

Neutrino oscillations: status and prospects of accelerator and reactor experiments

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OUTLINE

- **Neutrino oscillations**
- **Recent oscillation results**
 - **Accelerators: T2K, NOVA**
 - **IceCube**
 - **Reactors: Daya Bay, RENO, Double Chooz**
 - **Search for CP violation**
- **Future oscillation experiments**
 - **JUNO**
 - **DUNE**
 - **HyperKamiokande**
- **Light sterile neutrinos**
 - **current status: pro et contra**
 - **coming results**

ν oscillations and mixing

Standard Model: neutrinos are *massless* particles

3 families

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

U parameterization:

three mixing angles θ_{12} θ_{23} θ_{13}
 CP violating phase δ_{CP}

Physics beyond the Standard Model

atmospheric

link between
atmospheric and solar

solar

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

SuperK, K2K,
MINOS, T2K, NovA

T2K
MINOS

Daya Bay, RENO
Double Chooz

Solar experiments, SuperK
KamLAND

$$\theta_{23} \sim 45^\circ$$

$$|\Delta m_{32}^2| \cong |\Delta m_{31}^2| =$$

$$|\Delta m_{atm}^2| \approx 2.4 \times 10^{-3} \text{ eV}^2$$

$$\theta_{13} \approx 9^\circ$$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

$$\Delta m_{12}^2 + \Delta m_{23}^2 + \Delta m_{31}^2 = 0$$

$$\theta_{12} \approx 34^\circ$$

$$\Delta m_{21}^2 = \Delta m_{sol}^2 \approx 7.5 \times 10^{-5} \text{ eV}^2$$

two independent Δm^2

Physics beyond the Standard Model

Main goals of oscillation experiments

- CP violation in lepton sector

Strength of CP violation in neutrino oscillations

$$J_{CP} = \text{Im}(U_{e1} U_{\mu 2} U_{e2}^* U_{\mu 1}^*) = \text{Im}(U_{e2} U_{\mu 3} U_{e3}^* U_{\mu 2}^*)$$

$$= \cos\theta_{12} \sin\theta_{12} \cos^2\theta_{13} \sin\theta_{13} \cos\theta_{23} \sin\theta_{23} \sin\delta_{CP}$$

all mixing angles $\neq 0 \rightarrow$
 $\rightarrow J_{CP} \neq 0$ if $\delta_{CP} \neq 0$

First indication from T2K: $\delta_{CP} = -\pi/2$??

neutrinos

$$V_{MNS} \sim \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

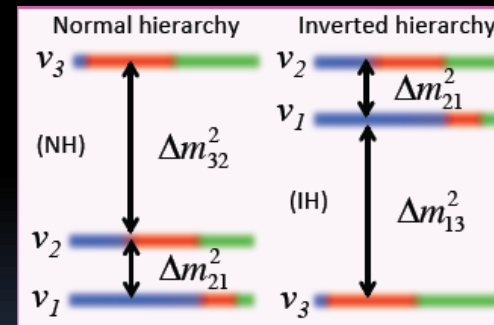
quarks

$$V_{CKM} \sim \begin{pmatrix} 1 & 0.2 & 0.001 \\ 0.2 & 1 & 0.01 \\ 0.001 & 0.01 & 1 \end{pmatrix}$$

Quark sector $J_{CP} \approx 3 \times 10^{-5}$

Lepton sector $J_{CP} \sim 0.02 \times \sin\delta_{CP}$

- Neutrino mass hierarchy



- θ_{23} – maximal? If not, what octant ($\theta_{23} > \pi/4$ or $\theta_{23} < \pi/4$)?

Neutrino cross sections

- Sterile neutrinos

Current experiments



about 500 members
59 institutions
from 11 countries

LONG-BASELINE NEUTRINO OSCILLATION EXPERIMENT



Super-K

Toyama
Kamioka Mine



JPARC
Tokai



JAPAN

Tokyo

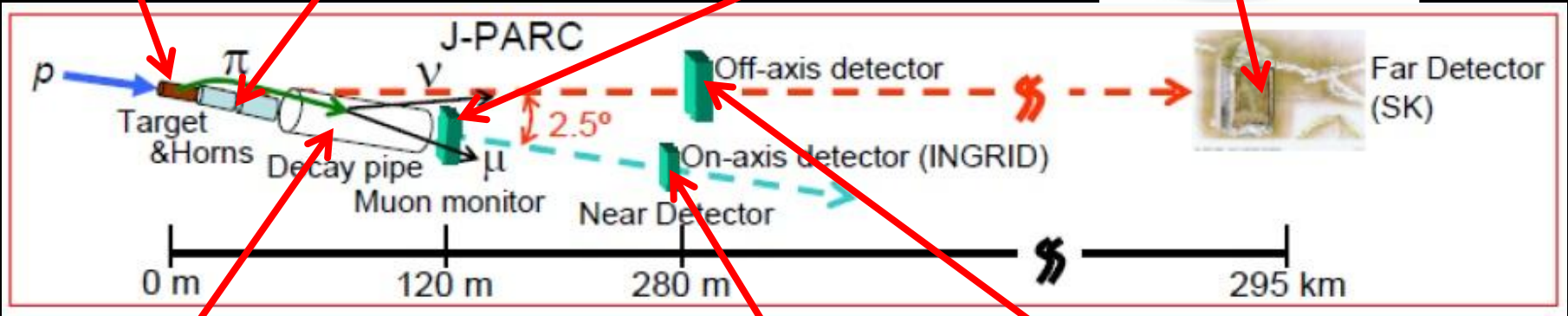
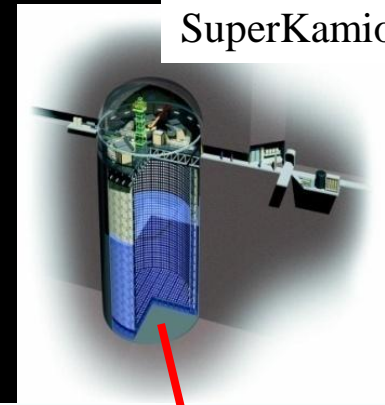
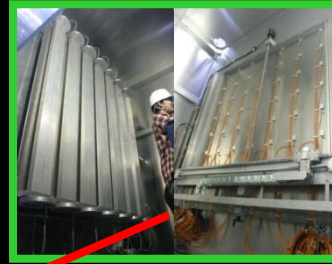
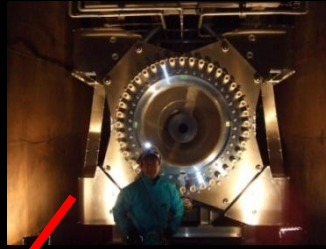
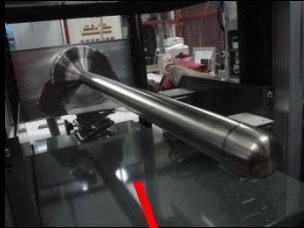


Tokyo/Narita Airport

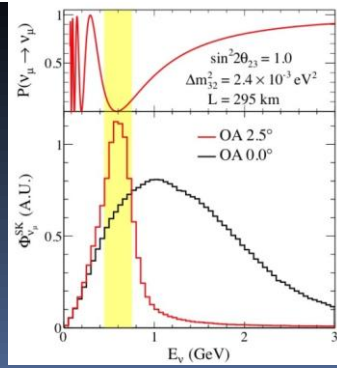
T2K experiment

Data taking since 2010

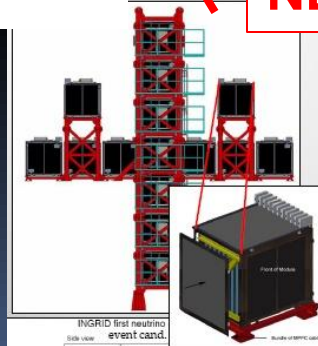
Far neutrino detector
SuperKamiokande



Off-axis neutrino beam

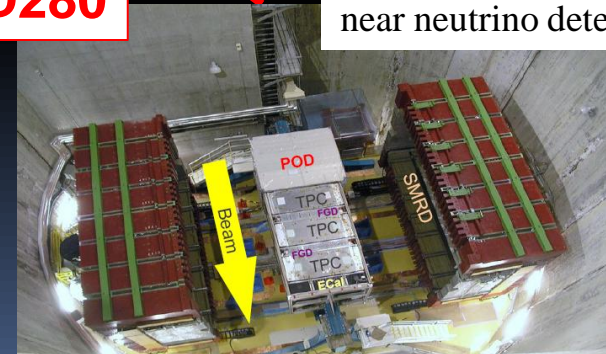


Neutrino monitor
INGRID



ND280

Off-axis near neutrino detector

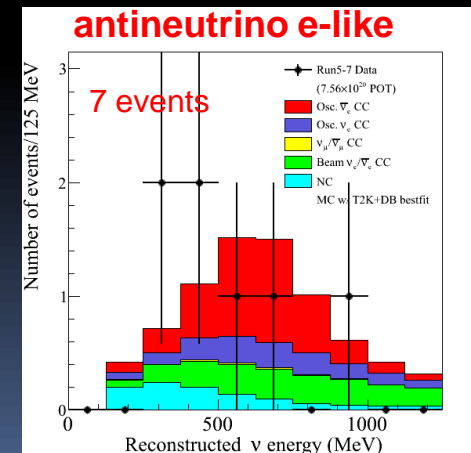
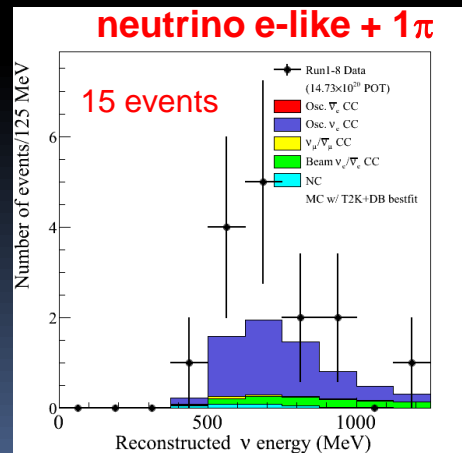
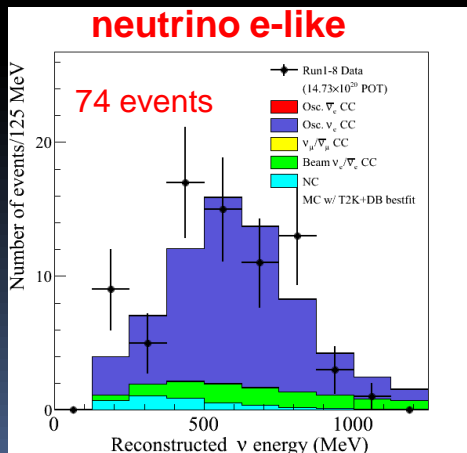


T2K: appearance

By Summer 2017

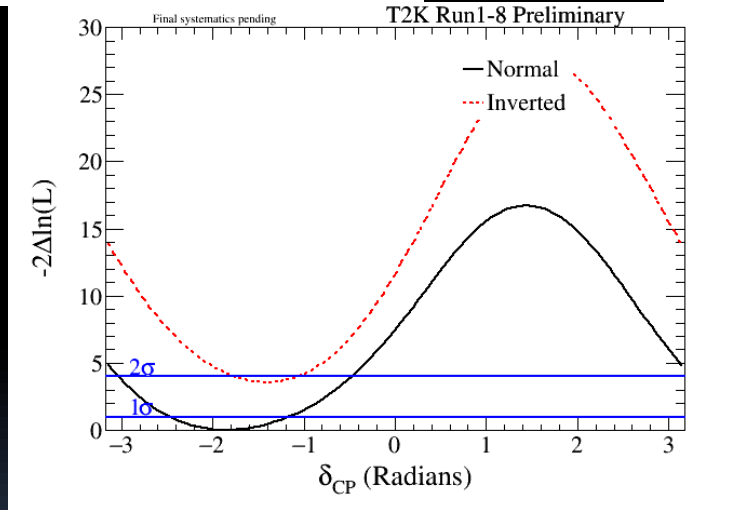
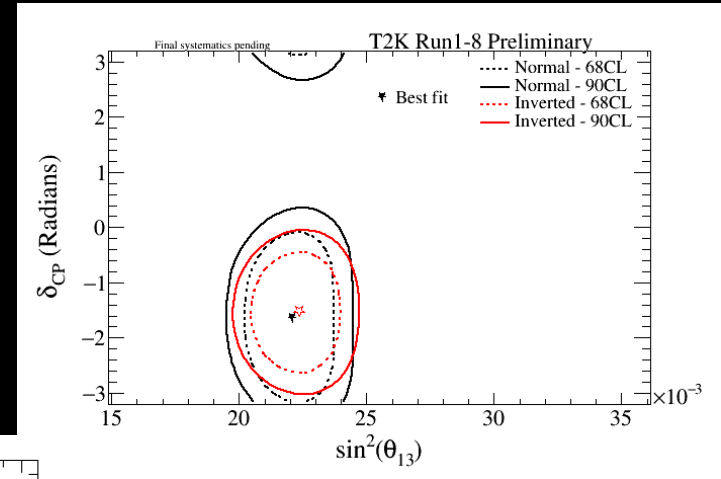
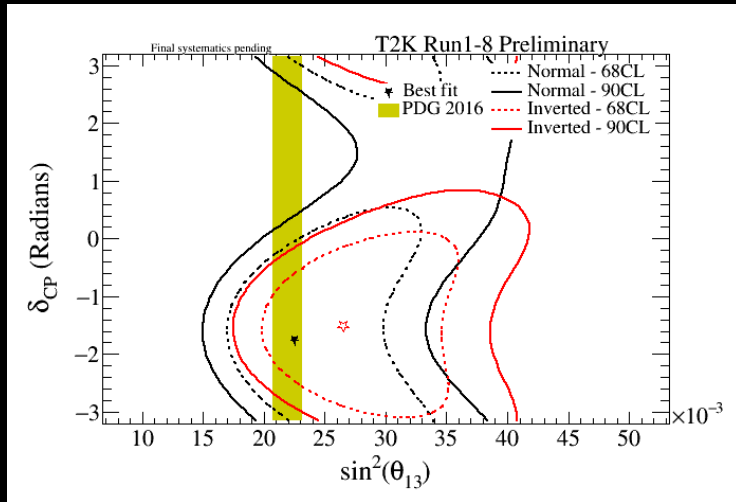
Neutrino beam 14.7×10^{20} POT
Antineutrino beam 7.6×10^{20} POT

Sample	Predicted events				Observed events
	$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = \pi/2$	$\delta_{CP} = \pi$	
CCQE 1-Ring e-like ν -mode	73.5	61.5	49.9	62.0	74
CC1 π 1-Ring e-like ν -mode	6.92	6.01	4.87	5.78	15
CCQE 1-Ring e-like $\bar{\nu}$ -mode	7.93	9.04	10.04	8.93	7
CCQE 1-Ring μ -like ν -mode	267.8	267.4	267.7	268.2	240
CCQE 1-Ring μ -like $\bar{\nu}$ -mode	63.1	62.9	63.1	63.1	68



T2K: CP

M.Antonova, talk at ICPPA2017

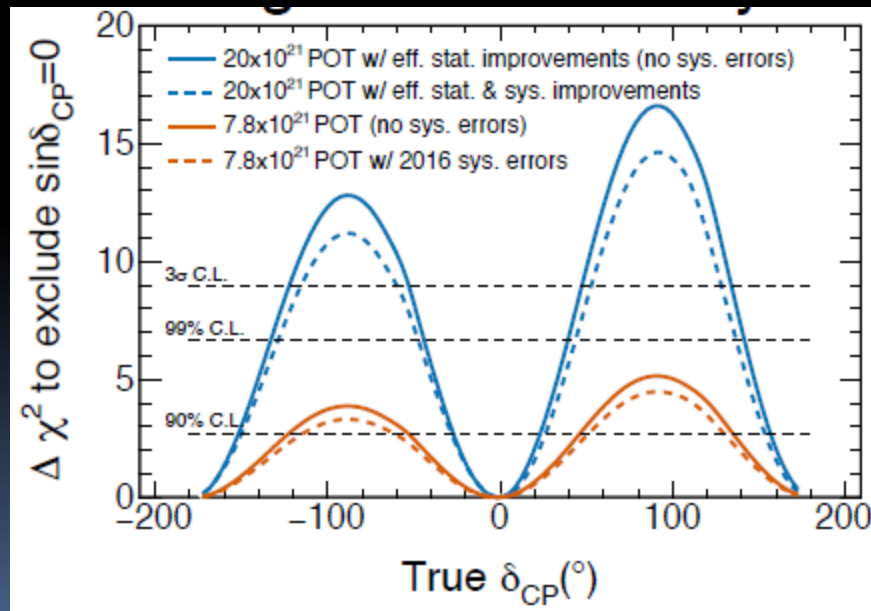


- CP-conservation hypothesis ($\sin\delta_{CP} = 0$ or π) outside 2σ interval
- Confidence intervals, rad: $[-2.49 - -1.23]$ (NH, 1σ) $[-2.98 - -.60]$ (IH, 2σ)
- T2K data favour $\delta_{CP} \sim -\pi/2$ and normal hierarchy

T2K: future plan

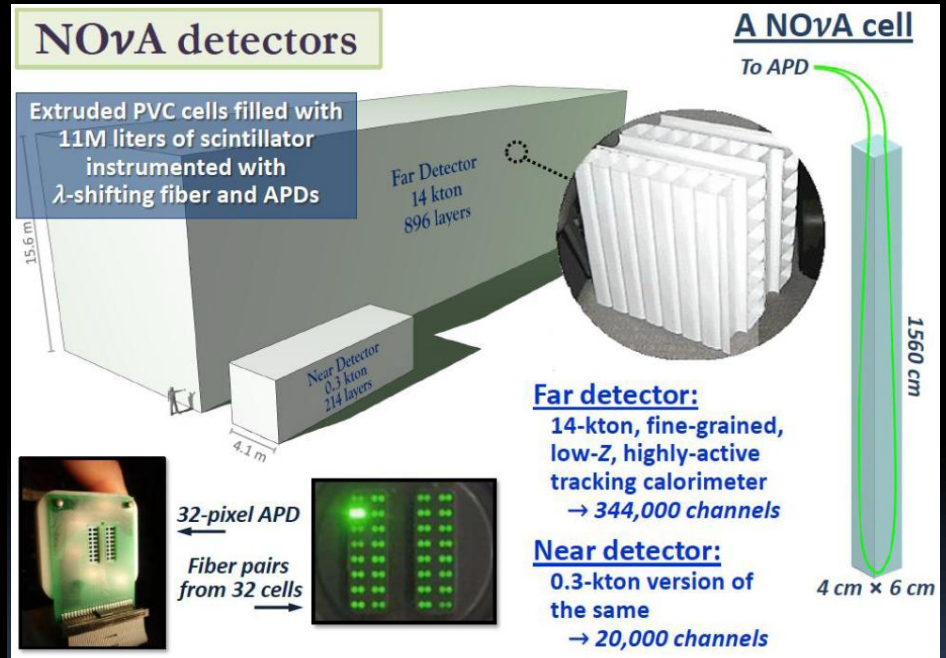
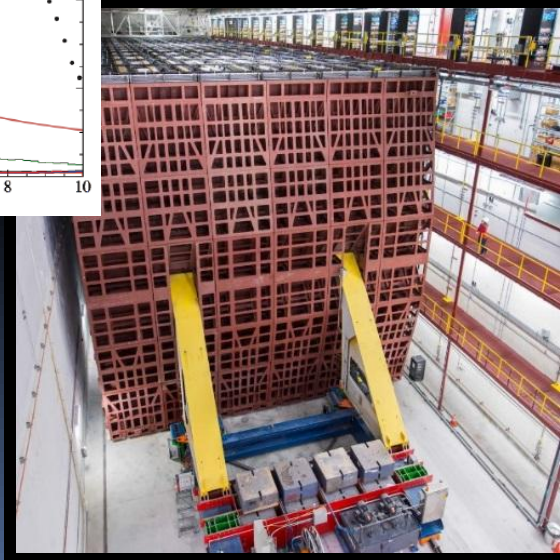
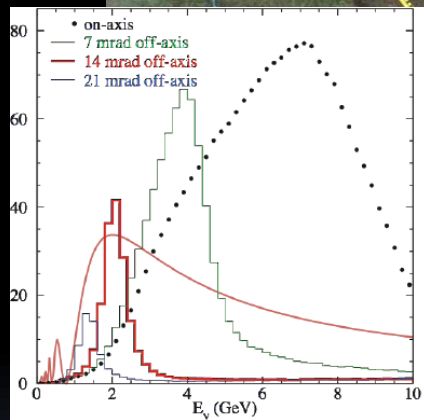
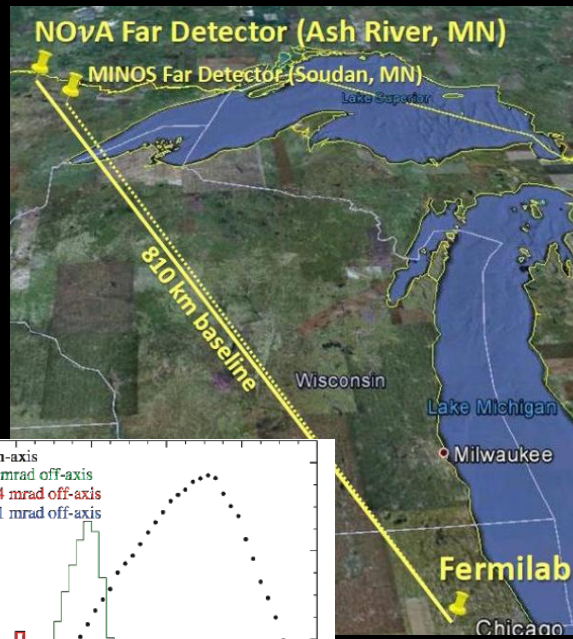
T2K expected to accumulate 7.8×10^{21} POT around 2021

- Upgrade of near detectors to improve systematic uncertainties **18% (2011) \rightarrow 9% (2014) \rightarrow 6% (2016) \rightarrow goal 4% (2020)**
- Plan to increase the beam intensity up to 1 MW in 2021
- Beam power up to 1.3 MW in \sim 2026
- T2K-II: proposed extension up to 2026 for 20×10^{21} POT
 3σ sensitivity to CP violation for $\delta_{CP} \sim -\pi/2$



NOVA

Neutrino beam from FNAL to Ash River
Baseline 810 km
Neutrino beam 14 mrad off-axis
Far detector : 14 kt fine-grained calorimeter
65% active mass
Near Detector: 0.3 kt fine-grained calorimeter



Taking data since Summer 2014
 Study of $\nu_\mu \rightarrow \nu_\mu$ and $\nu_\mu \rightarrow \nu_e$ oscillations

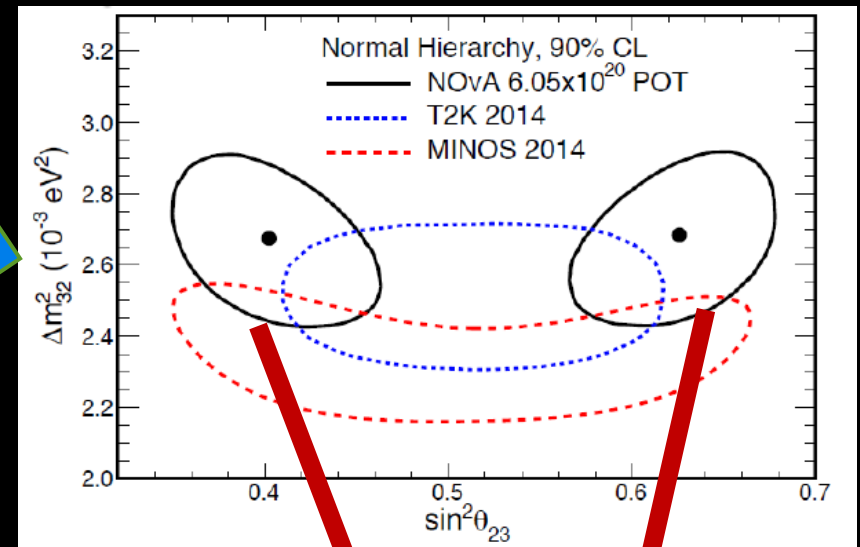
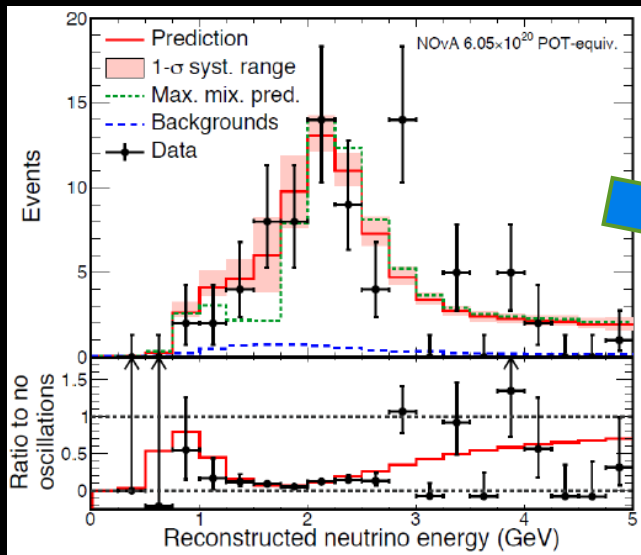
NOvA results

PRL 118 (2017)151802

ν_μ disappearance

473 \pm 30 events expected w/o oscillations
78 events observed

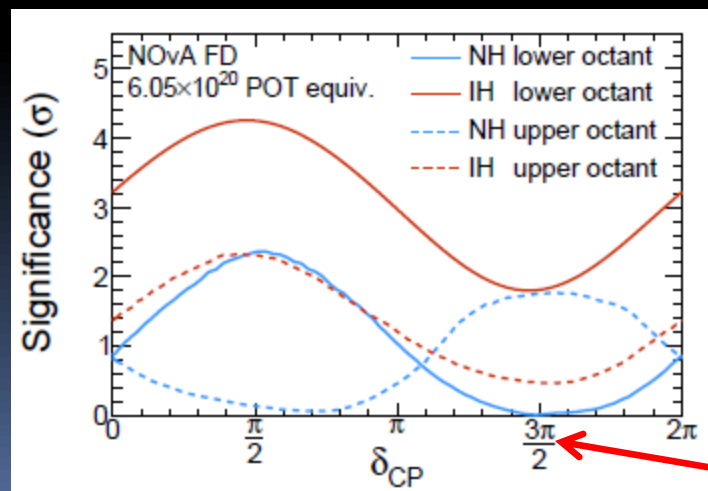
Neutrino mode 6.05x10⁵ POT



PRL 118 (2017) 231801

ν_e appearance

33 events detected
8.2 background



Maximal mixing
disfavored at 2.6 σ

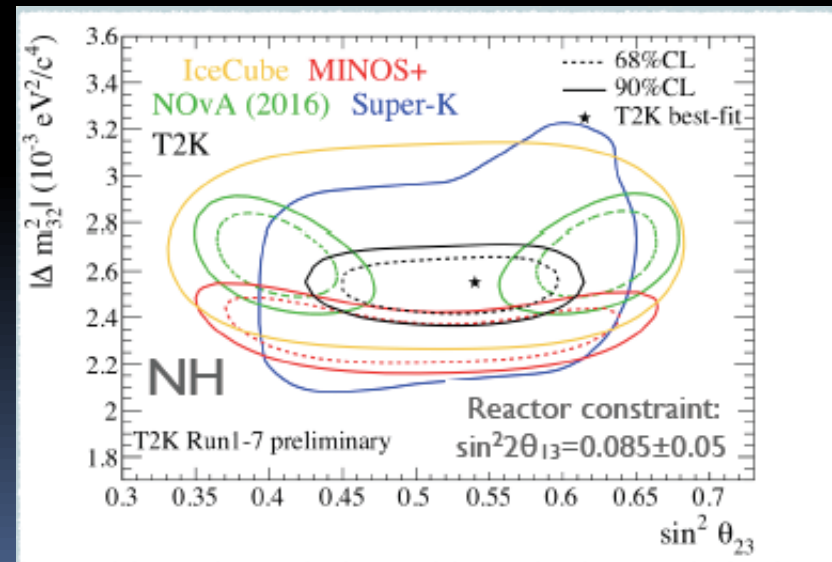
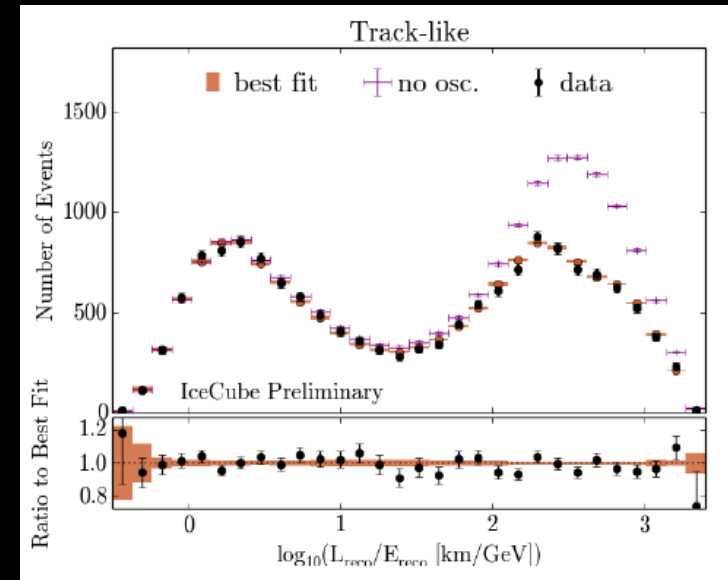
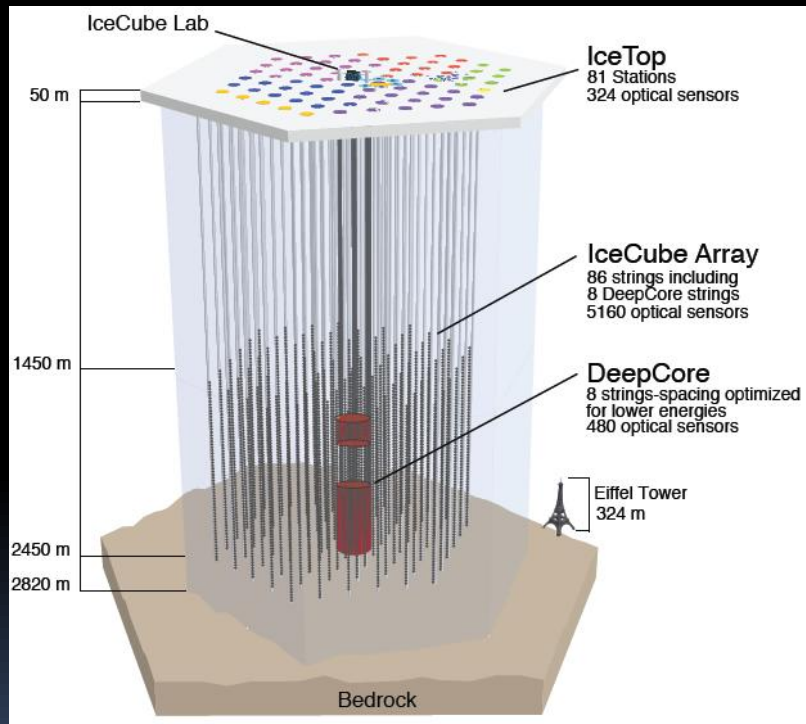
NOvA runs with
anti- ν_μ beam since
February 2017

$-\pi/2$ in T2K notation

IceCube

J.Hignight, talk at CoSSURF 2017

Neutrinos have the first maximum of disappearance at about 25 GeV



Reactor experiments

Daya Bay, China



17.4 GW

RENO, Korea



16 GW

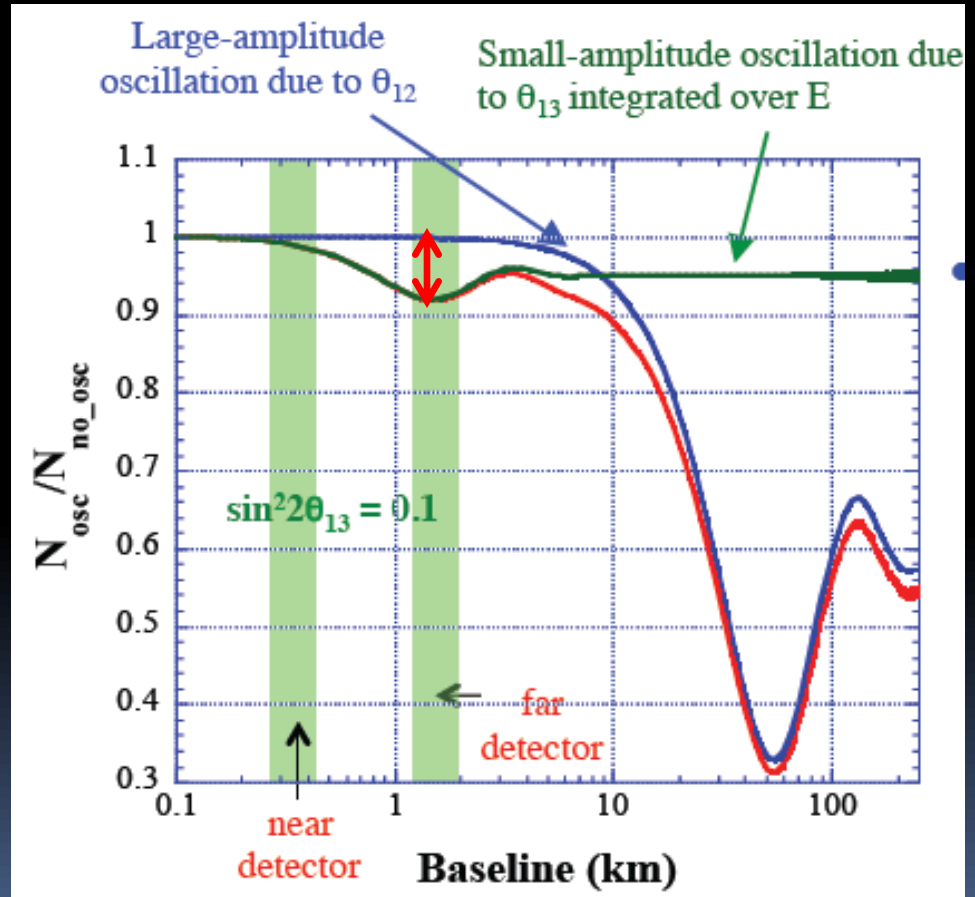
Double Chooz, France



8.5 GW

Measurement of θ_{13}

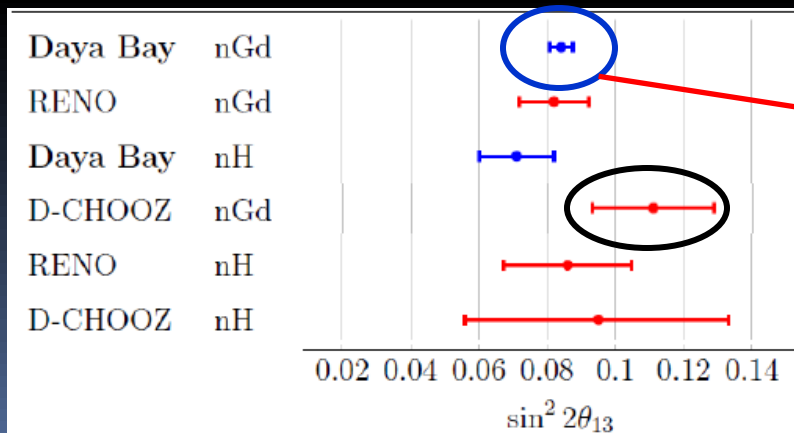
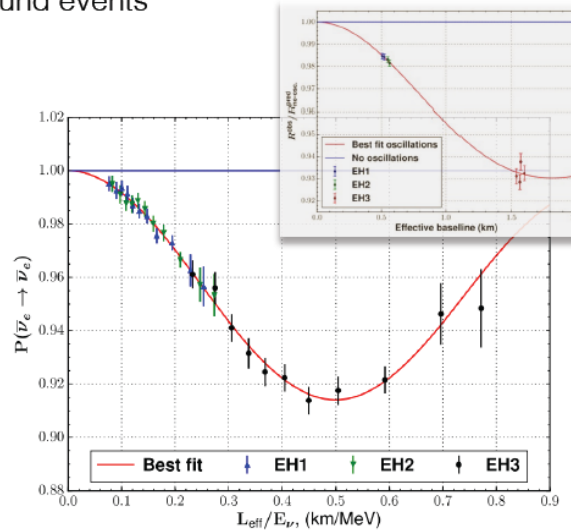
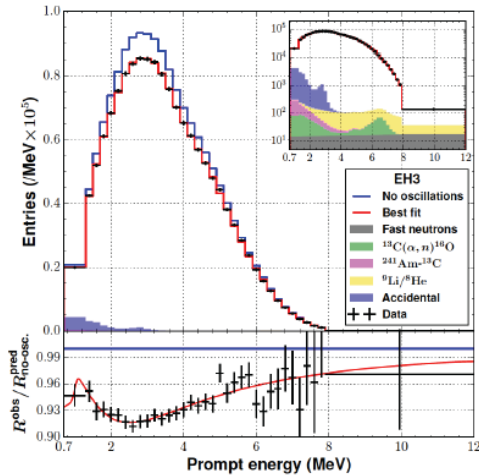
$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{ee}^2 L}{4E} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E} \right)$$



θ_{13} from reactors

- nGd 1230 days
 - 217 days of full 6-AD data set
 - 1013 days of 8-AD data sample
 - Improved energy response model and energy calibration
 - Reduced uncertainties in background events

Daya Bay



**$\sin 2\theta_{13} = 0.085 \pm 0.05$
 $\theta_{13} = 8.4 \text{ deg}$**

Future LBL Projects

- Reactor experiment JUNO
- Accelerator LBL experiment DUNE
- HyperKamiokande and T2HK

Reactor experiment JUNO

China



66 institutions
> 500 collaborators

Start data taking in 2020

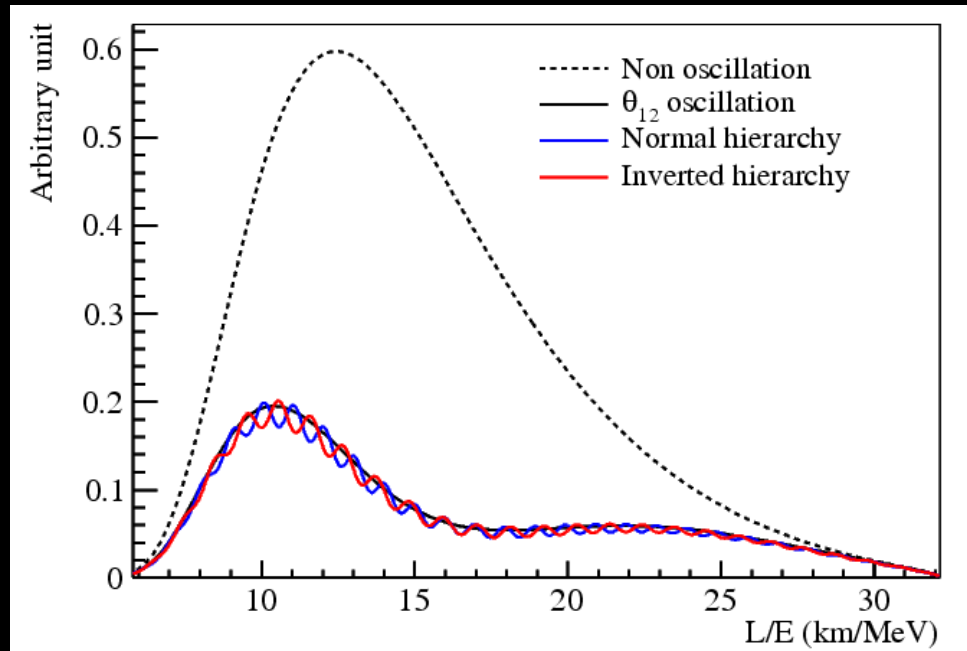
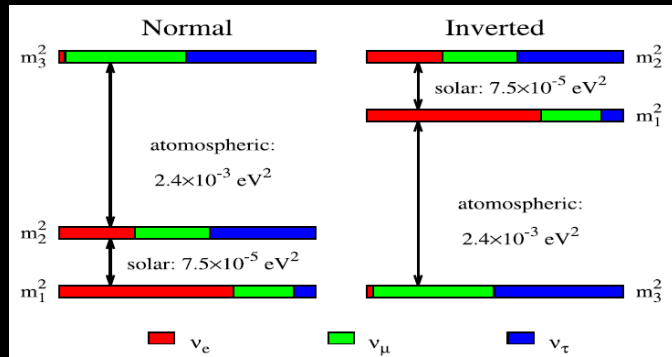
Main target:
Measurement of
neutrino mass hierarchy

- 700 m deep underground
- 36 GW reactor power
- 53 km baseline -> **oscillation maximum θ_{12}**
- 20 kton LS detector
- **3%** energy resolution at 1MeV
- **<1%** energy scale uncertainty

!?

JUNO targets

Main goal: determination of neutrino mass hierarchy



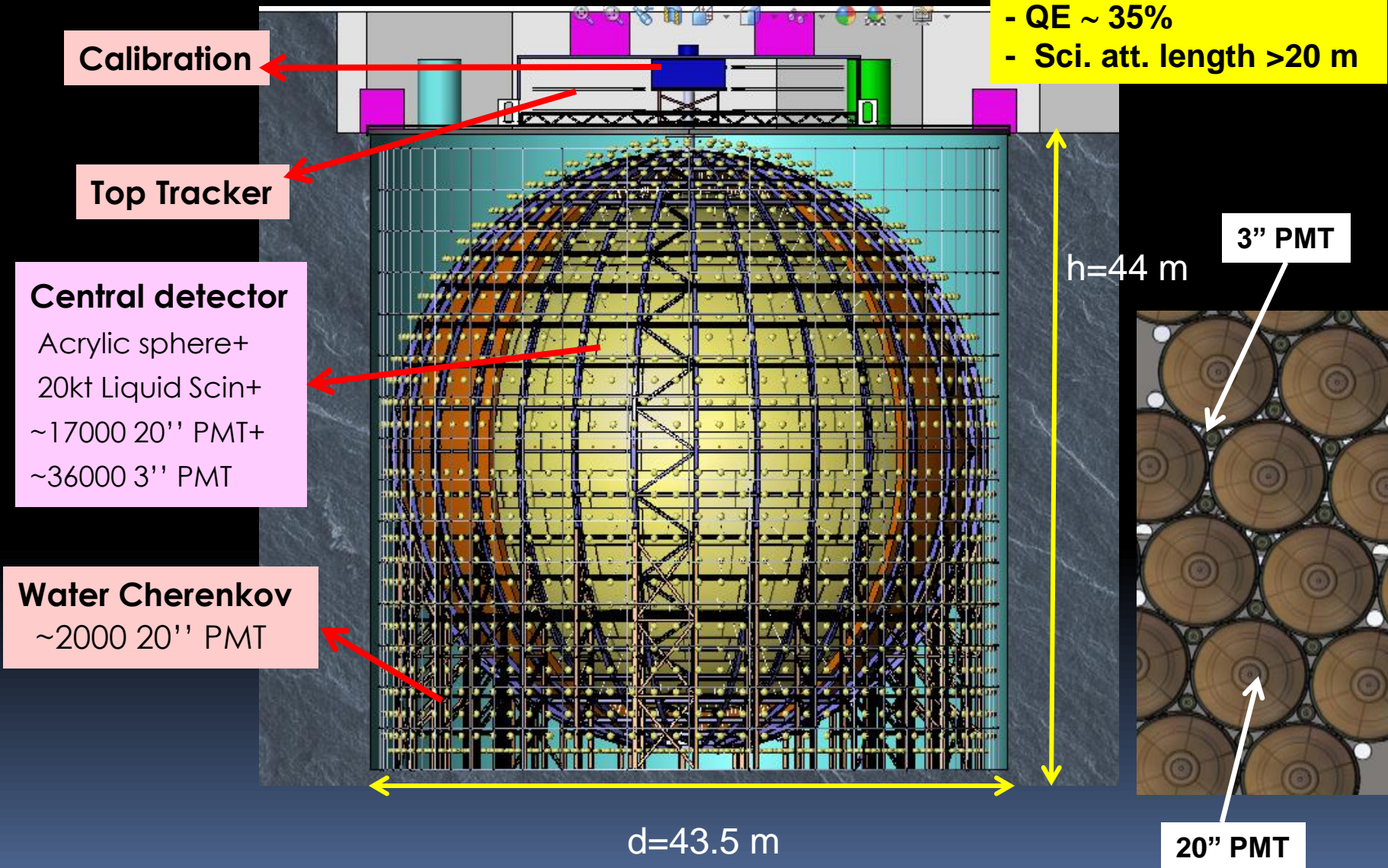
FRD 88, 013008(2013)	Hierarchy discrimination power	With info on $\Delta m_{\mu\mu}^2$ from LBL expts
Statistics only	4σ	5σ
Realistic case	3σ	4σ

Oscillation Parameter	Current accuracy (global 1σ) **	Dominant experiment(s)	JUNO Potentiality
Δm_{21}^2	2.3%	KamLAND	0.59%
$\Delta m^2 = m_3^2 - \frac{1}{2}(m_1^2 + m_2^2) $	1.6%	MINOS, T2K	0.44%
$\sin^2(\theta_{12})$	~4-6%	SNO	0.67%

+ Supernova neutrino
+ Geoneutrinos
+ Solar neutrinos

Detector JUNO

- Requirements:
- PMT coverage 75% of total surface
 - QE ~ 35%
 - Sci. att. length >20 m



LBNF/DUNE Project

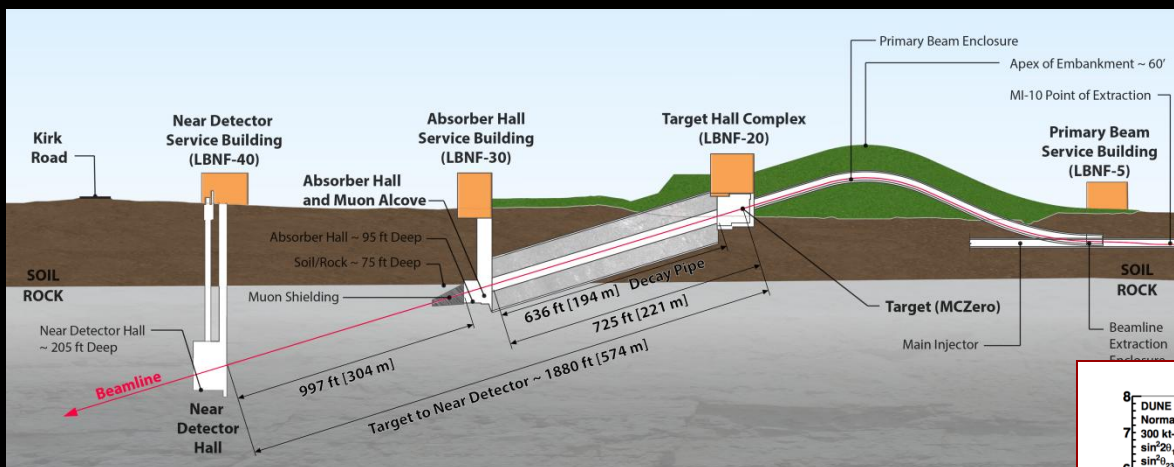
Groundbreaking ceremony at SURF - July 2017

Flagship FNAL project

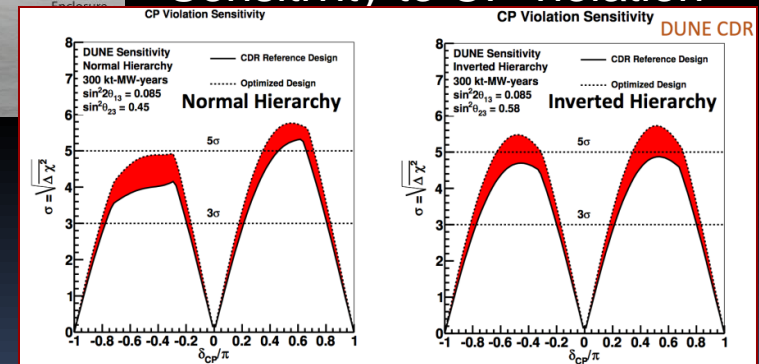
- Main goals:**
- discovery of CP violation in leptonic sector
 - neutrino mass hierarchy at $>5\sigma$ level
 - neutrino astronomy
 - proton decay search

30 countries
161 institutions
 ≥ 1000 collaborators

$E_p = 60-120$ GeV
Beam power 1.2 \rightarrow 2.4 MW
On axis neutrino beam
 $E_\nu \sim 1-6$ GeV
 $L=1300$ km from FNAL to SURF, S.Dakota

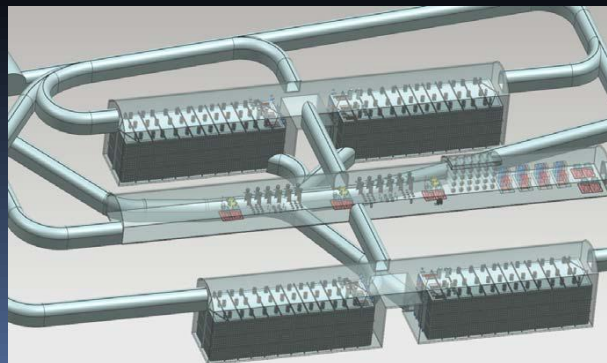


Sensitivity to CP violation



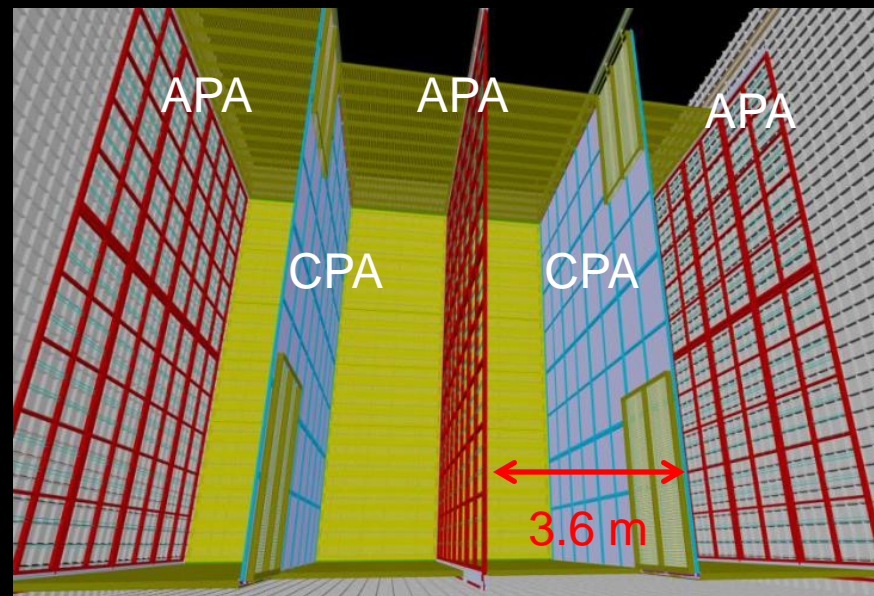
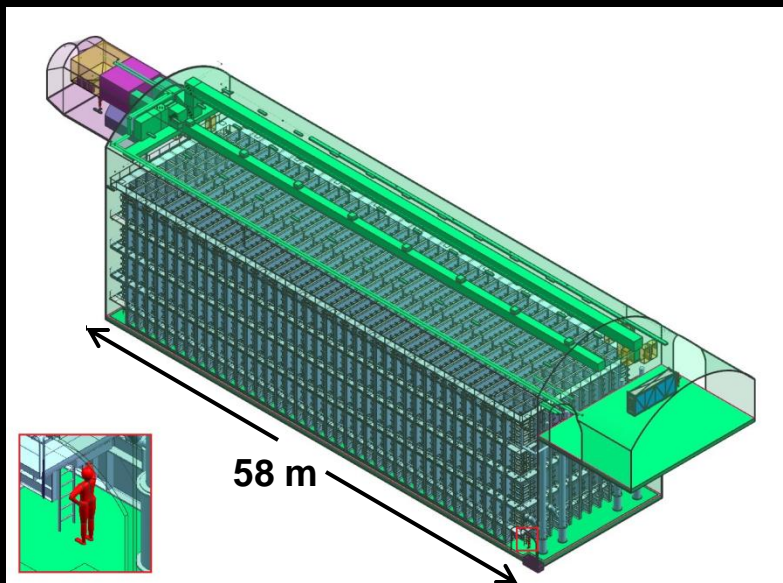
Far detector 40 kt (4 x 10kt) LAr TPC

Single and Dual phase detectors



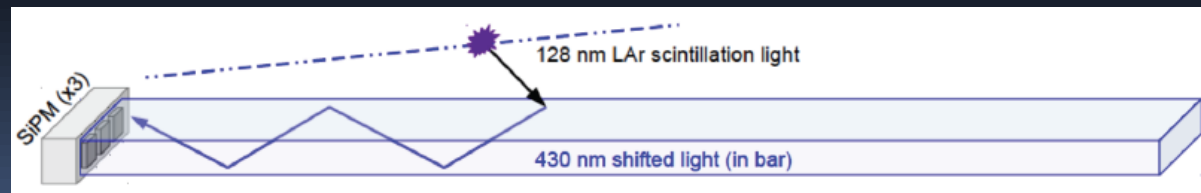
2021 – installation of 1st far detector
2024 – 2 modules operational
2026 – deliver neutrino beam

Single-phase LAr TPC



1st 10 kt module of DUNE - single-phase TPC
6m x 2.3 m anode and cathode planes 3.6 m spacing
Photon detectors – light guides + SiPMs embedded in APAs

J.Insler, talk at LLWI2017



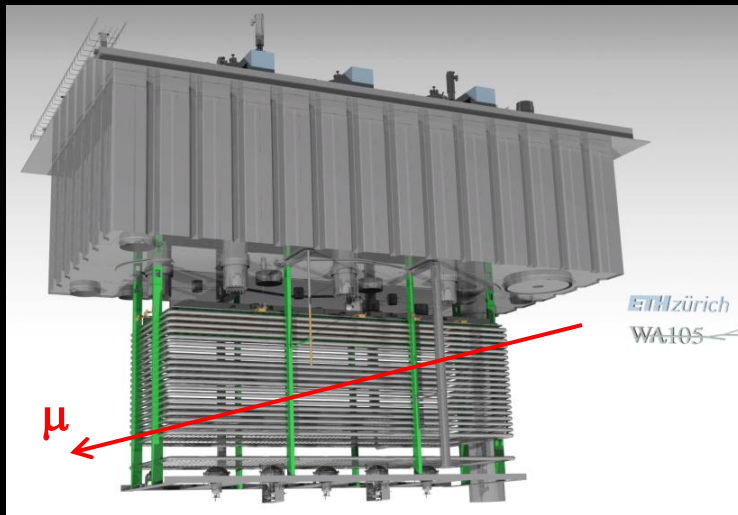
LAr detectors at CERN Neutrino Platform

NP02: WA105

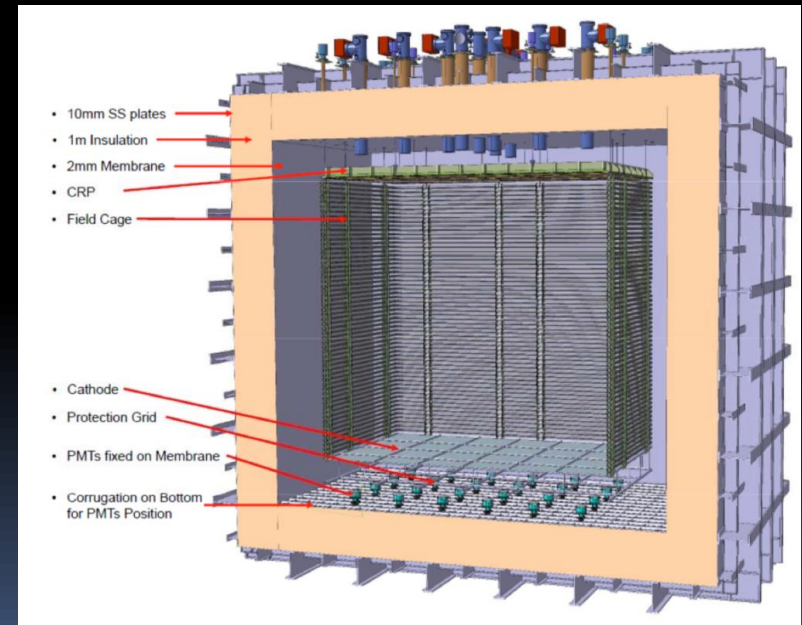
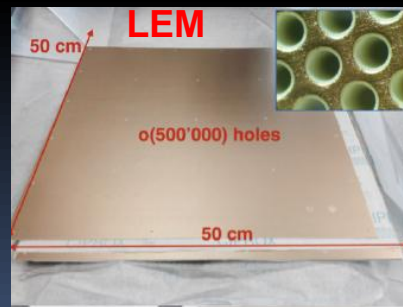
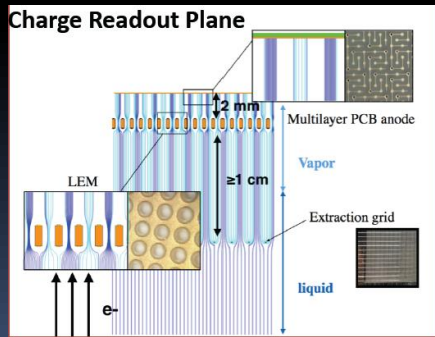
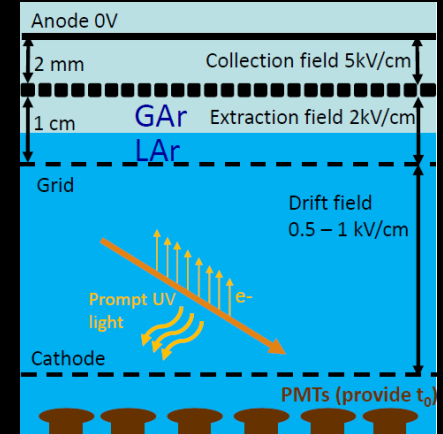
(DP demonstrator + ProtoDUNE DP)

S.Murthy, talk at TPC-2016

Demonstrator: 3x1x1 m³ – 5 tons



ProtoDUNE DP:
6x6x6 m³
300 tons active mass



Cosmic data taking

Measurements with test beam in 2018

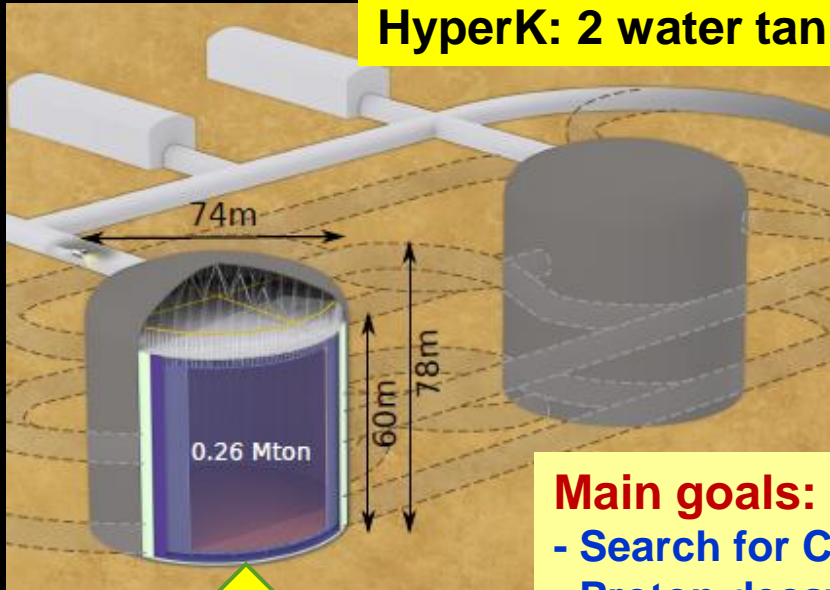
HyperKamiokande

Japan

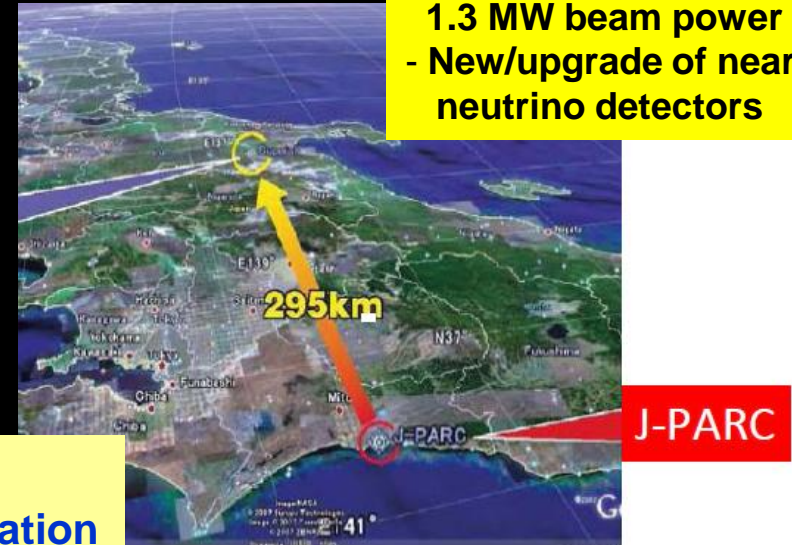
Included in the MEXT
Large Projects Roadmap
in August 2017

12 countries
70 institutes
~300 members
Expected data taking start 2026

HyperK: 2 water tanks



- Upgrade of JPARC to 1.3 MW beam power
- New/upgrade of near neutrino detectors



Main goals:

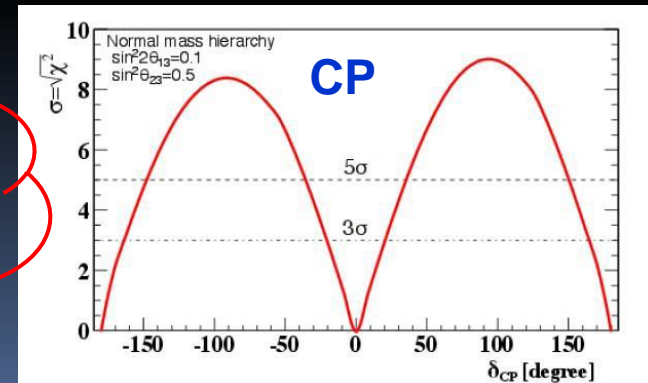
- Search for CP violation
- Proton decay
- Neutrino astrophysics

1 tank

60 m(H)x74m(D)
Total volume 260 kt
Fiducial volume 190 kt
~10xSuperK
PMT coverage 40%
40000 PMTs

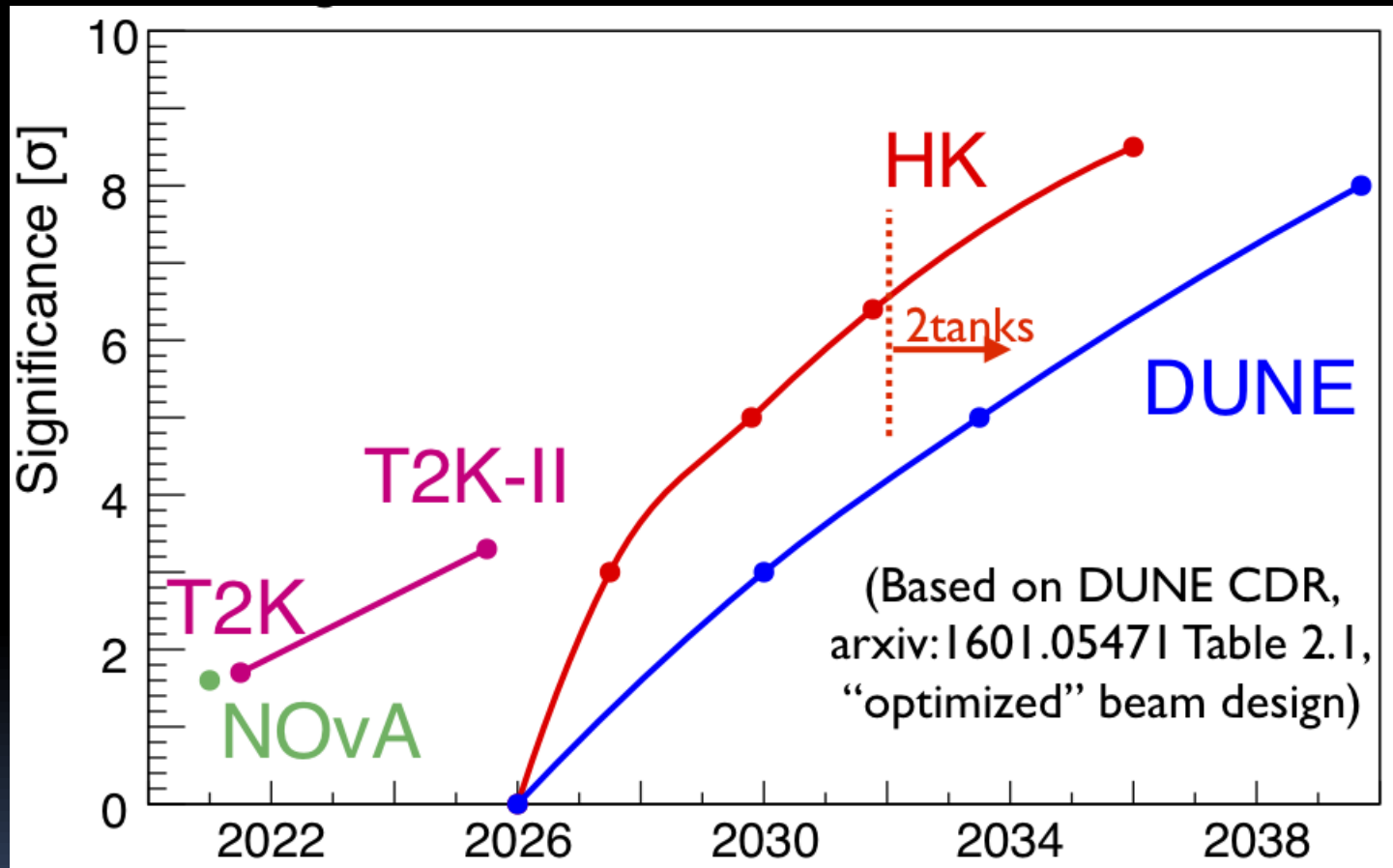
10 years of running:

- 8σ for $\delta_{CP} = -\pi/2$
- 80% coverage of δ_{CP} parameter space with $>3\sigma$
- $p \rightarrow \pi^0 e^+ > 10^{35} \text{ y}$



Expected sensitivity to CP

Significance for $\delta_{CP} = -\pi/2$



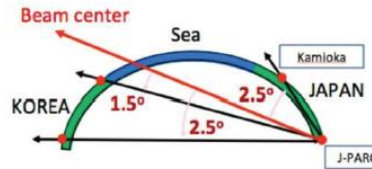
T2HKK

Second tank in Korea

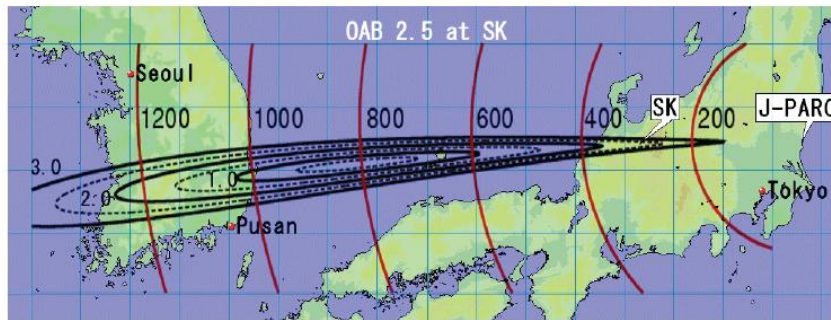
arXiv:1611.06118

Build second tank in Korea to enhance mass hierarchy and δ_{CP} sensitivities

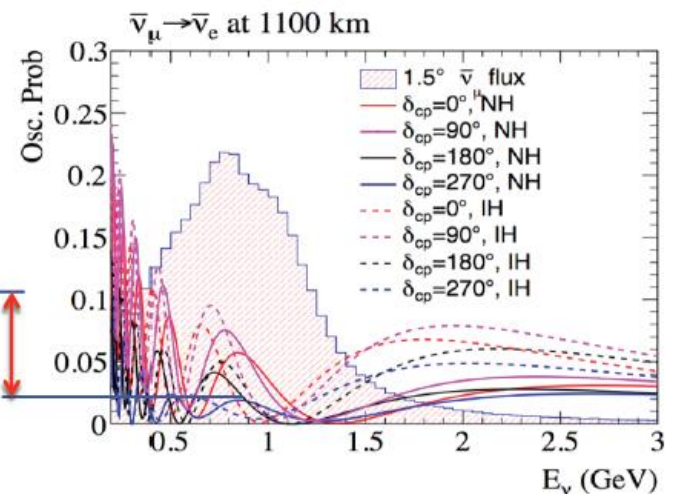
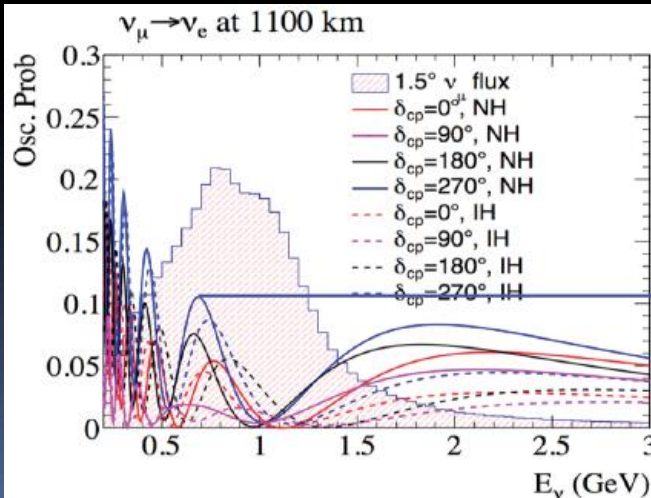
- 1000 – 1200 km baseline
- 1.3° – 3.0° off axis beam direction



Neutrino and antineutrino spectra in T2HKK cover 1st and 2nd oscillation maximum

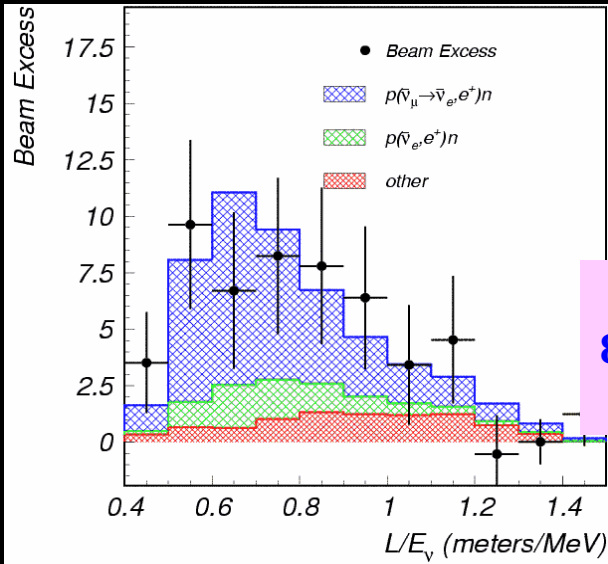


- $A_{CP} \sim 3$ times larger in 2nd maximum
- Sensitivity to MH

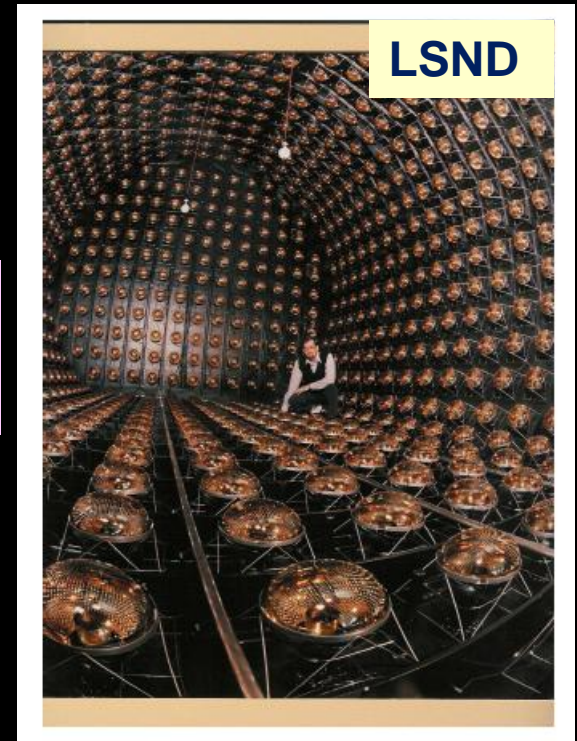
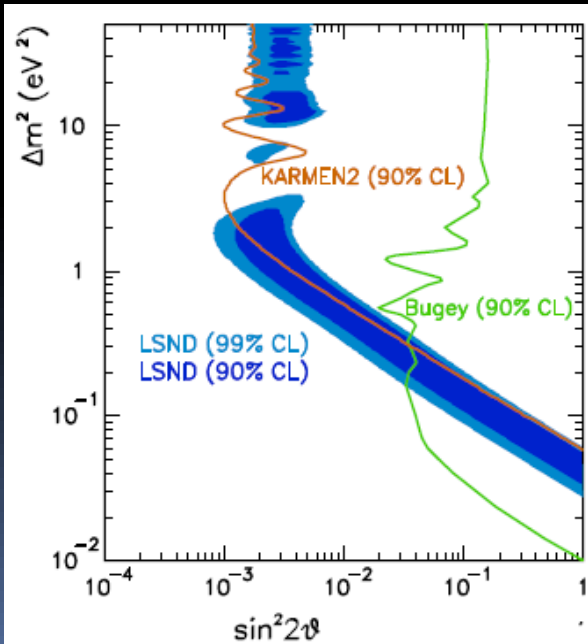


Sterile neutrinos

LSND



anti- $\nu_{\mu} \rightarrow$ anti- ν_e
 $87.9 \pm 22.4 \pm 6.0$ events
 Excess 3.8σ

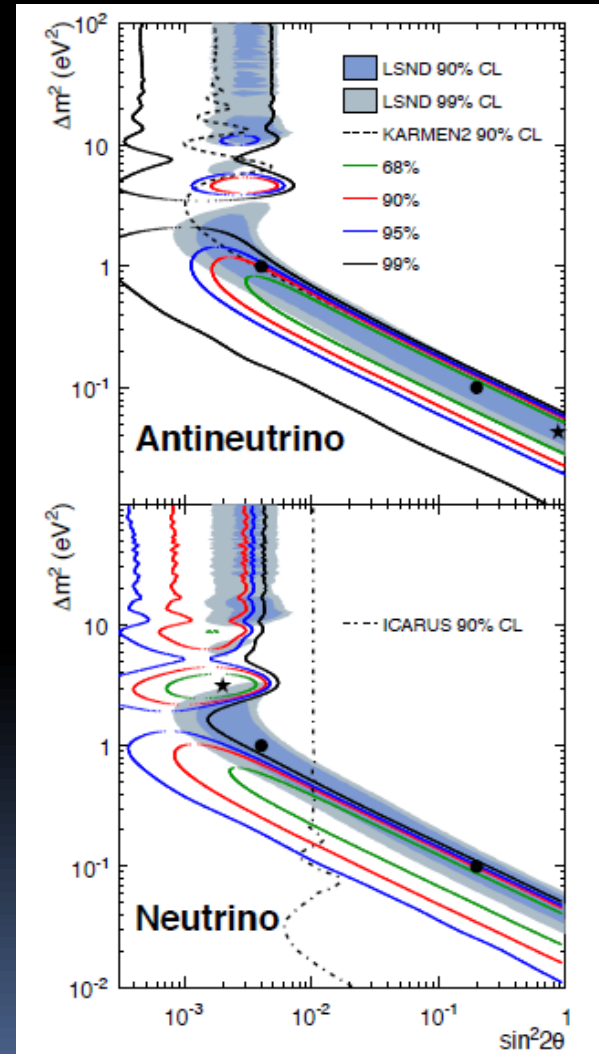
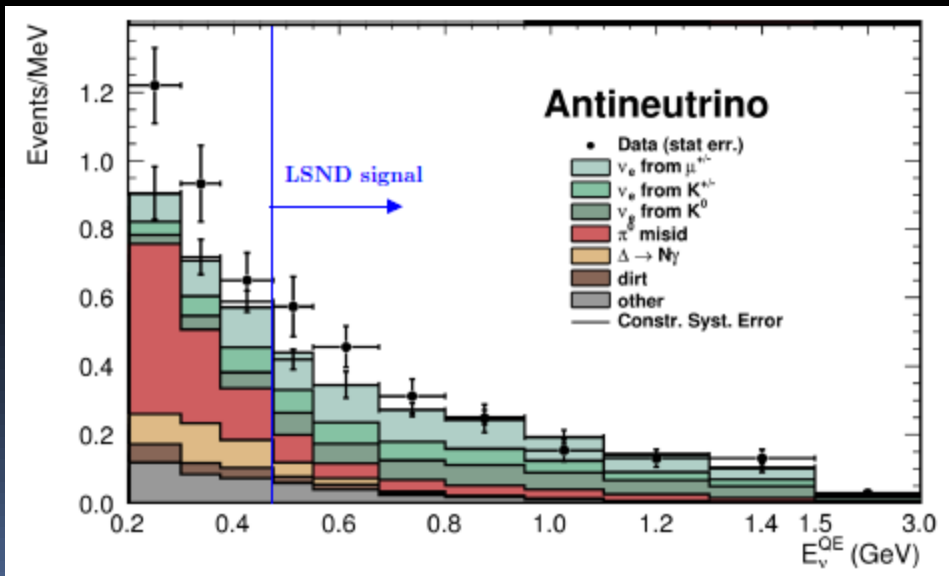
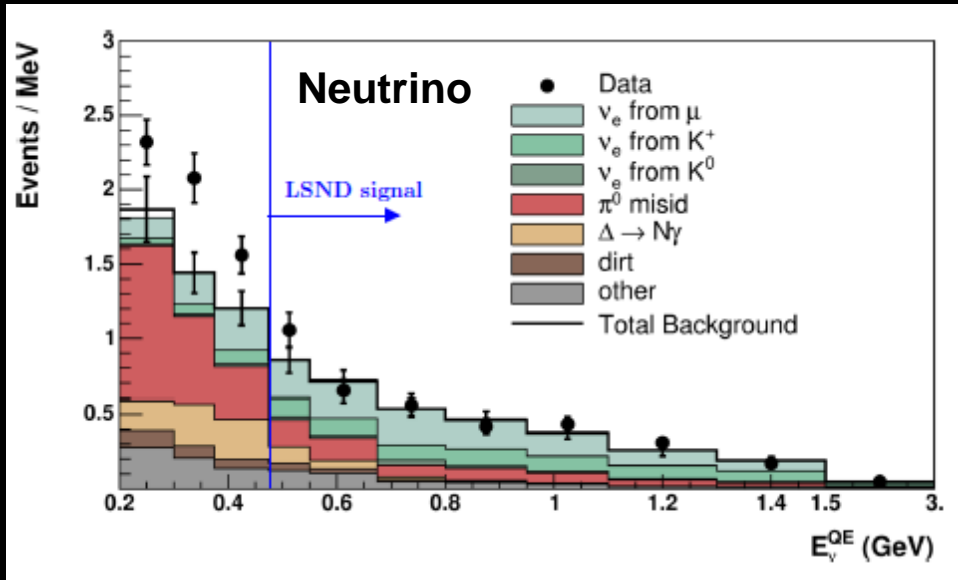


$$\begin{aligned}
 P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e) &= \sin^2(2\theta) \sin^2\left(\frac{1.27 L \Delta m^2}{E}\right) \\
 &= 0.245 \pm 0.067 \pm 0.045 \%
 \end{aligned}$$

$$0.2 \leq \Delta m^2 \leq 1 \text{ eV}^2 \quad 2 \times 10^{-3} \leq \sin^2 2\theta \leq 4 \times 10^{-2}$$

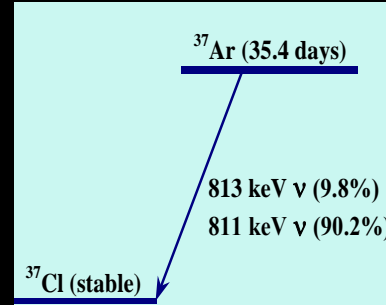
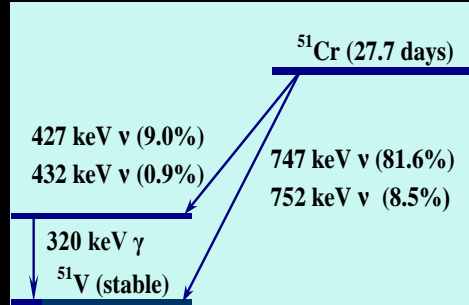
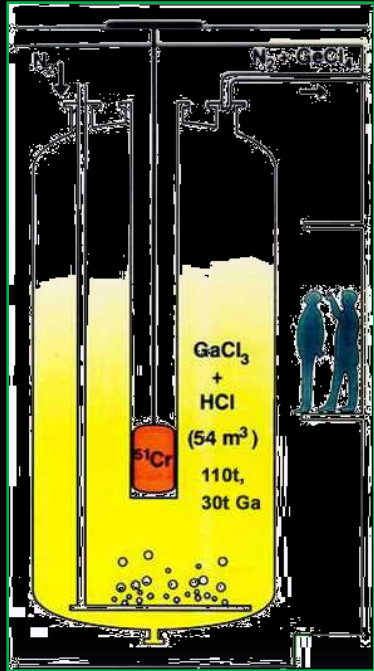
MiniBooNe

$\nu_\mu \rightarrow \nu_e$
 $\text{anti-}\nu_\mu \rightarrow \text{anti-}\nu_e$
 $L \approx 540 \text{ m}$ $E_\nu = 0.2\text{-}3 \text{ GeV}$

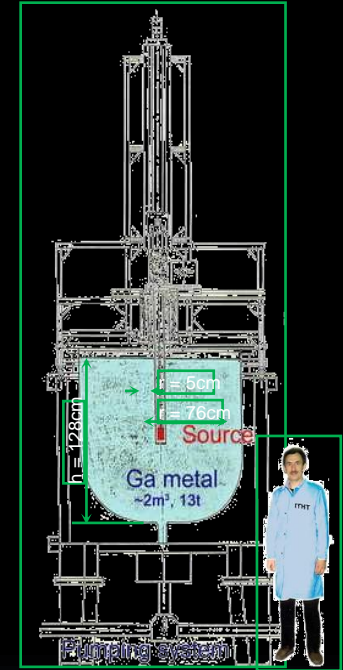


Gallium anomaly

GALLEX

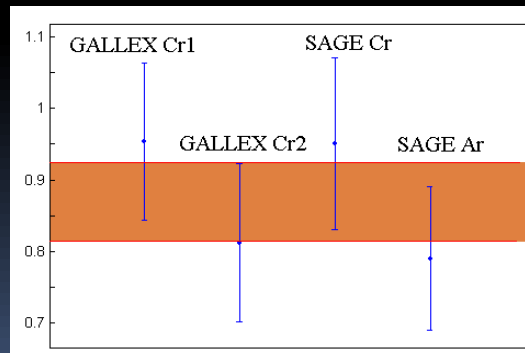


SAGE



Detection process: $\nu_e + {}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge} + e^-$

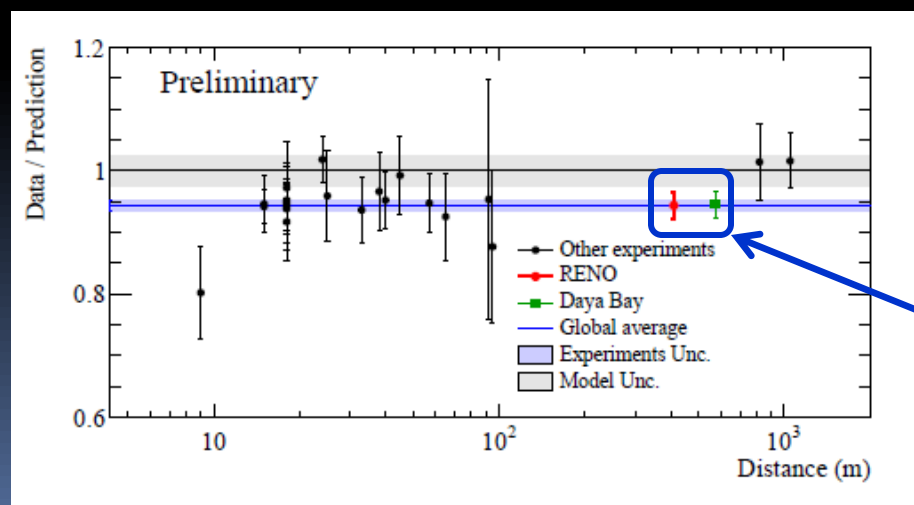
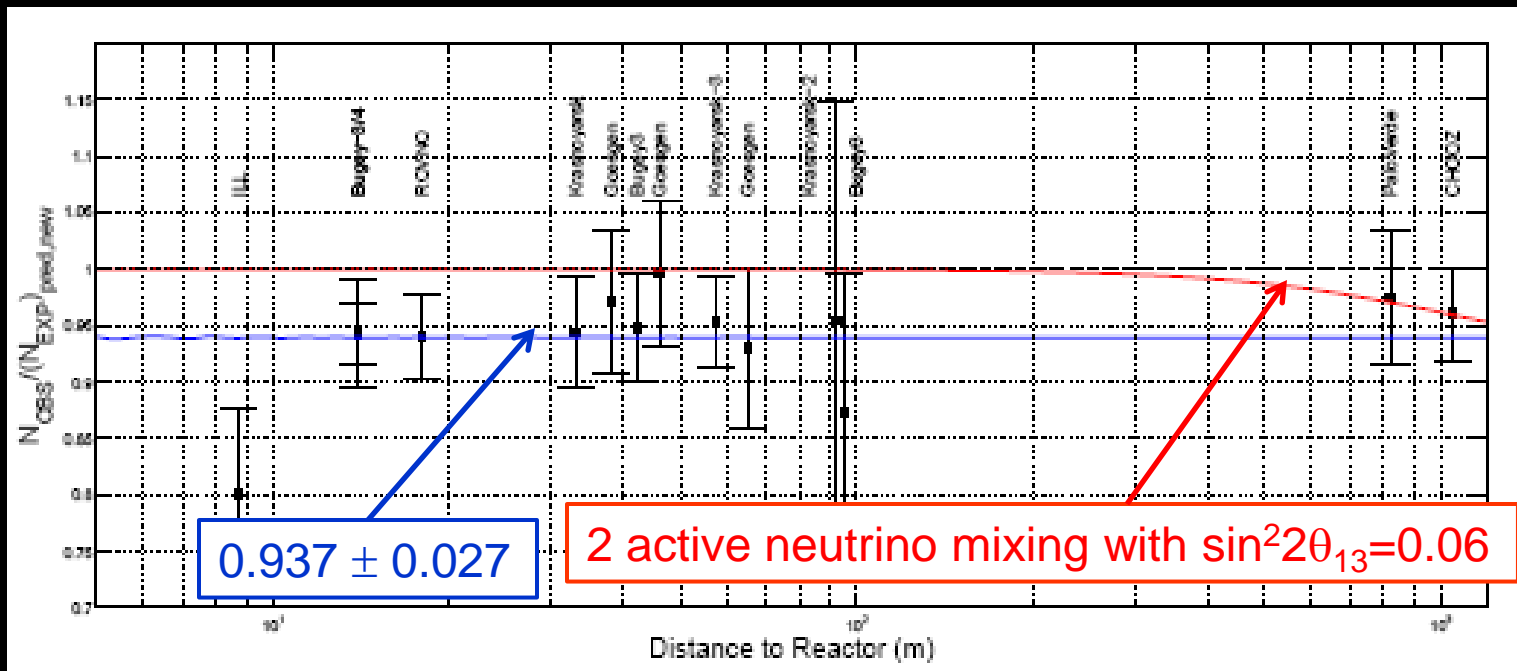
	GALLEX m(Ga)=30 t		SAGE m(Ga)=13 t	
Source	${}^{51}\text{Cr}$ -1	${}^{51}\text{Cr}$ -2	${}^{51}\text{Cr}$	${}^{37}\text{Ar}$
Intensity (Mci)	1.714	1.868	0.517	0.409
$R = (p_{\text{exp}}/p_{\text{theory}})$	0.95 ± 0.11	0.81 ± 0.11	0.95 ± 0.12	0.79 ± 0.10
R_{comb}	0.88 ± 0.08		0.86 ± 0.08	



$$R = p_{\text{exp}}/p_{\text{theory}} = 0.87 \pm 0.05$$

Reactor anomaly

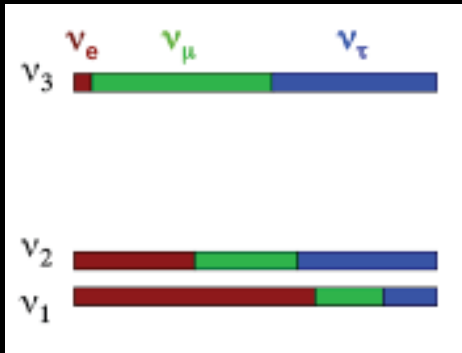
anti- $\nu_e \rightarrow \text{anti-}\nu_e$



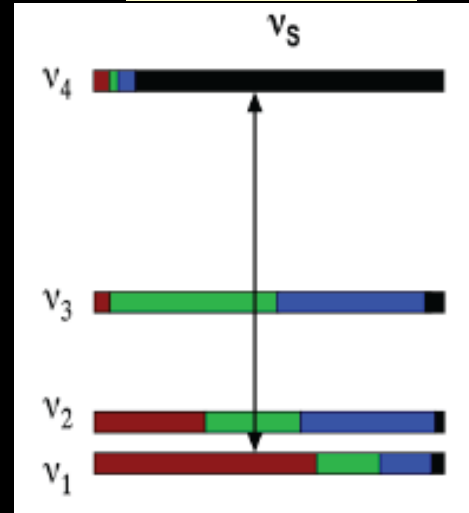
S.Gariazzo et al
arXiv:1703.08860

Sterile neutrino?

3ν



3ν + 1s



$\Delta m_{14}^2 \sim 1 \text{ eV}^2$

?



PNMS matrix

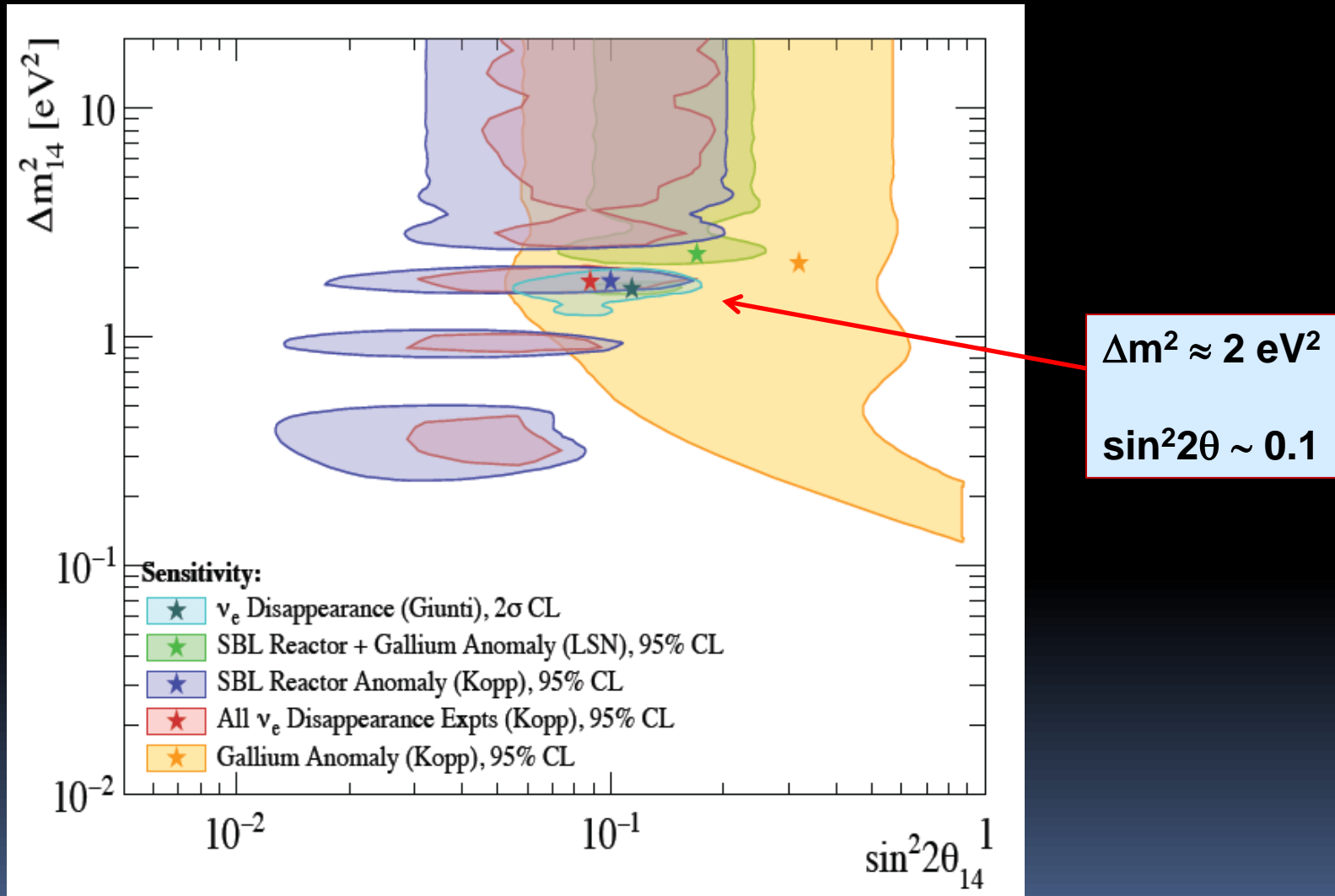
$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{bmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s1} & U_{s1} & U_{s1} \end{pmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{bmatrix}$$

$$\begin{aligned} |U_{e4}|^2 &= \sin^2 \theta_{14} \\ |U_{\mu4}|^2 &= \sin^2 \theta_{24} \cdot \cos^2 \theta_{14} \\ |U_{\tau4}|^2 &= \sin^2 \theta_{34} \cdot \cos^2 \theta_{24} \cdot \cos^2 \theta_{14} \end{aligned}$$

ν_e and anti- ν_e disappearance

arXiv:1512.02202

Global fit of reactor and Gallium data



Hunting for sterile neutrinos

Accelerator

MINOS

SBN at FNAL

Reactor *(running or under construction)*

Daya Bay, RENO, Double Chooz

DANSS

Neutrino-4

NEOS

STEREO

Solid

PROSPECT

NuLat

Neutrino sources

BEST

SOX

Atmospheric neutrinos

SuperKamiokande

IceCube

Sterile ν 's: Daya Bay + MINOS + Bugey-3

- Daya Bay data

- Constrains Δm_{41}^2 (mainly 10^{-4} to 10^{-1} eV^2) and $\sin^2 2\theta_{14}$

- Bugey-3 data

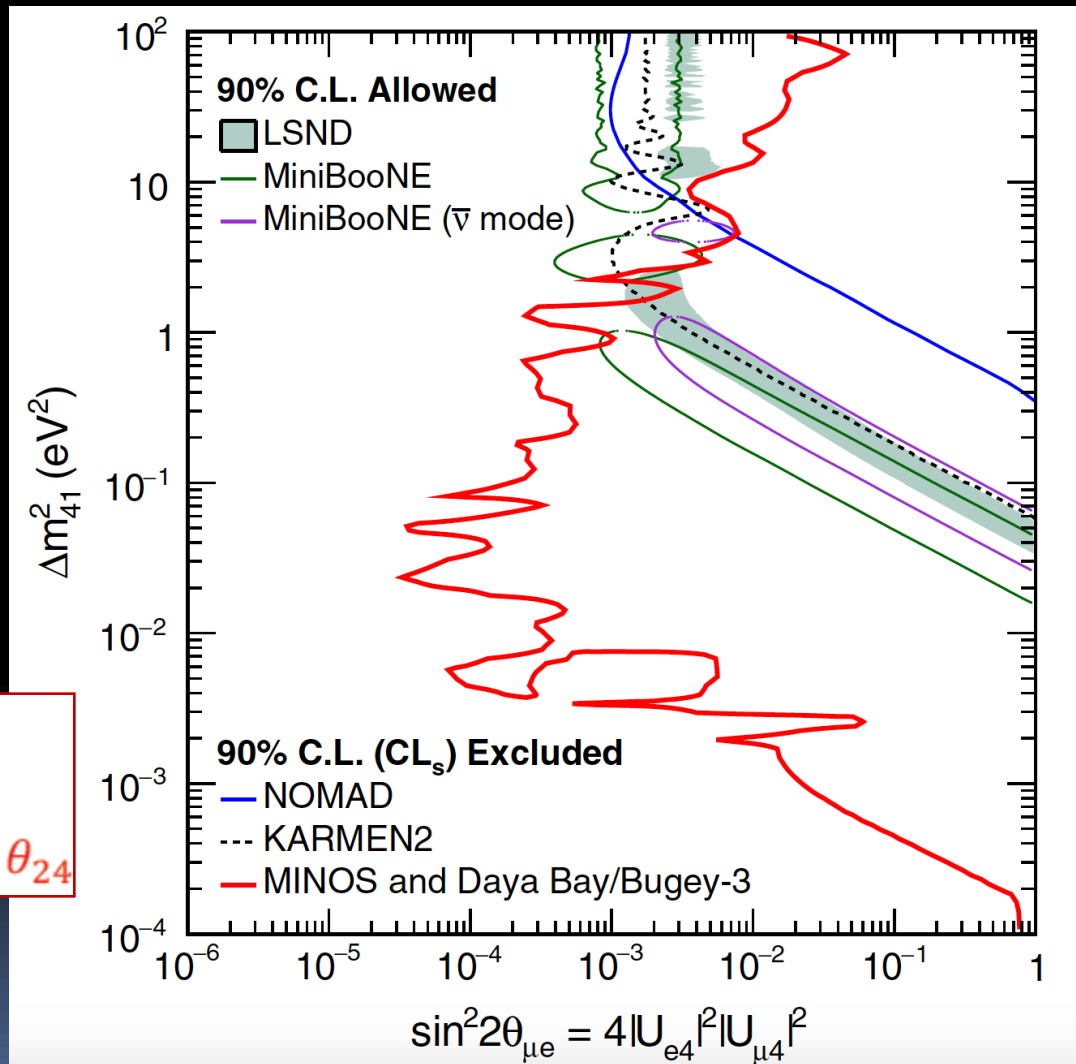
- constrains Δm_{41}^2 (mainly 10^{-1} to 10 eV^2) and $\sin^2 2\theta_{14}$

- MINOS data

- Constrains Δm_{41}^2 (mainly 10^{-3} to 10^2 eV^2) and $\sin^2 \theta_{24}$

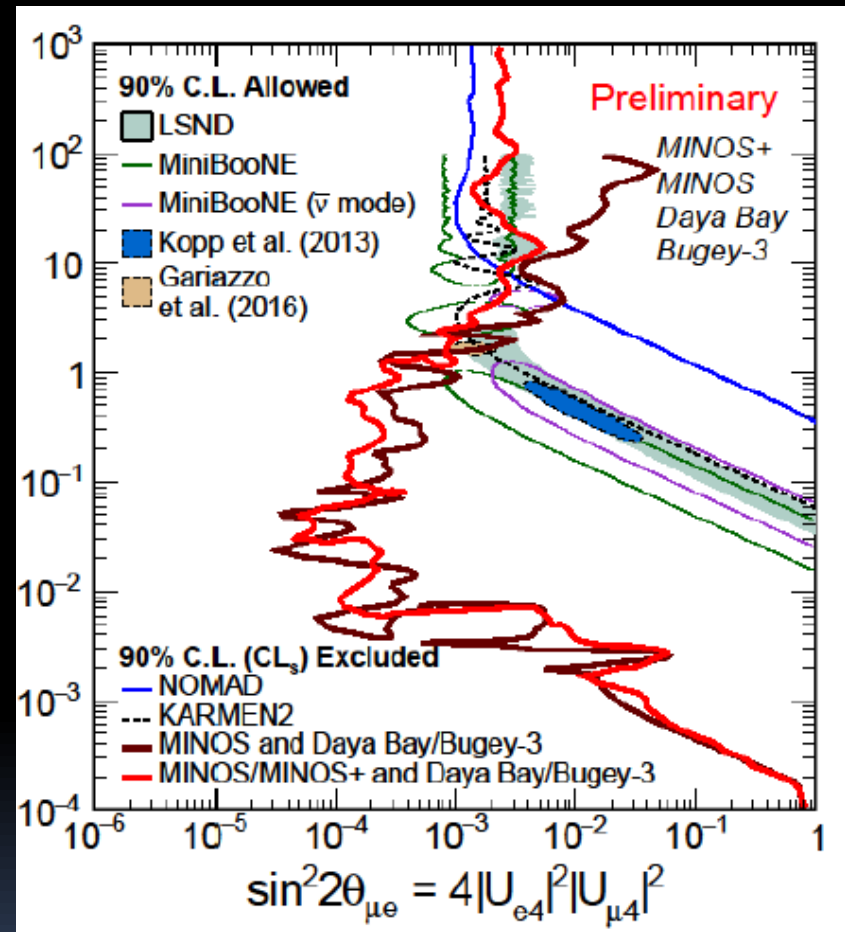
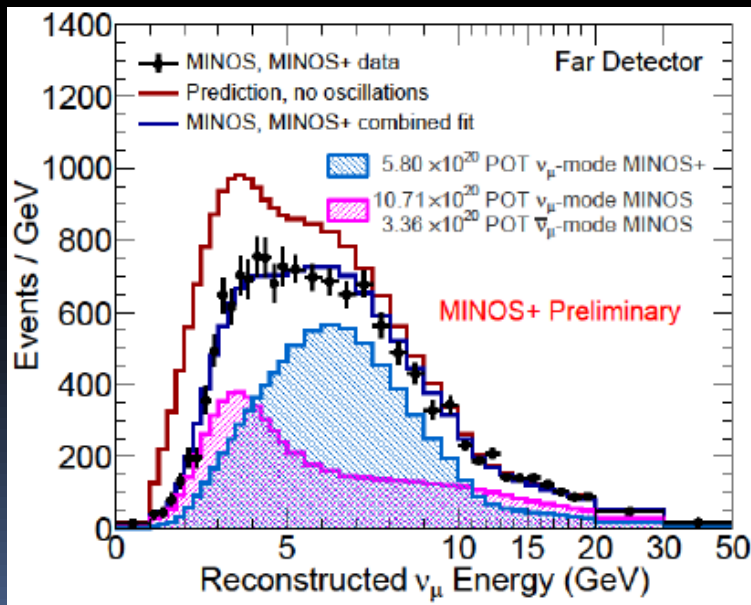
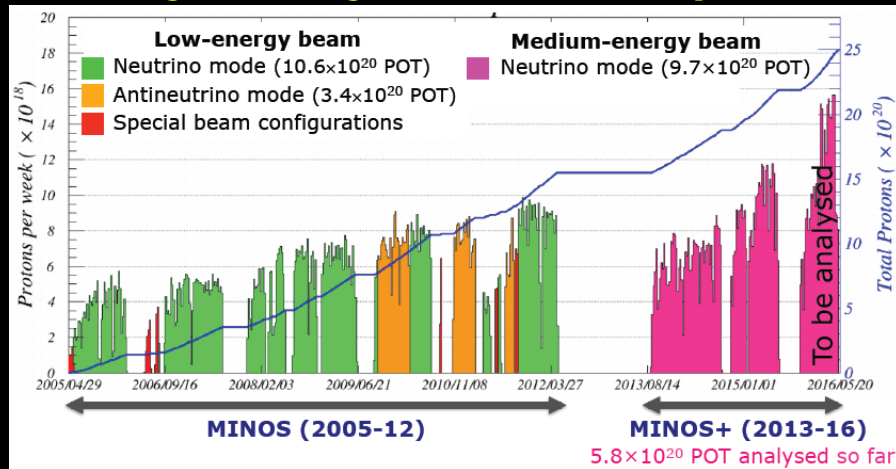
- Combined all three

- Constrains Δm_{41}^2 and $\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \cdot \sin^2 \theta_{24}$



Sterile ν 's: Daya Bay + MINOS/MINOS+ + Bugey-3

A.Blake, talk at
NuFact 2017,
September 2017



**No evidence for 3+1
sterile neutrino oscillations**

Sterile ν 's: IceCube

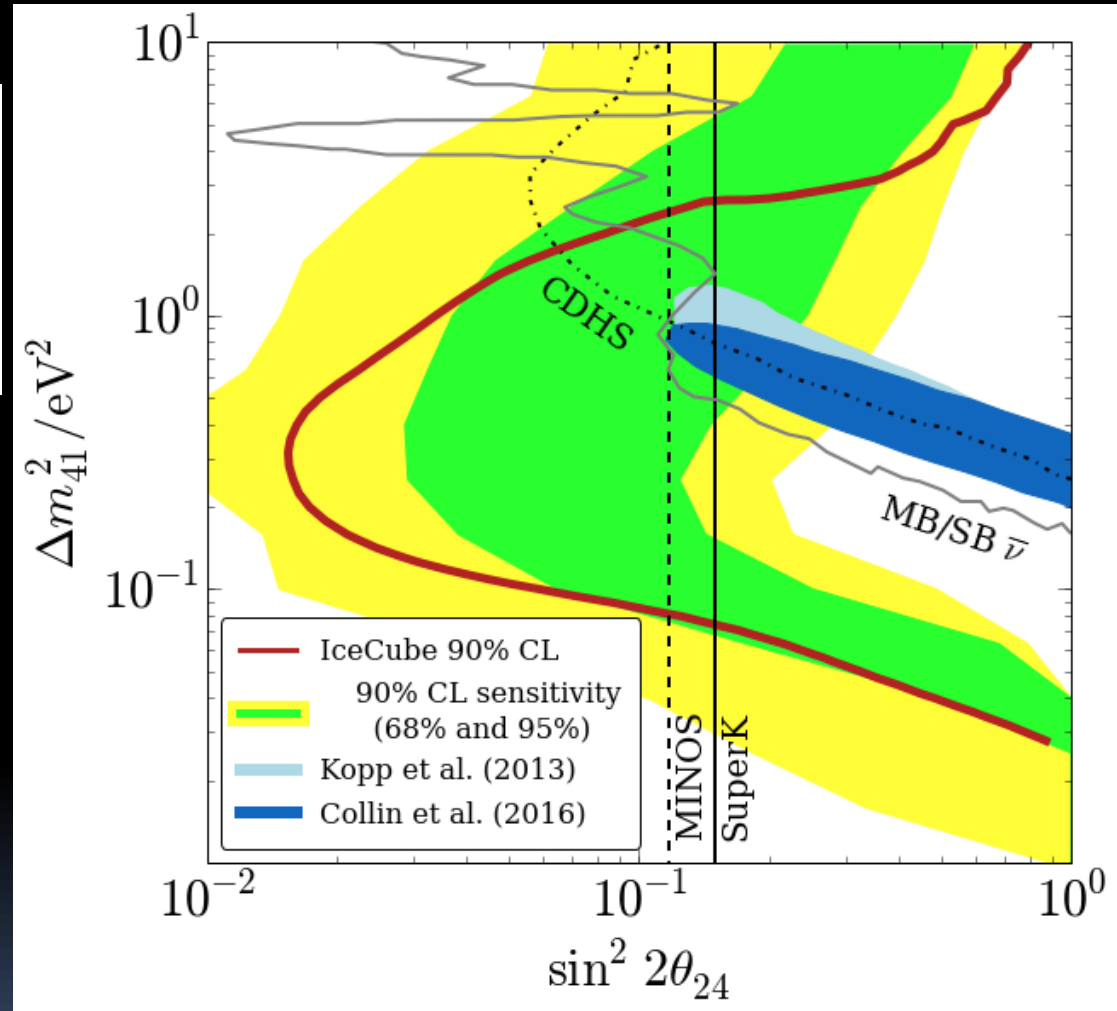
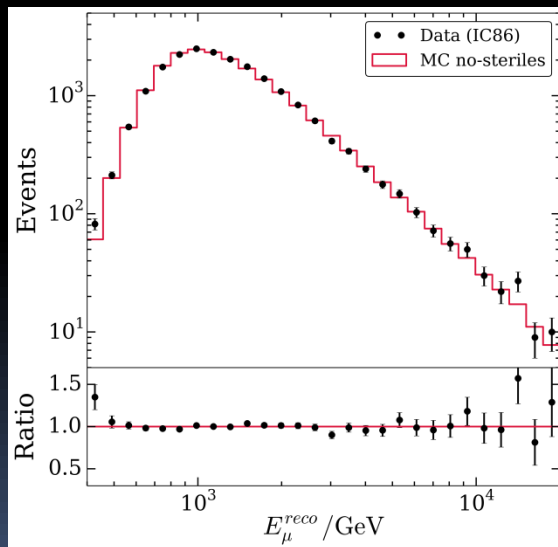
PRL 117 (2016) 071801

$E_\nu = 320 \text{ GeV} - 20 \text{ TeV}$

sterile neutrinos produce distortions of $\nu_\mu + \text{anti-}\nu_\mu$ flux (energy and angle) in the range

$$0.01 \leq \Delta m^2 \leq 10 \text{ eV}^2$$

- 1 year of data
- statistics limited



Result compatible with no-sterile hypothesis

NEOS: reactor anti- ν disappearance

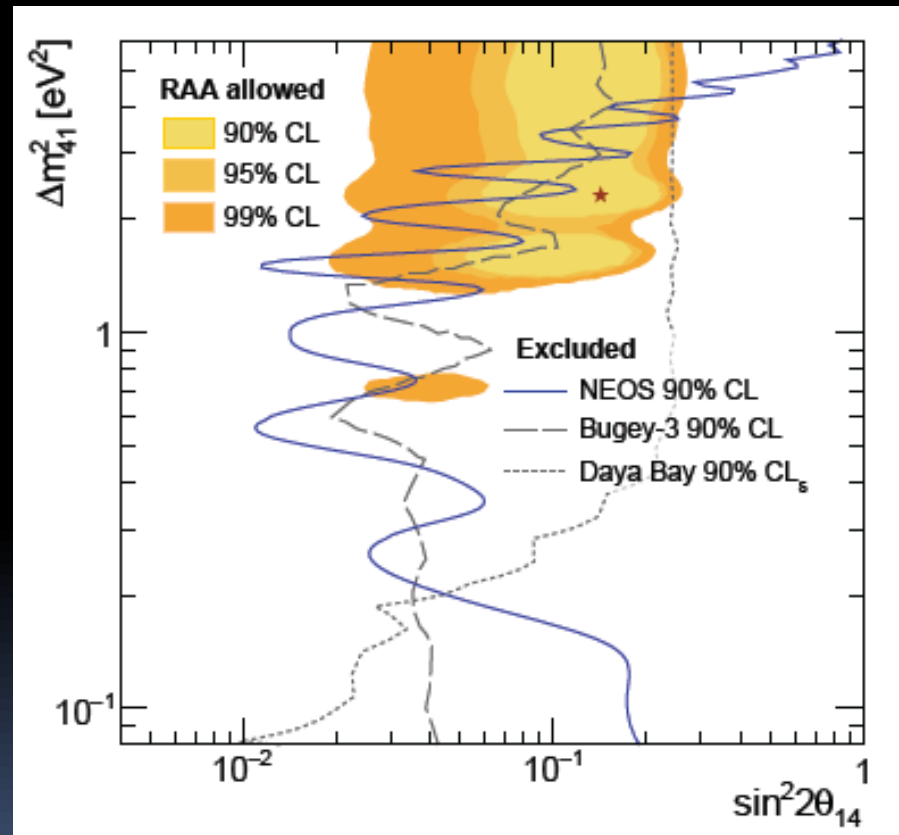
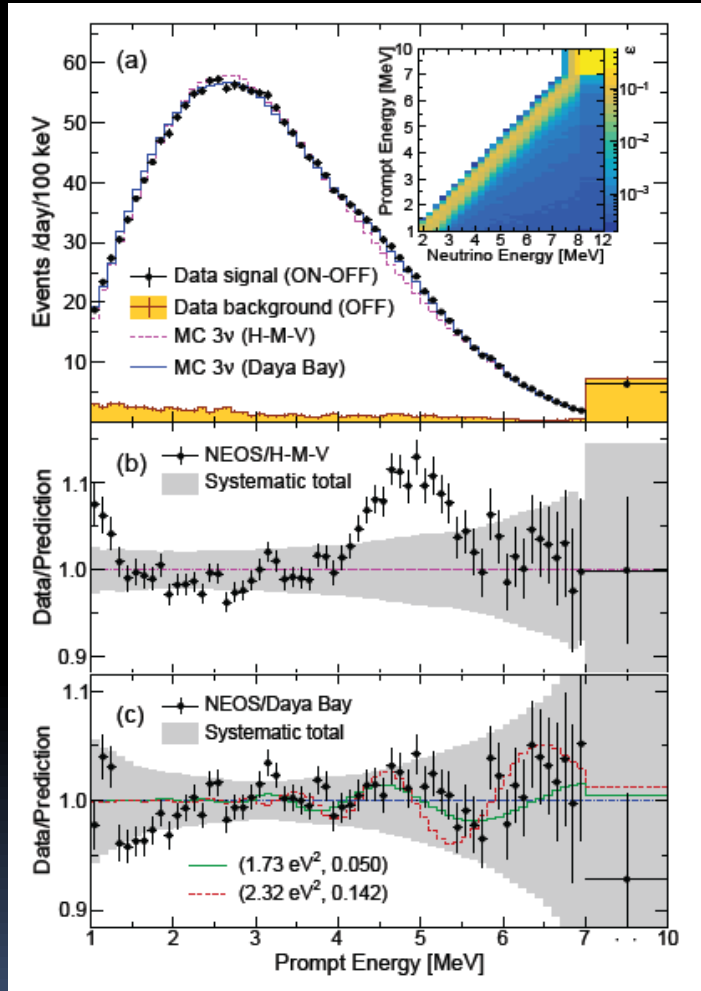
PRL 118 (2017) 121802

Korea, Reactor 2.8 GW

Core: \varnothing 3.1 m h=3.8 m

Detector 1t LS + Gd, 24 m from reactor core

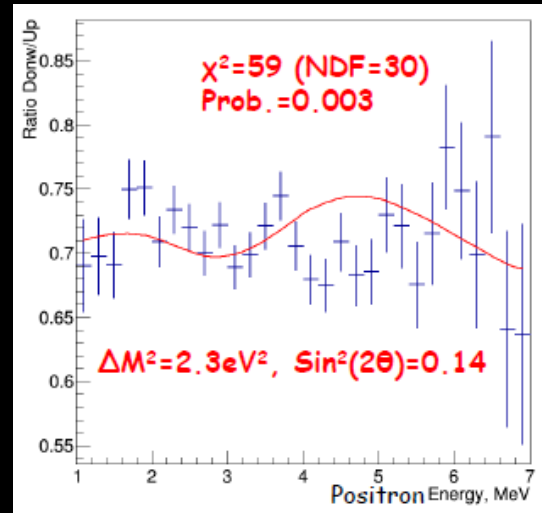
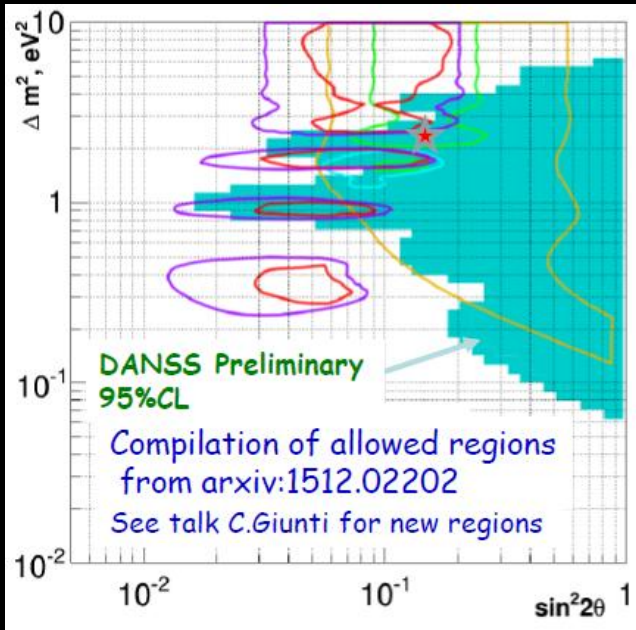
S/N \sim 22



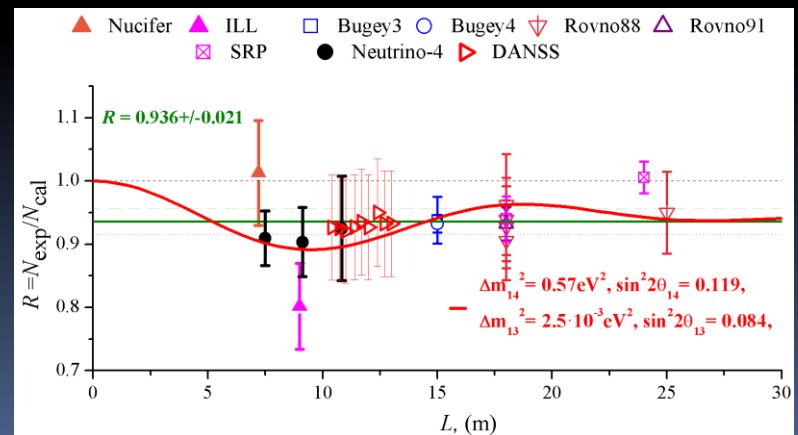
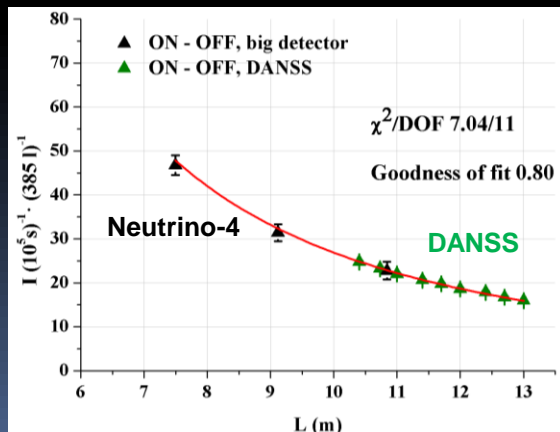
No evidence for sterile neutrino with mass \sim 1 eV

DANSS and Neutrino-4

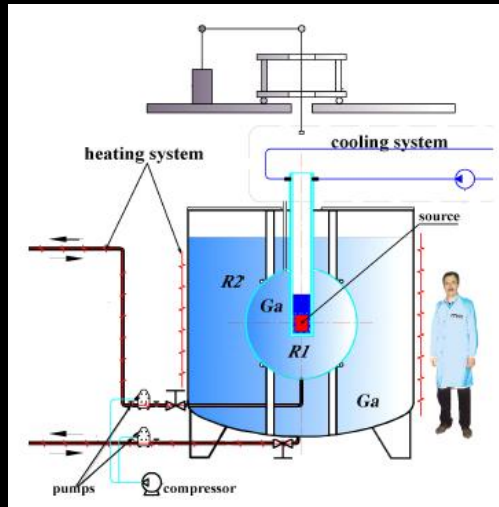
M.Danilov, talk at 52 Moriond EW 2017



R. Samoilo, talk at ICPPA 2017



Source experiments

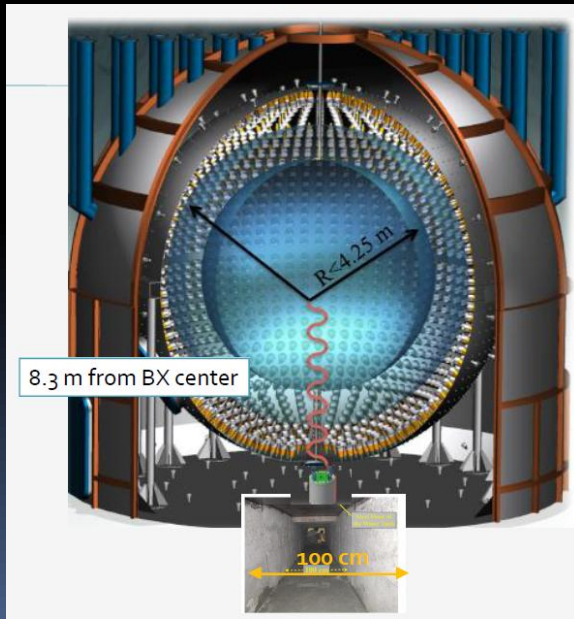
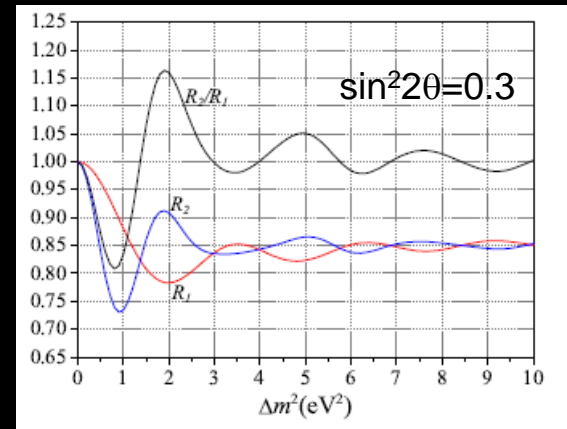


BEST

3 MCi ^{51}Cr source

Two-zone 50 t
liquid Ga metal target

J.Phys.Conf.Ser. 798 (2017) 012113



SOX

Ultra-low radioactive background

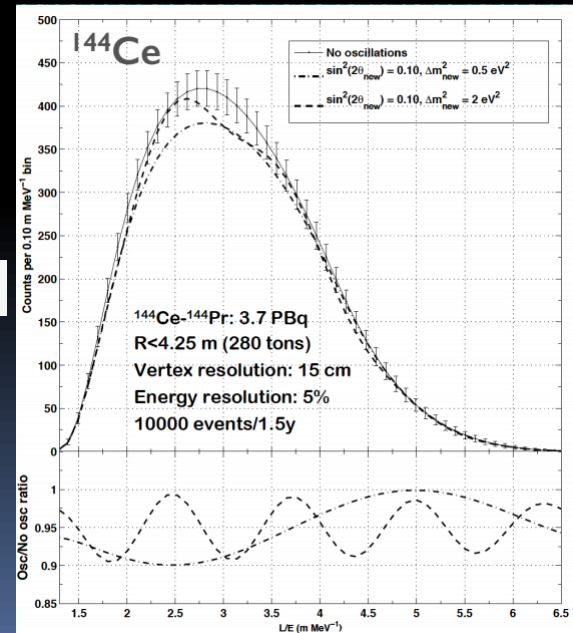
- Spatial resolution: 12 cm @ 2 MeV
- Energy resolution: ~3,5% @ 2 MeV

^{144}Ce - ^{144}Pr $\bar{\nu}_e$ source (100-150 kCi)

Source will be produced
at Mayak, Russia

Start data taking in 2018

PRD 91 (2015) 072005

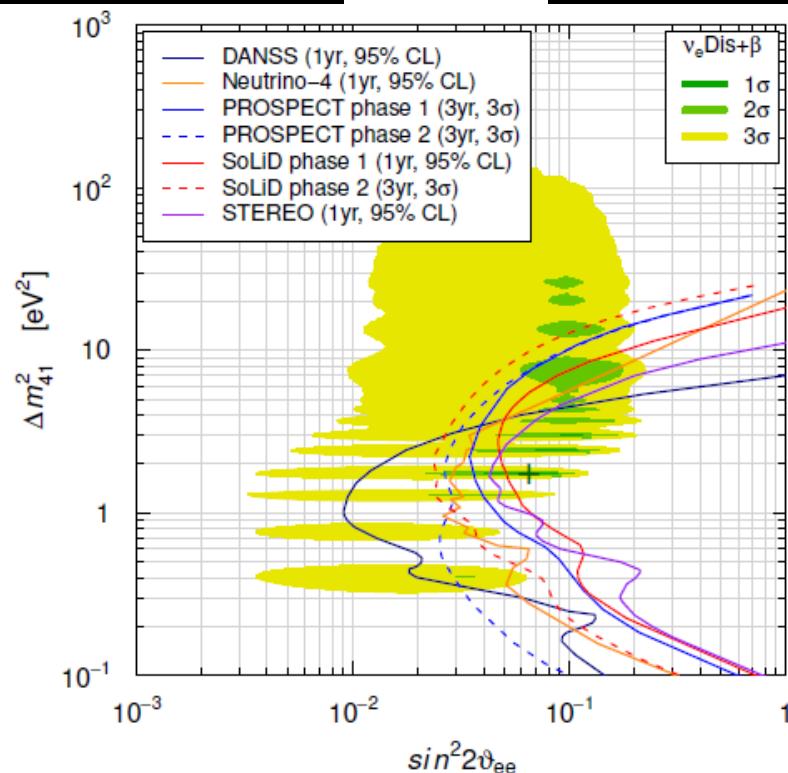
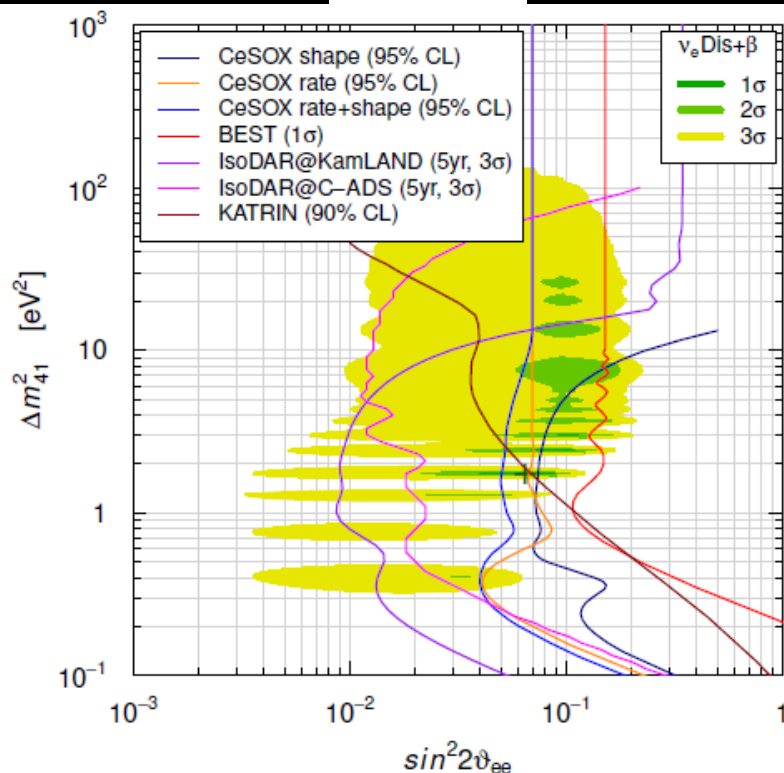


Expected sensitivity ν_e and anti- ν_e disappearance

C.Giunti, Moriond EW-2017

source

reactor



CeSOX (Gran Sasso, Italy) $^{144}\text{Ce} \rightarrow \bar{\nu}_e$
BOREXINO: $L \simeq 5\text{-}12\text{m}$ [Vivier@TAUP2015]

BEST (Baksan, Russia) $^{51}\text{Cr} \rightarrow \nu_e$
 $L \simeq 5\text{-}12\text{m}$ [PRD 93 (2016) 073002]

IsoDAR@KamLAND (Kamioka, Japan)
 $^8\text{Li} \rightarrow \bar{\nu}_e$ $L \simeq 16\text{m}$ [arXiv:1511.05130]

IsoDAR@C-ADS (Guangdong, China)
 $^8\text{Li} \rightarrow \bar{\nu}_e$ $L \simeq 15\text{m}$ [JHEP 1601 (2016) 004]

DANSS (Kalinin, Russia) $L \simeq 10\text{-}12\text{m}$ [arXiv:1606.02896]

Neutrino-4 (RIAR, Russia) $L \simeq 6\text{-}11\text{m}$ [JETP 121 (2015) 578]

PROSPECT (ORNL, USA) $L \simeq 7\text{-}12\text{m}$ [arXiv:1512.02202]

SoLid (SCK-CEN, Belgium) $L \simeq 5\text{-}8\text{m}$ [arXiv:1510.07835]

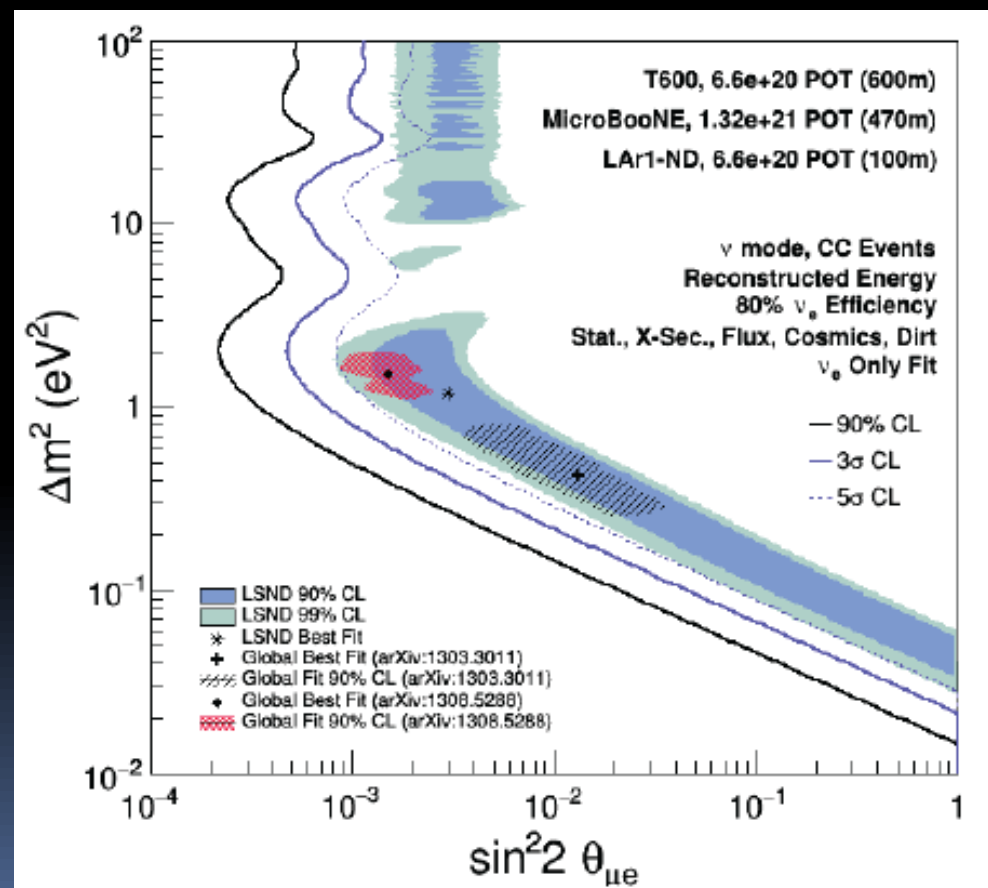
STEREO (ILL, France) $L \simeq 8\text{-}12\text{m}$ [arXiv:1602.00568]

KATRIN (Karlsruhe, Germany) $^3\text{H} \rightarrow \bar{\nu}_e$ [Drexlin@NOW2016]

Short baseline experiments: three LAr detectors

Detector	Distance from BNB Target	LAr Total Mass	LAr Active Mass
LAr1-ND	110 m	220 t	112 t
MicroBooNE	470 m	170 t	89 t
ICARUS-T600	600 m	760 t	476 t

Data taking to start in 2020



Sterile ν 's: « + » and « - »

+



LSND/MinBooNe

Reactor anomaly + Daya Bay RENO flux measurements
Ga anomaly

-



MINOS Disappearance

MINOS/Daya Bay/Bugey combined result
IceCube
NEOS, DANSS, Neutrino-4

?



More results from
accelerator, reactor, source experiments coming soon

Conclusion

Current LBL experiments T2K + NOvA

main goals: CP violation (3σ), Mass Hierarchy, θ_{23}

Next generation experiments: discovery/measurement of CP violation, determination of Mass Hierarchy

JUNO (MH) *under construction*

DUNE (CP, MH) *approved*

HyperK and T2HK (CP) *included in roadmap*

Light sterile neutrinos: crucial tests in experiments with neutrinos from accelerators, reactors, atmosphere and neutrino sources **in progress**

Thank you!