



# SOLAR NEUTRINO CAPTURE CROSS-SECTION FOR <sup>76</sup>Ge NUCLEI \*

FAZLIAKHMETOV, ALMAZ <sup>1,2</sup> LUTOSTANSKY, Yury <sup>3</sup>

INZHECHIK, LEV<sup>1</sup> TIKHONOV, VICTOR<sup>3</sup> KOROTEEV, GRIGORY<sup>1</sup> VYBOROV, ANDREI<sup>1,2</sup>

<sup>1</sup>Moscow Institute of Physics and Technology, Russia <sup>2</sup>Institute for Nuclear Research, Russia

<sup>3</sup>National Research Center Kurchatov Institute, Russia

\*posted on <a href="https://arxiv.org/abs/1810.07452">https://arxiv.org/abs/1810.07452</a>



MAJORANA experiment

GERDA experiment

 $0\nu\beta\beta\ decay:$   $^{76}_{32}Ge \rightarrow ^{76}_{34}Se + 2e^-, \quad Q_{0\nu\beta\beta} = 2039\ keV$ 

v backround:  $v_e^{\text{solar}} + {}^{76}Ge \rightarrow {}^{76}As + e^- + n\gamma$  ${}^{76}As \rightarrow {}^{76}Se + e^- + \overline{v_e} + m\gamma$ 



background events in MAJORANA experiment (2017)

### MAJORANA, GERDA $\rightarrow$ LEGEND



Scheme of transitions



Scheme of transitions

$$\sigma_{total}(E_v) = \sigma_{discrete}(E_v) + \sigma_{resonances}(E_v) \quad (= 0, if E_v \le Q_{EC})$$
(1)  
levels

$$\sigma_{discrete}\left(E_{v}\right) = \sum_{k} \frac{G_{F}^{2} cos^{2} \theta_{c}}{\pi} p_{e} E_{e} F(Z, E_{e}) \left[B(F)_{k} + \left(\frac{g_{A}}{g_{V}}\right)^{2} B(GT)_{k}\right]$$
(2)

$$\sigma_{\rm res}(E_{\nu}) = \frac{1}{\pi} \int_{\varepsilon_{\rm min}}^{\varepsilon_{\rm max}} G_{\rm F}^2 \cos^2\theta_{\rm C} p_e E_e F(Z, E_e) S(E) dE$$
(3)

 $G_F$  – the weak coupling constant,  $\theta_c$  – the Cabibbo angle  $p_e/E_e$  – the outgoing electron momentum / total energy  $F(Z, E_e)$  – the Fermi function  $\frac{g_A}{g_V}$  – the ratio of the axial vector and vector coupling constants  $B(F)_k$ ,  $B(GT)_k$  – the Fermi (Gamow-Teller) response k-th excited state of  ${}^{76}_{33}As$  \*  $S_\beta(E)$  – nuclear beta strenght function  $E_{sep}$  – neutron separation energy,  $E_{sep}$  = 7.3 MeV  $\varepsilon_{min}$  = 5 MeV,  $\varepsilon_{max} = E_{sep}$  $\nu_{e \ solar}$ +

$$v_{e \ solar} + {}^{76}_{32}Ge \rightarrow {}^{76}_{33}As + e^{-1}$$

S(E), MeV<sup>-1</sup>

*Charge-exchange strength function of the* <sup>76</sup>*Ge*(<sup>3</sup>*He, t*) <sup>76</sup>*As reaction S*(*E*)



S(E), MeV<sup>-1</sup>

*Charge-exchange strength function of the* <sup>76</sup>*Ge*(<sup>3</sup>*He, t*) <sup>76</sup>*As reaction S*(*E*)



S(E), MeV<sup>-1</sup>

*Charge-exchange strength function of the* <sup>76</sup>*Ge*(<sup>3</sup>*He, t*) <sup>76</sup>*As reaction S*(*E*)



*Cross-sections for*  ${}^{76}Ge(v_e, e){}^{76}As$  *reaction* 



*Flux density of incident neutrinos* (*BS*05 *model*)



#### Rate of solar neutrino capture

Capture rate of solar neutrinos [*SNU]	рер	hep	<sup>13</sup> N	<sup>17</sup> F	<sup>15</sup> 0	<sup>7</sup> B	Total capture rate
R <sub>discr</sub>	1.369	0.0451	0.102	0.021	0.828	13.54	15.9
R <sub>total</sub>	1.369	0.090	0.102	0.021	0.828	21.17	23.58
R <sub>discr</sub> + R <sub>GTR</sub>	1.369	0.070	0.102	0.021	0.828	17.46	19.85
R <sub>GTR</sub> / R <sub>total</sub>	0%	28%	0%	0%	0%	19%	17%
R <sub>cont</sub> / R <sub>discrete</sub>	0%	99%	0%	0%	0%	56%	48%

\*SNU =  $10^{-36} \frac{1}{nucleon \cdot s}$ 

#### Summary:

- $\bullet$   $\sigma(E_v)$  were determined using S(E) for discrete and continuous states
- **\*** The effect of resonance structure on  $\sigma(E_v)$  was investigated
- ✤ 17% contribution of only GTR
- ✤ Account of continuous states increases R by up 50%

## It is planned to:

- Consider other types of resonances and states behind S<sub>n</sub>
- Consider neutrino oscillations

★ Calculate the contribution of the secondary process  ${}^{76}As \rightarrow {}^{76}Se + e^- + \overline{\nu_e}$  (account the design of detecting elements)

# THANK YOU FOR YOUR ATTENTION!