





Thunderstorm investigations based on the data obtained by the URAGAN muon hodoscope and Doppler weather radar DMRL-C

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Muon hodoscope URAGAN

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- provides registration of muons with a high spatial and angular resolution (about 1 cm and 0.8°, respectively) in a wide range of zenith angles (from 0° to 84°)
- track parameters are reconstructed in real time and accumulated in a two-dimensional matrix for one minute interval. Such a matrix is a "muon picture" of the upper hemisphere

characteristics reflecting the distortion of the distribution of the muon flux used in this analysis are:

- vector of local anisotropy (A) which indicates the average direction of muon arrival recorded by the detector
- vector of relative anisotropy (R) which is the difference between the local anisotropy vector in the current time and averaged over a long period anisotropy vector and its projection on the horizontal plane (R_{hor})

Streamer tube chambers Foam plastic
Movable platform

Doppler weather (meteorological) radar DMRL-C



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Displaying distribution of various meteorological data such as precipitation intensity and velocity and direction of movement of cloud systems

- a view range of 250 300 km
- a maximum detection height of about 20 km
- a high spatial resolution (0.5 1 km)
- Detection of dangerous weather phenomena such as thunderstorms

The research

- Muon pictures (muonographs) and meteorological maps obtained every ten minutes by means of two facilities during thunderstorm periods of spring and summer of 2018 (~27 events) are compared with each other
- Wavelet analyses of obtained data are performed
 - Results are illustrated by thunderstorm events occurred in Moscow on July 14, 2018 and August 30, 2018

14.07.2018

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The thunderstorm was observed at the Domodedovo airport meteorological station at 14:00 (UTC time) and at the VDNKh meteorological station from 14:00 to 17:00









Comparison (11:30 – 13:10)



Wavelet analyses of URAGAN data



Wavelet analyses of URAGAN





30.08.2018 - 31.08.2018









Comparison (20:30 – 05:20)



Wavelet analyses of URAGAN data



Wavelet analysis: Rhor

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Conclusions

- Results obtained from the analysis of data provided by MH URAGAN are consistent with DMRL-C data
- Characteristics of the muon flux provide a good reaction to the passage of a thunderstorm
- Wavelet analysis of different time series shows that before and during the passage of a thunderstorm there are periodic disturbances in the characteristics of the muon flux. Periods of those disturbances are ranging from 60 to 180 minutes

Thank you for your attention!

Characteristics reflecting the distortion of the distribution of the muon flux

- Vector of local anisotropy (A) indicates the average direction of muon arrival recorded by the detector. This value is calculated as the sum of unit vectors, each of which has a reconstructed direction of a single muon track, normalized by the total number of muons.
- The vector of relative anisotropy (r) is the difference between the local anisotropy vector in the current time and averaged over a long period anisotropy vector

$$\vec{R} = \vec{A} - \vec{A}_{N}$$
, $R_{hor} = \sqrt{R_{South}^2 + R_{East}^2}$

- In this analysis, the value of the vector **r** module and its projection on the horizontal plane (r_{hor}) , the projections onto the horizontal axis South (r_{south}) and East (r_{east}) , as well as the projection on the vertical axis (r_z) are used
 - Registration of muon flux variations at the Earth's surface from different directions allows to obtain a picture of the conditions in the upper atmosphere, to trace the dynamics of their changes and to identify disturbed regions

14.07 Wavelet analysis:

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14.07 Wavelet analysis: A



30.08 Wavelet analysis: Isum



30.08 Wavelet analysis: A



Muon diagnostics

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- Formation of the muon flux: in the atmosphere at an altitude of 10 - 20 km
- \cdot The top of a typical thunderstorm cell: at a height of 8 - 12 km

The muon flux is modulated by in the basic changes thermodynamic parameters of the atmosphere. ΛN

$$\frac{\mathbf{v}_{\mu}}{T} = \beta_p \Delta p + \beta_T \Delta T$$

The higher the temperature, the:

- higher pion generation point
- longer geometric path of the muon to the surface of the Earth
- higher probability of decav

The greater the pressure, the greater muons absorption (due to an increase the amount of substance)



Wavelet analyses of URAGAN data

