ We published **world's best limit:  $T_{1/2}^{0\nu} > 1.07 \times 10^{26}$  yr at 90% C.L.** ( $m_{\beta\beta}$  is **61-165meV** depending on choice of NME).

## Ongoing:

- ❑ **A new mini-balloon** for 800kg of  $^{136}\text{Xe}$  will be deployed into KamLAND this year.
- ❑ During the **800kg** phase we may test **Yanagida's prediction** for  **$m_{\beta\beta} = 47 \pm 1 \text{meV}$** .

## Future:

- ❑ We work on future **KamLAND2-Zen** project to cover most of the **IH mass region** down to  **$m_{\beta\beta} = 20 \text{meV}$** .

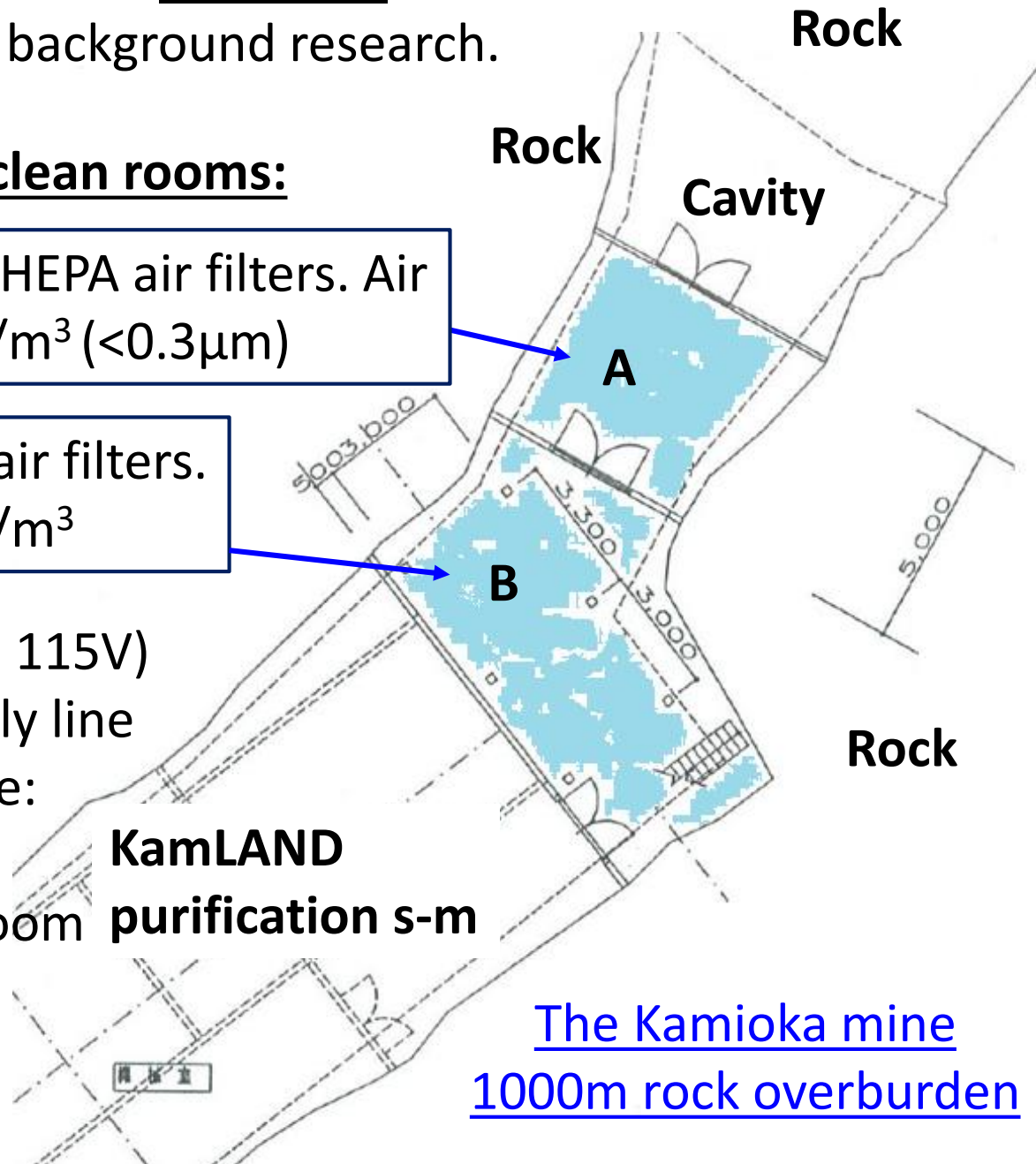
In 2012, I began development of two clean rooms (A, B) for ultra-low background research.

**Current conditions at the clean rooms:**

**Room A:** 60m<sup>3</sup>/min ULPA/HEPA air filters. Air quality: 100-300 particles/m<sup>3</sup> (<0.3μm)

**Room B:** 70m<sup>3</sup>/min HEPA air filters. Air quality: 2000 particles/m<sup>3</sup>

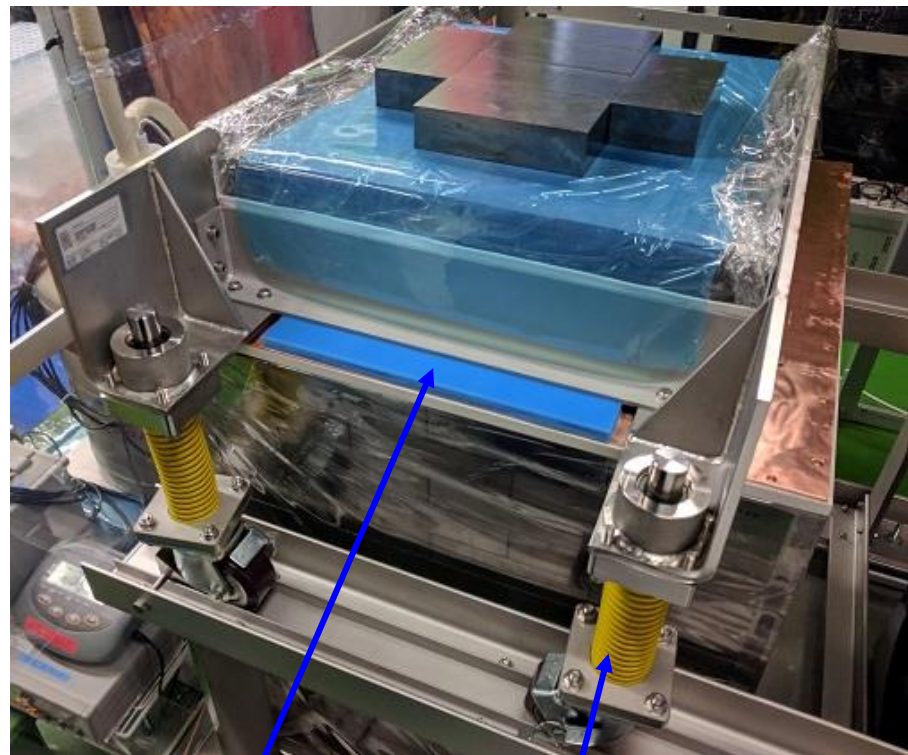
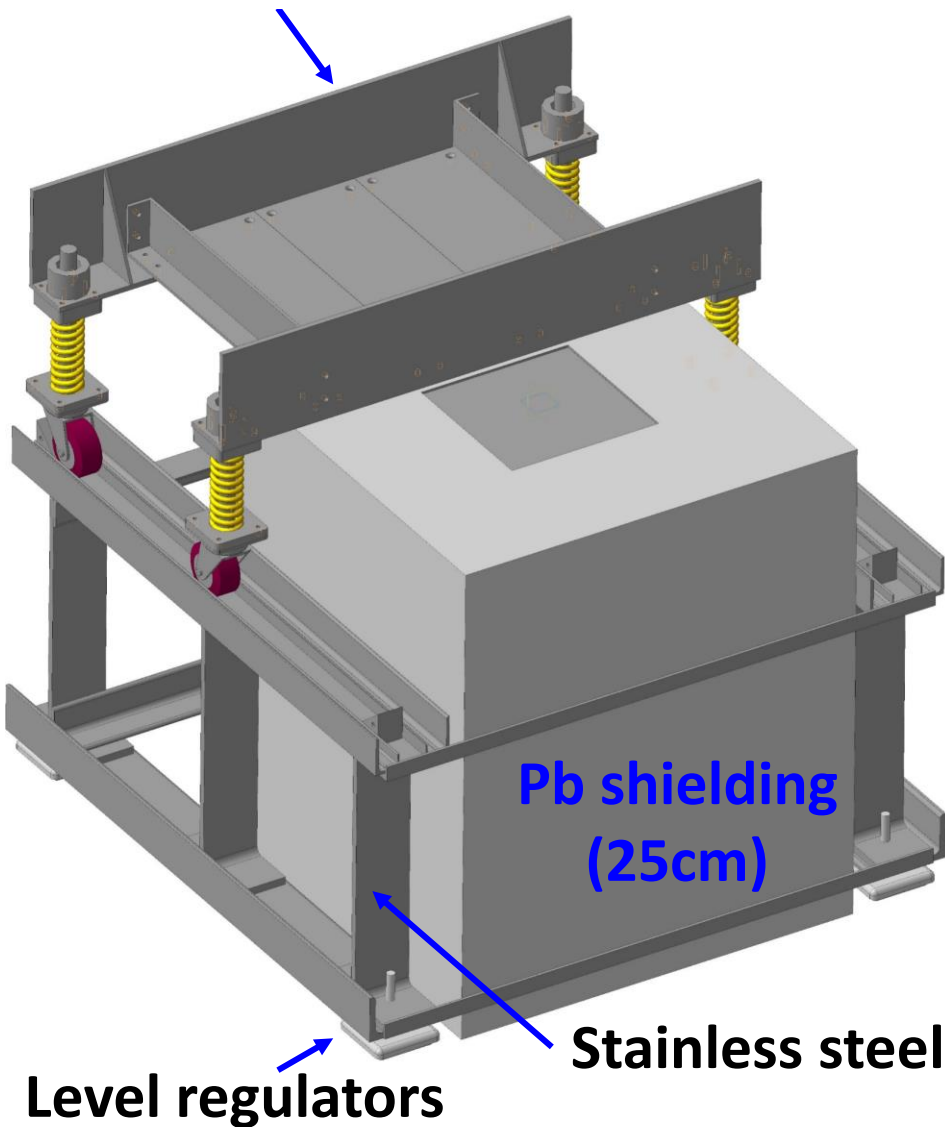
- 17kWatt AVR unit (100V, 115V)
- Boiled off Nitrogen supply line
- Radon-less air supply line: (5-10m<sup>3</sup> per hour)
- Air cond. units in each room



The Kamioka mine  
1000m rock overburden

# Our HPGe detector shielding design

**Cart for Pb bricks made of radio-pure stainless steel**



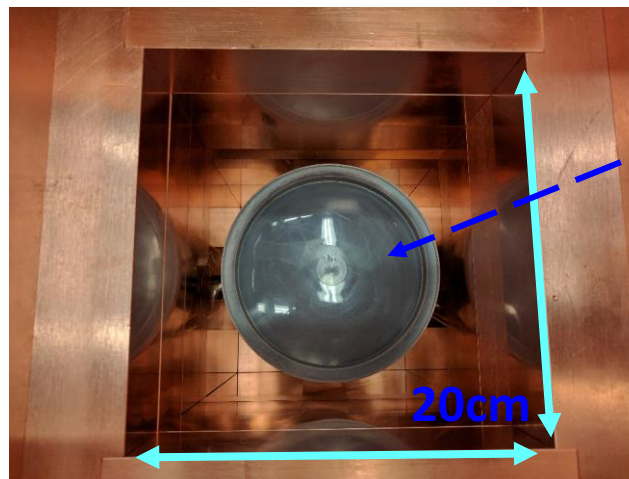
**plastic sealant**

**Springs used as a lifting mechanism**

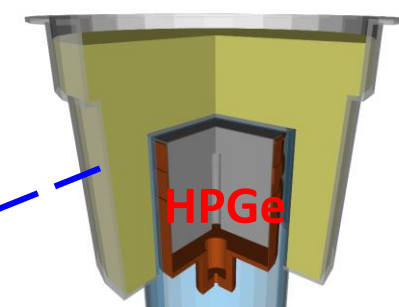
# The HPGe detector located underground (2700m.w.e.)

Isotope	Energy, keV	Background, Events / day
$^{40}\text{K}$	1460.8	10.0
$^{60}\text{Co}$	1332.5	1.6
$^{234\text{m}}\text{Pa}$	1001	3.1
$^{228}\text{Ac}$	911.2	3.2
$^{214}\text{Bi}$	609.3	2.3
$^{208}\text{Tl}$	583.2	3.3
$^{214}\text{Pb}$	351.9	3.7
$^{235}\text{U}$	185.7	22.5
$^{234}\text{Th}$	92.6	25.0

25cmPb + 5cm Cu

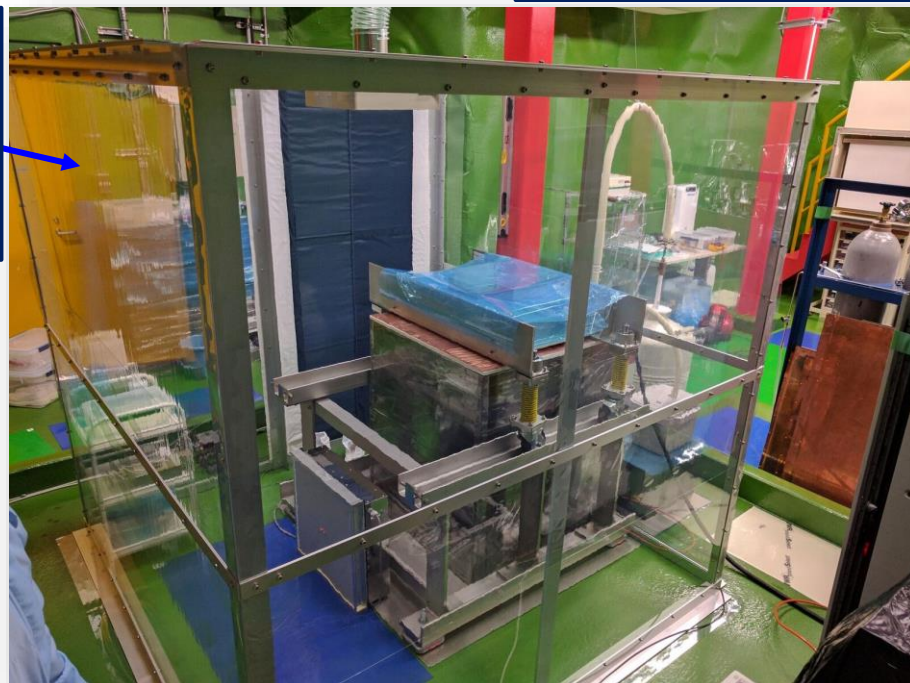


1.3L container



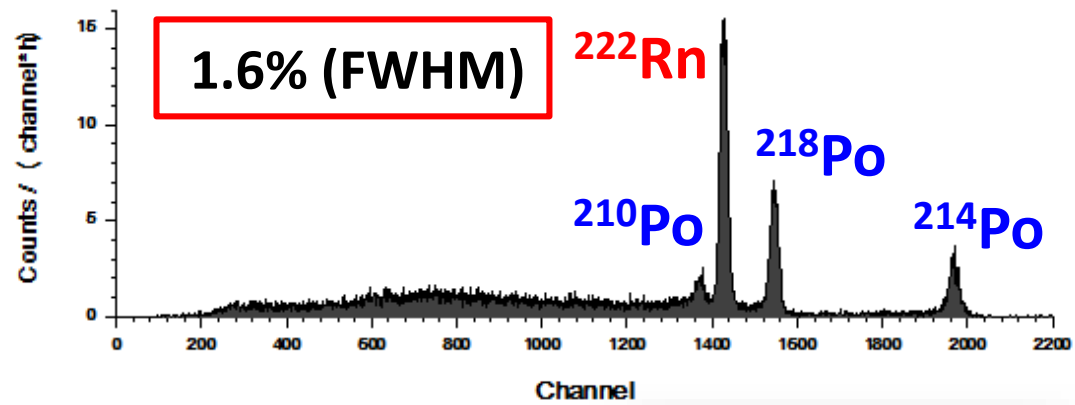
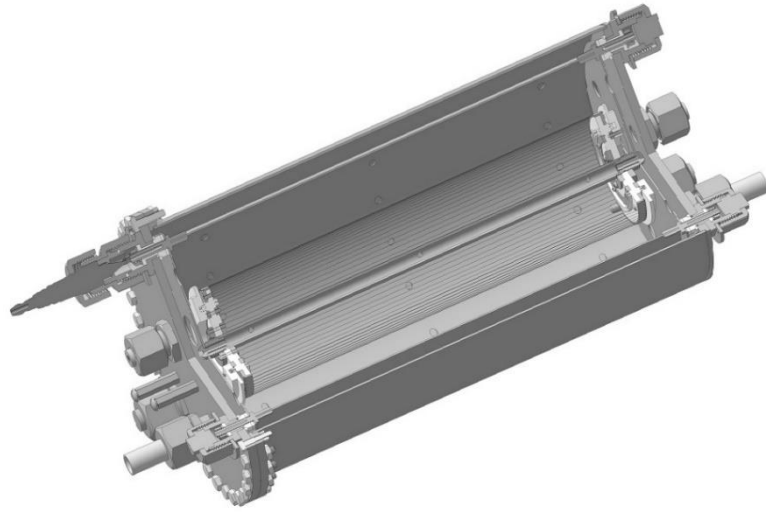
Enough space for e.g. 4-inch PMT

HPGe p-type  
75% rel. eff.  
**Room B**



- Shielding materials were kept **10+ yrs underground**
- Clean-room tent + radon-less air s-m**
- Detector's inner volume is purged with boiled off Nitrogen (**5.5L/min**)

# BNO high resolution ion-pulse ionization chamber



Direct detection of the Radon  $\alpha$ -decay in the air.

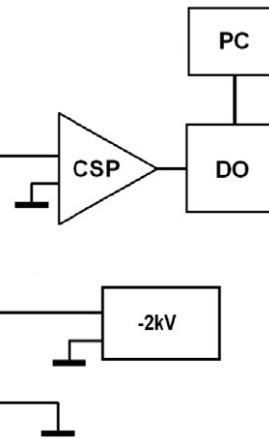
The 1<sup>st</sup> layer of acoustic shielding: a 5mm-thick acrylic box.



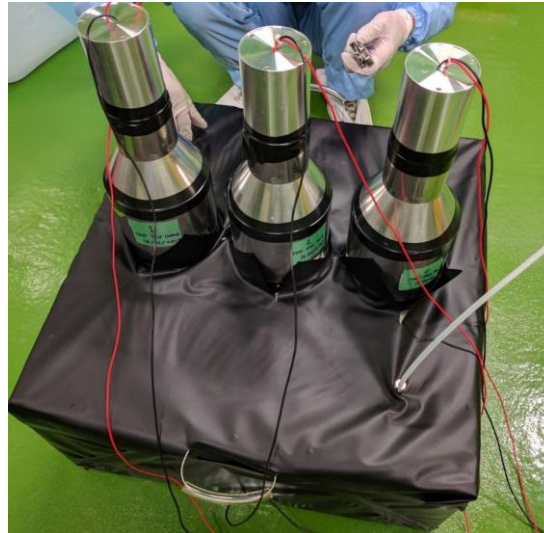
[JSPS grant: 16K05371](#)

Anode  
Cathode  
Insulators

Air sample



Currently, we build a **second layer of acoustic shielding** to minimize interference between radon measurements and our every day activities at the mine.

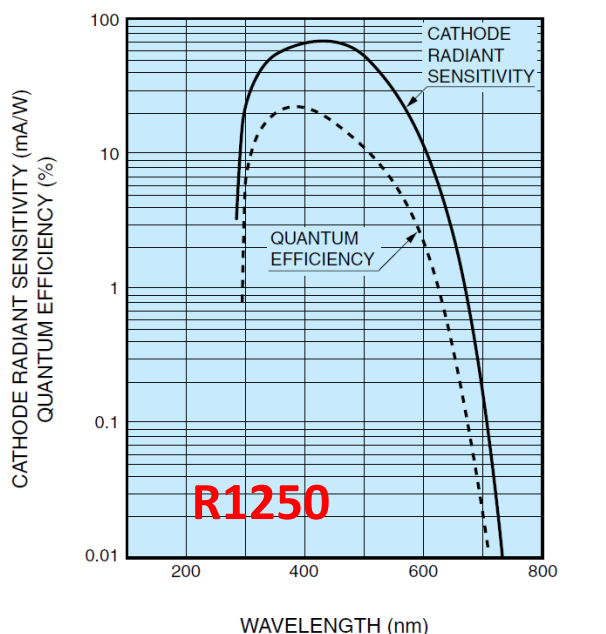
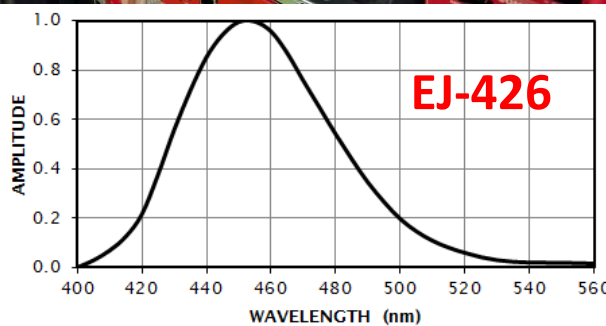


**Conical light reflectors** covered by **Tyvek** sheets. The Al box made air tight using the **Teflon sealant**.

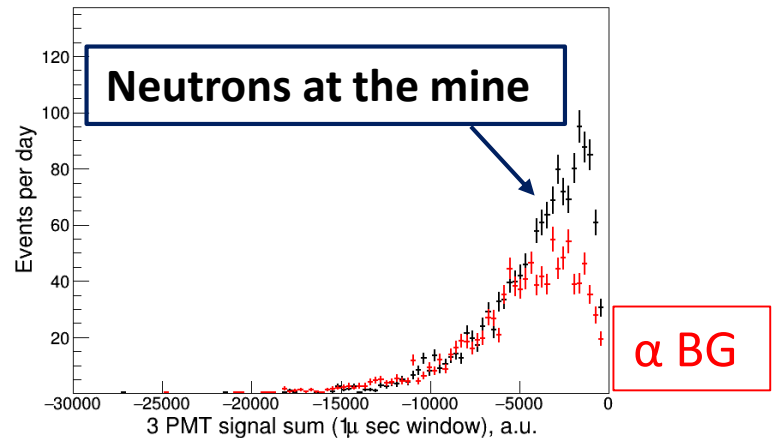
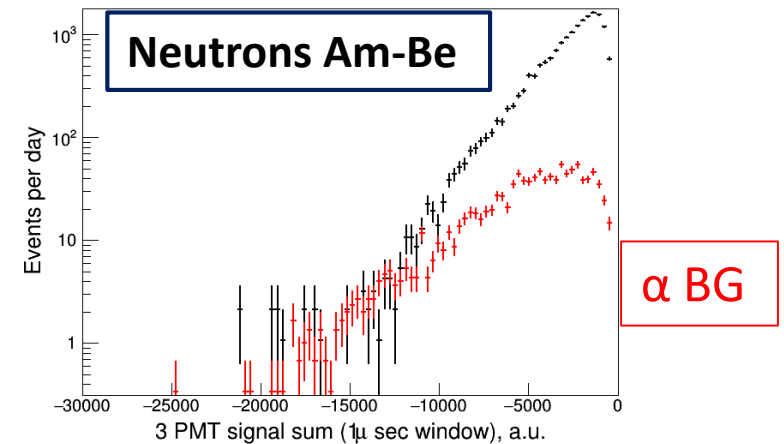
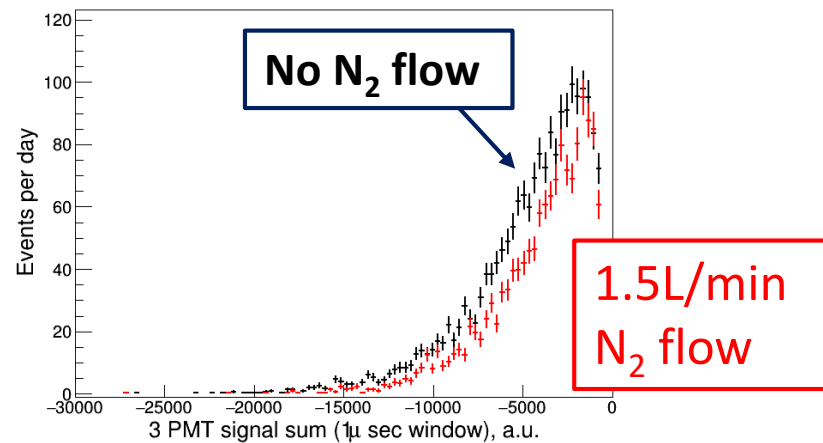
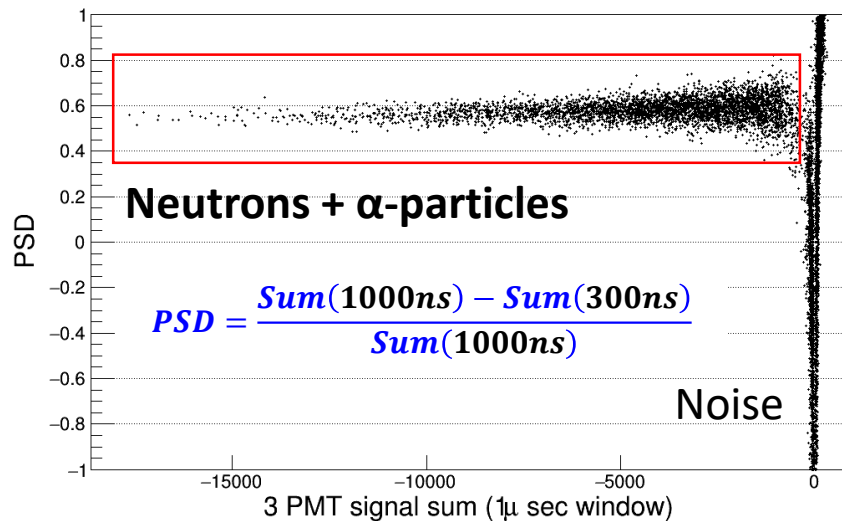
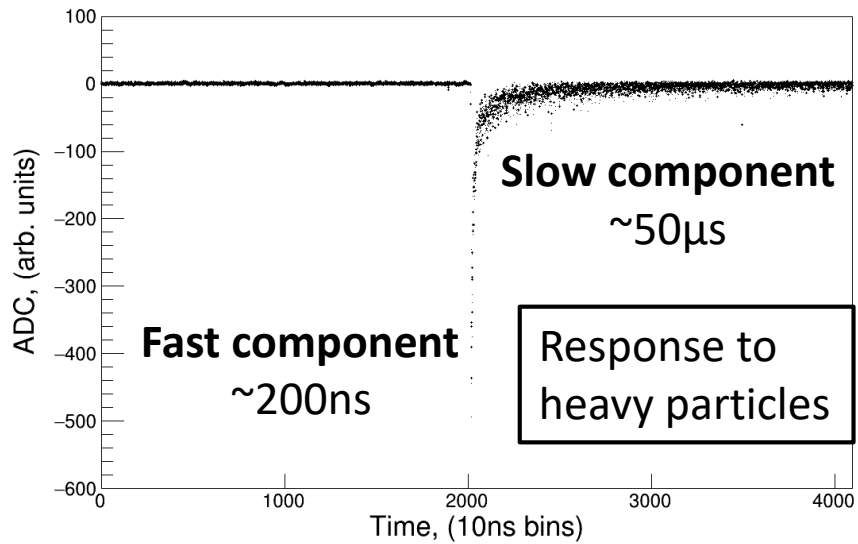
**JSPS grant: 16K05371**

**Hamamatsu R1250 PMTs**  
**CAEN 4 channel waveform digitizer N6724F (100MHz)**

- Two 0.32mm-thick **EJ-426** scintillator sheets (**0.25m × 0.5m each**) laminated by 0.25mm-thick clear polyester sheets from both sides were used.
- The EJ-226 is a homogeneous mixture of **LiF** and **ZnS:Ag** (mass ratio **1:2**), a **95% enriched <sup>6</sup>Li** was used. The ZnS:Ag light yield is **95000photons/MeV**.
- **${}^6\text{Li} + n \rightarrow \alpha + {}^3\text{H} + 4.78\text{MeV}$**  ( $\sigma = 941\text{barn}$ );
- The thermal detection efficiency is **34%**;
- The detector is **non-sensitive to  $\gamma$ -rays**;



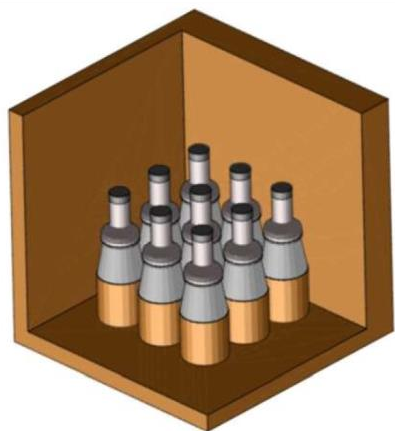
# The thermal neutron flux variation monitoring at the Kamioka mine



Impurity	DAMA/LIBRA	Our result
natK [ppb]	< 20	125
Th-chain [ppt]	0.5 ~ 7.5	0.3 ± 0.5
<sup>226</sup> Ra [μBq/kg]	21.7 ± 1.1	58 ± 4
<sup>210</sup> Pb [μBq/kg]	24.2 ± 1.6	30 ± 7



- The **R13444X** is **4-inch experimental ultra-low background Hamamatsu PMT** with a metal body. Optical window is made of Synthetic Silica with the Bialkali photocathode. Spectral response maximum at 420nm (200-650nm range); **QE@420nm : 34.9% and 33.38%**. **Gain at +1500V :  $5 \times 10^6$ ; TTS: 13ns**

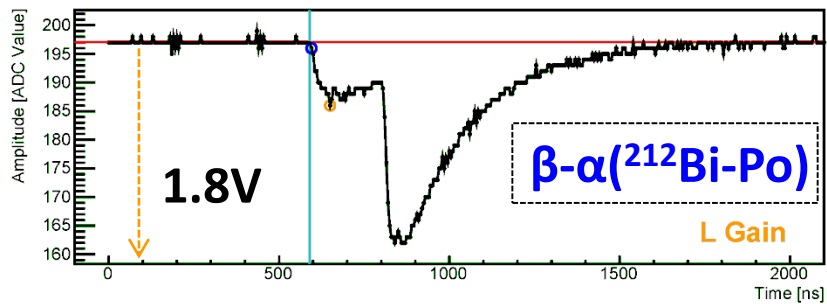
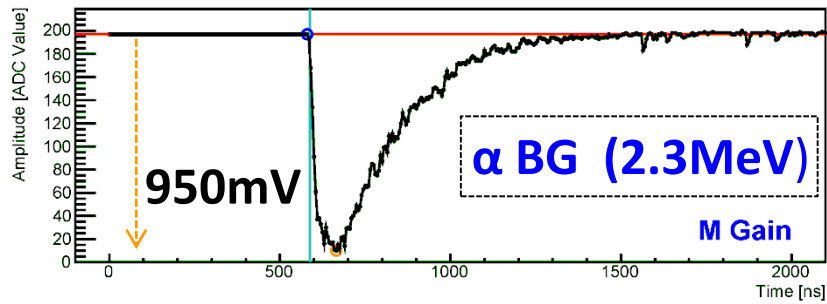
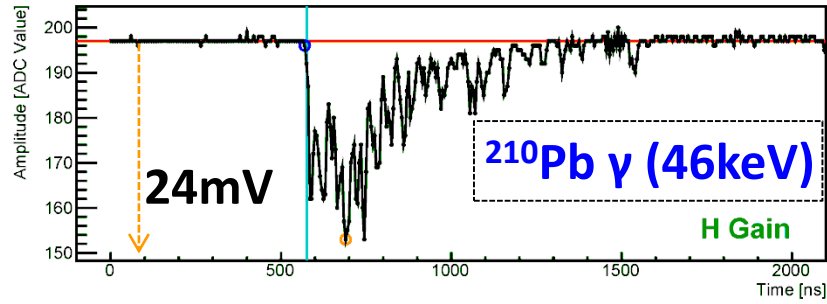
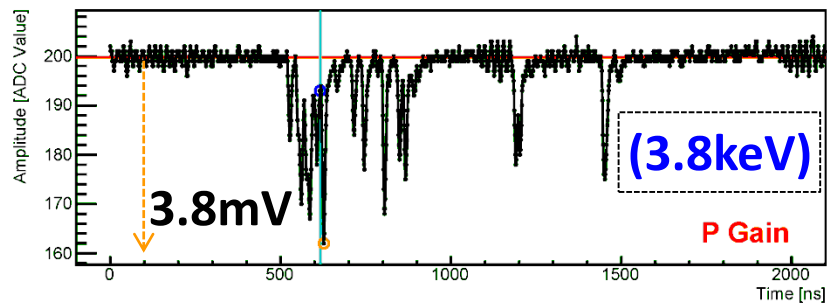


**Shielding: Old lead + 600kg of copper melted at Mitsubishi Materials** from freshly manufactured electroformed Cu sheets.

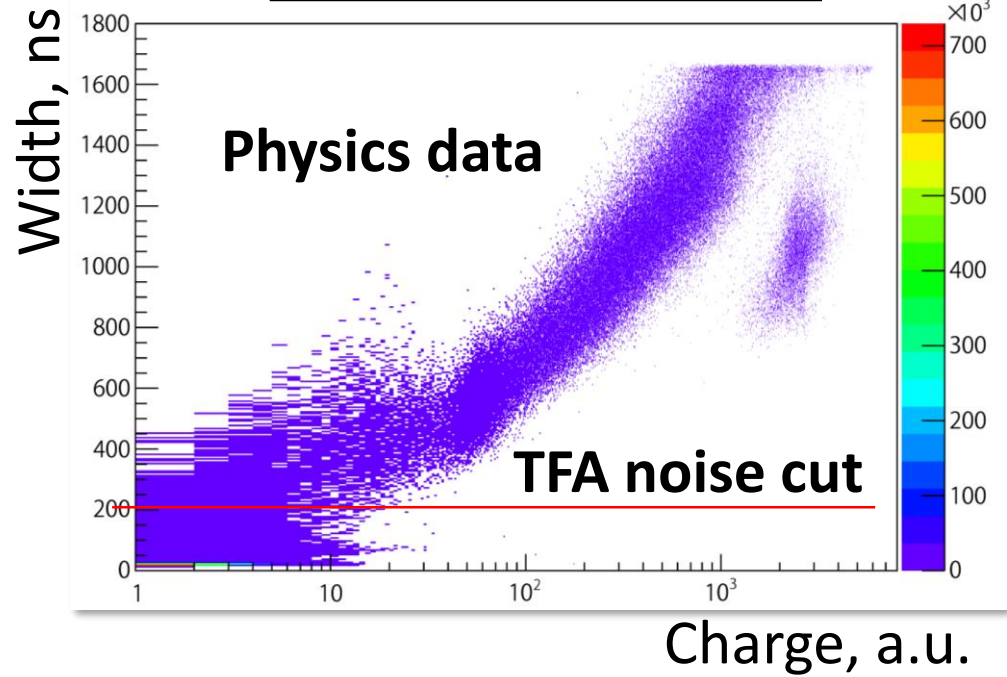
**DM detector (Phase 1): 5-inch cylindrical NaI(Tl) crystals viewed by R13444 PMTs**

We developed radio-pure NaI(Tl) crystals using the Bridgman method. For the new **5x5inch crystal** we expect **radio-purity similar to that of the DAMA/LIBRA crystals.**

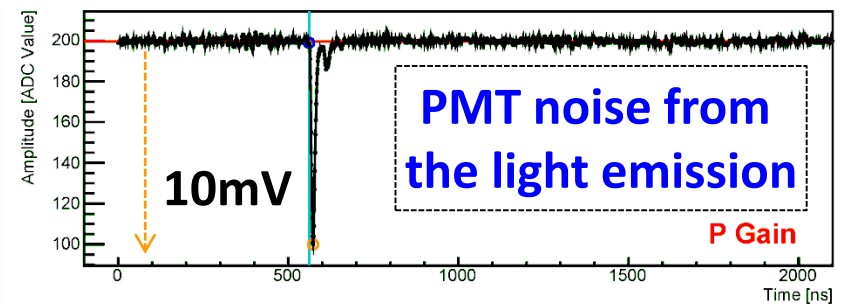




# The NaI(Tl) data



PMT noise >99.9% of data, from single to several narrow pulses (trains) was effectively cut by TFA.

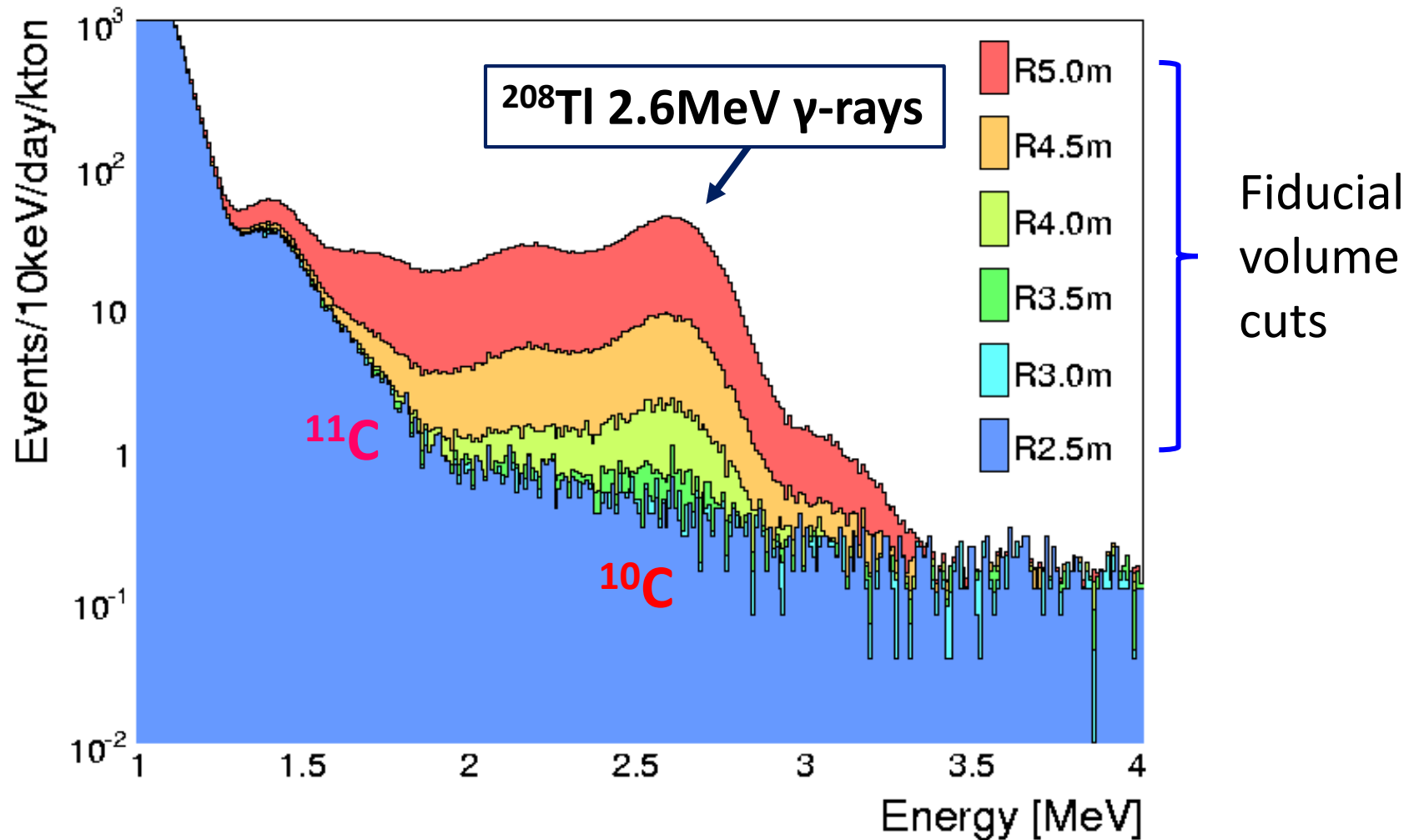


# Summary for the NaI(Tl) research

- ❑ We developed **5×5-inch highly radio-pure NaI(Tl) crystals** for rare event searches.
- ❑ In year 2018, we plan to start construction of the **NaI(Tl) DM detector** in a conventional Pb+Cu passive shielding that can be used for testing the **DM observation claim made by DAMA/LIBRA collaborations**. In more distant future, that detector could be deployed into the ultra-low background central region of KamLAND currently reserved for the KamLAND-Zen  $0\nu\beta\beta$  experiment.
- ❑ We intend to use our crystals as a detector for **precision studies of the coherent neutrino scattering**. Possible neutrino sources are JPARC, and IsoDAR.

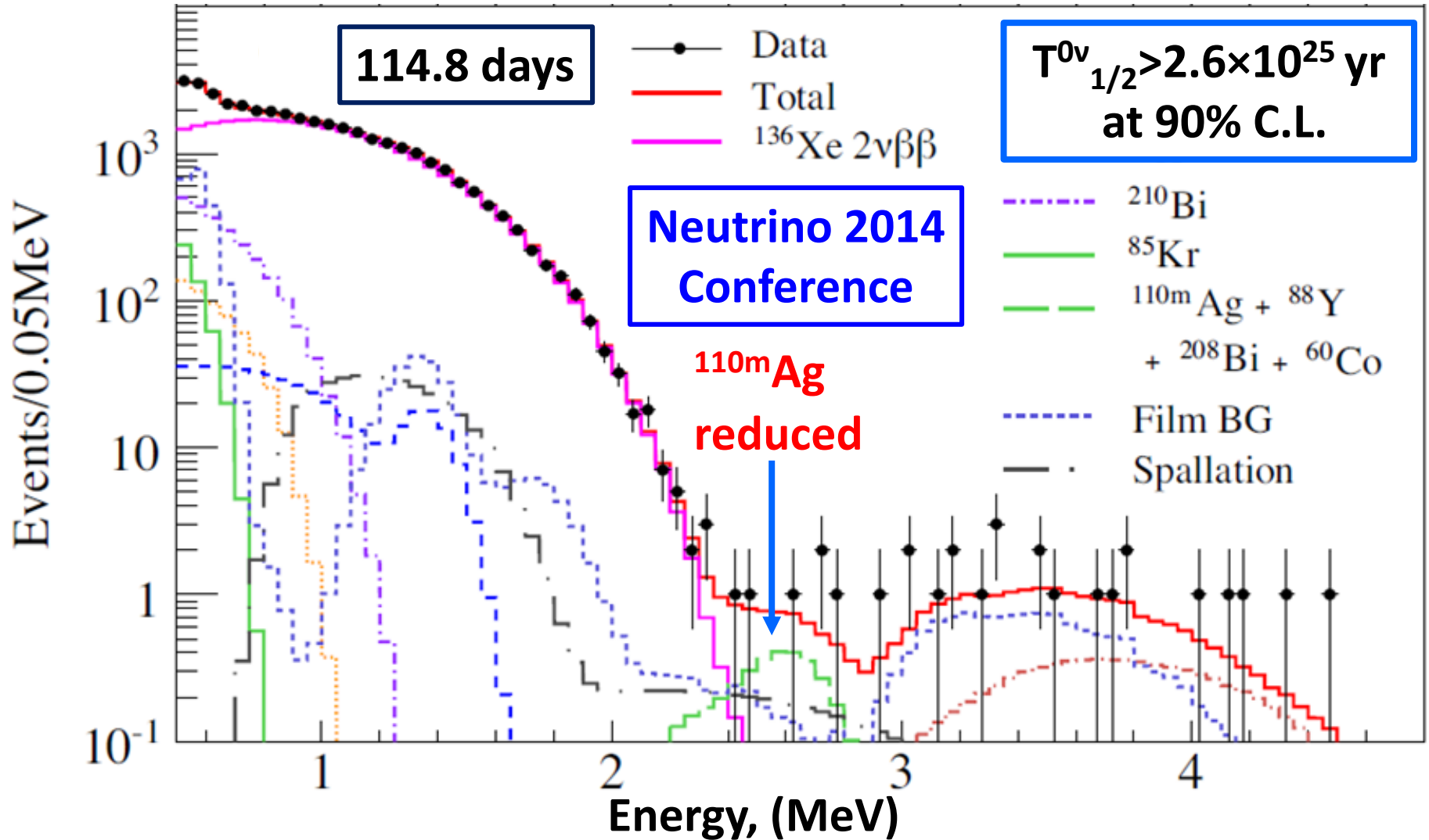
**BACKUP SLIDES**

# Use of KamLAND for rare event searches



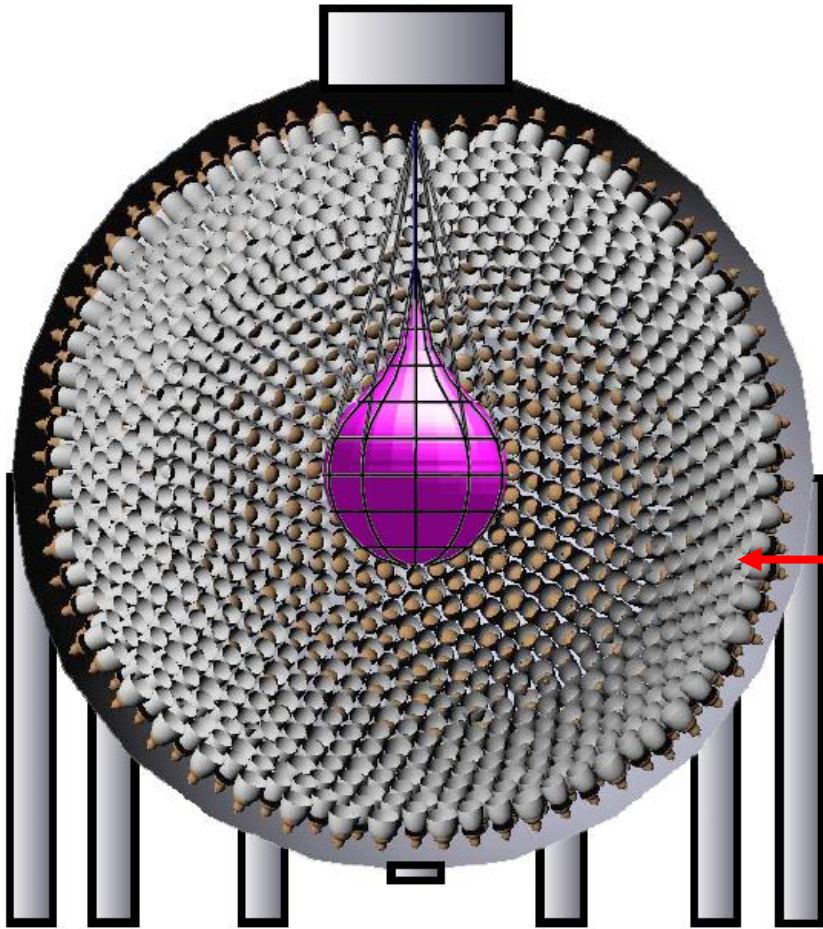
- **Existence of ultra-low background region** at the center of KamLAND makes searches for Dark Matter,  $0\nu\beta\beta$  signals possible.

# KamLAND-Zen 400: start of Phase II (year 2014)



During **Phase II** the same mini-balloon was used but amount of enriched xenon was increased from **320kg** to **383kg**.

# KamLAND2-Zen to cover the IH mass region



We need to detect **more light** to improve energy resolution → reduce the  **$2\nu\beta\beta$  tail background**.

**Sensitivity target:  $m_{\beta\beta} \sim 20\text{meV}$**



**Enriched xenon mass > 1000kg**

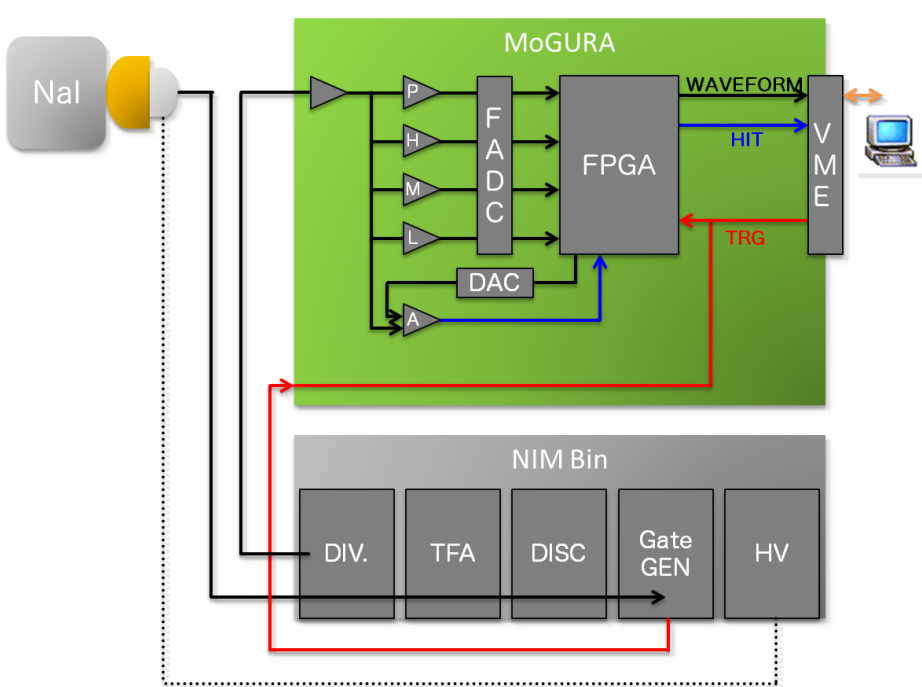
**Gain in number of detected photons**

(after upgrade to KamLAND2)

**New LAB scintillator: 1.4 times**

**High QE PMTs: 1.9 times**

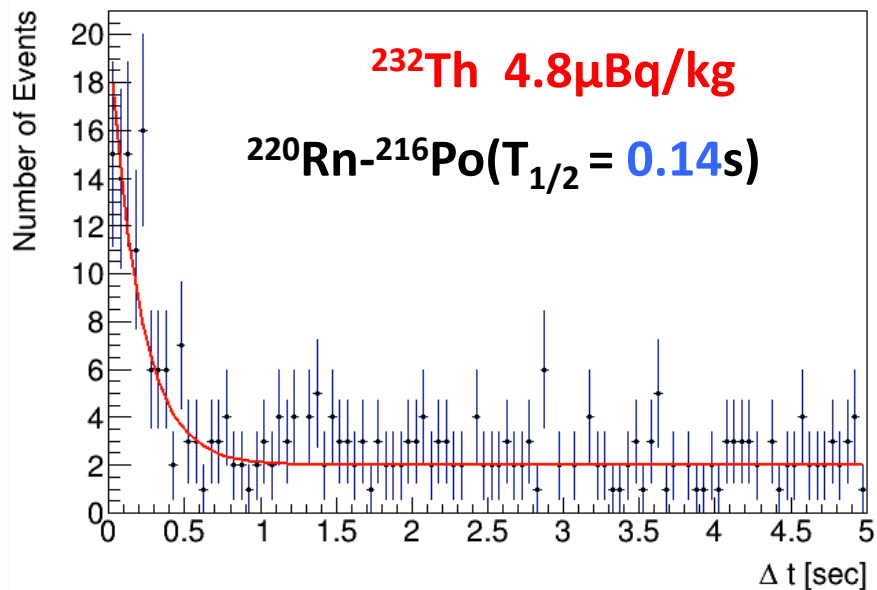
**Light collecting cones: 1.8 times**



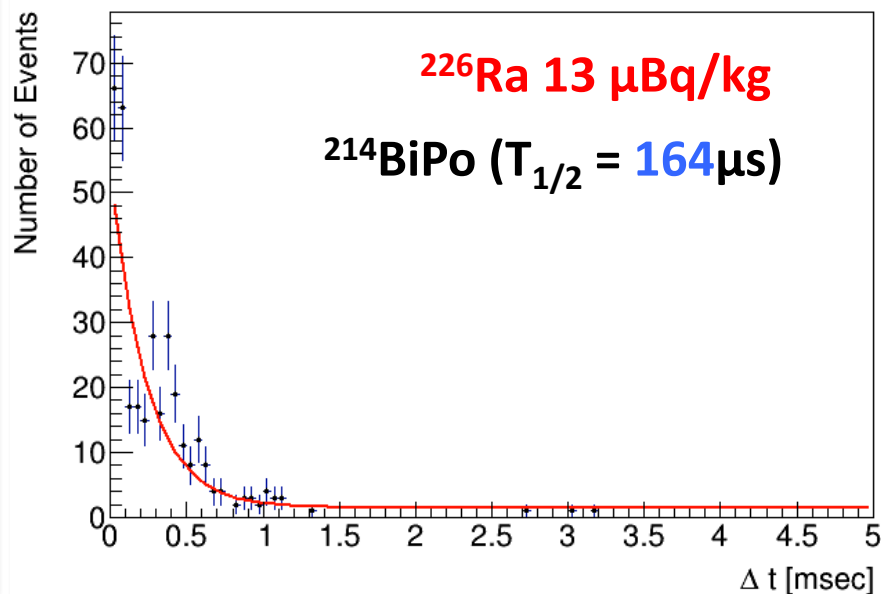
## MoGURA based DAQ

- **12ch** input VME 9U board
- **FADC**
  - **P** : 0.1mV/LSB, 8bit, 1GSPS
  - **H** : 0.5mV/LSB, 8bit, 200MSPS
  - **M** : 5 mV/LSB, 8bit, 200MSPS
  - **L** : 50 mV/LSB, 8bit, 200MSPS
  - 0.1mV ÷ 10V
- **FPGA**
  - Up to 10 μsec waveform buffer
- **HIT**
  - Analog discri.: >5mV
  - Digital discri.: >0.5mV
- **TFA** for PMT noise cut

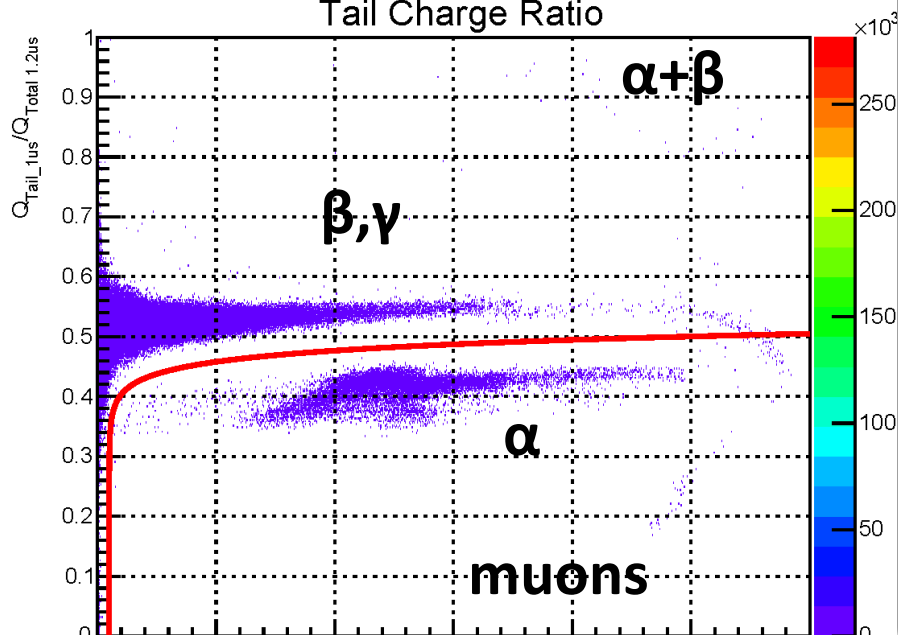
$\alpha$ - $\alpha$  interval ( $^{220}\text{Rn} \rightarrow ^{216}\text{Po}$ )



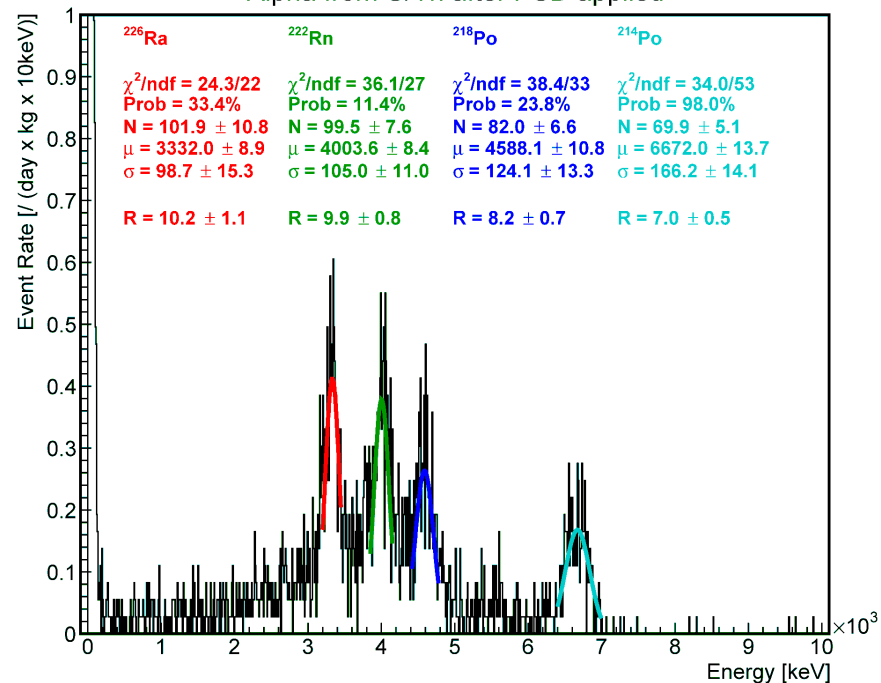
$\beta$ - $\alpha$  interval ( $^{214}\text{Bi} \rightarrow ^{214}\text{Po}$ )



Tail Charge Ratio

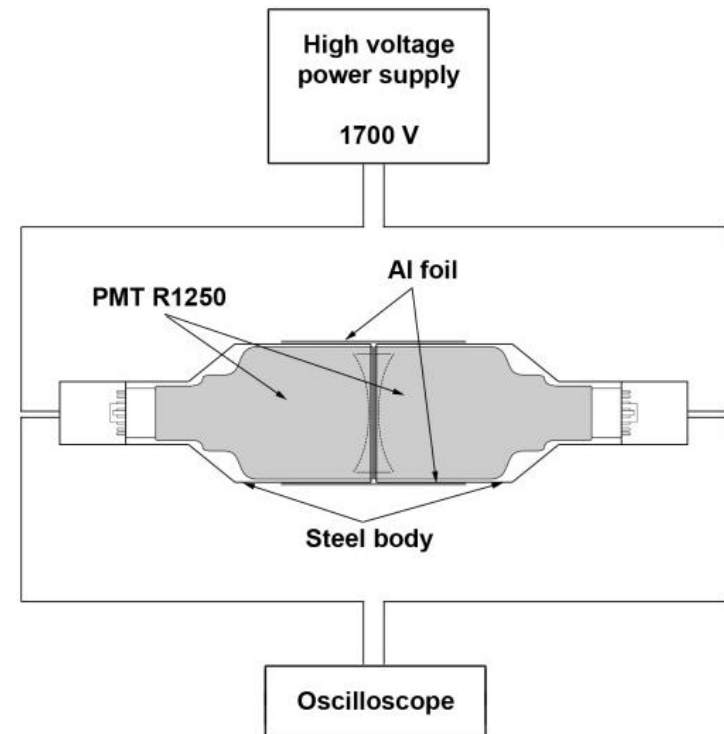


Alpha from U/Th after PSD applied





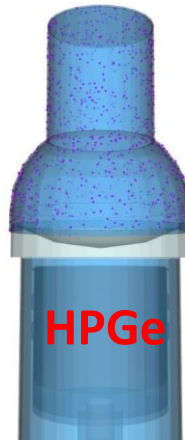
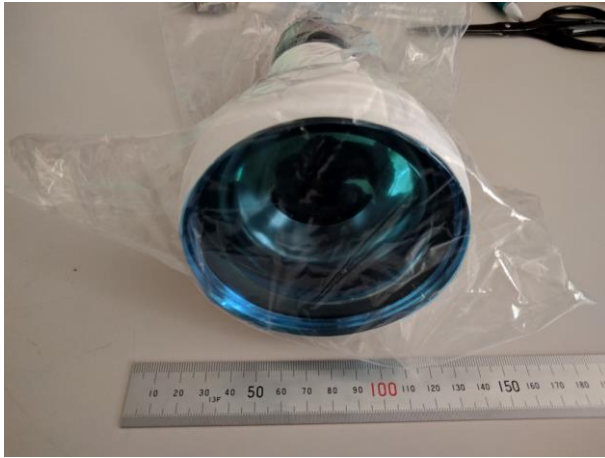
# Problem of high intensity fast PMT noise



Coincident noise pulses in **face-to-face** optically coupled PMTs.

During development of the NaI(Tl) DM detectors we faced a problem of **intense high amplitude fast PMT noise** (~10ns wide pulses). The noise was observed for several Hamamatsu PMT types: R6091, R11065-20 3-inch PMTs, R13444X 4-inch PMT, R1250 5-inch PMT. Use of different DAQ hardware, opened or underground locations had no effect on this noise. **Existence of the noise was admitted by the Hamamatsu Photonics.** Most likely cause - **fast flashes of light in the PMTs.**

# Experimental ultra-low background PMT R13444X



The PMT is made of about **37 components**. Materials: **Ni-Fe alloy, SUS304, ceramics, Al (5N), Ag, Ni, ...** . Hard to make a realistic GEANT model based limited information available.

Need **~1 month** long measurement at the HPGe detector **per PMT**, we are taking data now. Preliminary, confirmed presence of  **$^{40}\text{K}$  at few tens mBq/PMT,  $^{60}\text{Co}$  at few mBq/PMT**. Search for radio-pure voltage divider parts with Hamamatsu.

- The **R13444X** is the largest (4-inch) ultra-low background Hamamatsu PMT with a metal body. Optical window is made of Synthetic Silica with the Bialkali photocathode. I purchased **two PMTs** this spring to check if it could be used as replacement for 3-inch Hamamatsu ultra-low background R11165-20 PMT.
- Spectral response maximum at 420nm (200-650nm range); **QE@420nm : 34.9% and 33.38%**. **Gain** at +1500V :  $5 \times 10^6$ ; **TTS**: 13ns