

Searches for alpha condensate states in relativistic nuclear fragmentation

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One of key aspects of nuclear structure is presence of degrees of freedom in which quartets of spin-paired protons and neutrons behave as constituent clusters, manifested in intensive formation of α -particles in a wide variety reactions and decays. Transition to study of ensembles of α -particles just above binding thresholds makes it possible to reveal a role of the unstable ⁸Be and Hoyle 3 α -state (HS) and search for their analogues.

A decay energy ⁸Be $\rightarrow 2\alpha$ is only $E_{th}(^{8}Be) = 91.8$ keV, and width $\Gamma(^{8}Be) = 5.57 \pm 0.25$ eV. The ⁸Be nucleus is an indispensable decay product of HS. The HS state is a second ¹²C excitation and a first at $E_{th}(HS) = 378$ keV above the 3 α threshold. A value $\Gamma(HS) = 9.3 \pm 0.9$ eV corresponds in order of magnitude to a width of the decay $\pi^{0} \rightarrow 2\gamma$.



An isolation of HS from higher ¹²C excitations points to it as a ⁸Be 3*a*-analogue. Synthesis of ¹²C is possible through the fusion $3a \rightarrow a^8Be \rightarrow {}^{12}C(0^+_2) \rightarrow {}^{12}C$ (+2 γ or e⁺e- with a probability of the order of 10⁻⁴). Further synthesis is possible in the sequence ${}^{12}C^{12}C \rightarrow {}^{12}C^{12}C(0^+_2) \rightarrow {}^{16}O^8Be$ determining the ¹²C and ¹⁶O abundances. These facts assume importance of heavier analogs.



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Modern interest is motivated by the concept of an α -particle Bose–Einstein condensate (α BEC). A manifestation of α BEC can be α -particle excitations of nuclei immediately above the binding thresholds of α -particles. Coexisting with fermionic excitations, they are considered on the basis of the mean field of the bosonic type, formed by a gas of almost ideal bosons in the S state at an average density four times lower than usual. The ⁸Be(0⁺) ground state is described as 2 α BEC, and HS as 3 α BEC. The ⁸Be \rightarrow 2 α and HS \rightarrow ⁸Be α decays can serve as signatures for cascade decays of more complex α BEC forms. The 0⁺₆ excitation of the ¹⁶O nucleus at 660 keV above the 4 α threshold, considered as 4 α BEC, can sequentially decay ¹⁶O(0⁺₆) $\rightarrow \alpha$ HS or ¹⁶O(0⁺₆) $\rightarrow 2^{8}$ Be(0⁺).



Hair - 60 μm AgBr Crystal - 0.2 μm Atom - 10⁻⁴ μm Proton - 10⁻⁹ μm



The Heavy Nuclei of the Primary Cosmic Radiation

H. L. BRADT AND B. PETERS University of Rochester, Rochester, New York (Received September 9, 1949)



SHOWER OF PHAPARTICLES PROTON **PI 11**8 1

The Study of **Elementary Particles** by the Photographic Method

An account of The Principal Techniques and Discoveries illustrated by An Atlas of Photomicrographs

C.F. POWELL P. H. FOWLER and D. H. PERKINS

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PERGAMON PRESS LONDON - NEW YORK - PARIS - LOS ANGELES

1959

1974 JINR Synchrophasotron¹²C @ 3.65 A GeV



Baldin Aleksand Mikhailovich

In the 70s exposures of NTE stacks to light nuclei at the Synchrophasotron (JINR) and Bevalac (LBL) began, and in the 90s - medium and heavy ones at AGS (BNL) and SPS (CERN). Even then, prospects of studying the nuclear structure in the limiting fragmentation region were recognized.

 $^{24}Mg \rightarrow 6\alpha$

The use of a technically simple and inexpensive method of NTE in beams of relativistic nuclei provides flexibility and uniformity at a search stage, and in a theoretical aspect, transparency of interpretation. It is very valuable to demonstrate similarity of conclusions based on relativistic invariance.

During the dissociation of relativistic nuclei in a narrow solid angle of fragmentation, ensembles of He and H nuclei are intensively generated. When they are detected, there are no thresholds, and energy losses are minimal.

According to the widths, decays of ⁸Be and HS occur at ranges several thousand atomic sizes and must be identified by a minimum invariant mass. Due to minimum energy, their decays of should appear as pairs and triplets of relativistic He fragments with the smallest opening angles. The same collimation along the interaction axis should be exhibited by more complex states of α BEC.





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Clusters in Nuclei, Volume 3



"Tomography" of the Cluster Structure of Light Nuclei via Relativistic Dissociation



In general, energy of a few-particle system Q is $Q = M^*$ - M. M^* is the invariant mass defined by the sum of all products of 4-momenta $P_{i,k}$ fragments $M^{*2} = \sum (P_i \cdot P_k)$. Subtraction of mass M is a matter of convenience. The 4-momenta $P_{i,k}$ are determined in the approximation of conservation of the initial momentum per nucleon. Then, the definition of Q comes down to determining the angles between the fragment emission directions.



The distribution over Q_{2a} of 2a-pairs ($N_{2a} = 500$), including "white" 2a-stars ($N_{ws} = 198$), indicates the condition $Q_{2a}(^{8}Be) < 0.2$ MeV. Then the fraction of decays $^{8}Be 36 \pm 3$ and $41 \pm 5\%$, respectively. Two "influxes" deserve to be noted around 0.6 and 3 MeV. The first of them corresponds to decay through the $^{9}Be 2.43$ MeV excitation, and the second corresponds to decay from the first excited state $^{8}Be_{2+}$.



Selected under the cleanest conditions, the criterion $Q_{2\alpha}(^{8}Be) < 0.2$ MeV includes the accepted approximations, the kinematic ellipse of the ⁸Be decay, and the resolution of angular measurements. Its application allows us to determine the ⁸Be contribution to the statistics of "white" stars equal to $45 \pm 4\%$ for $^{12}C \rightarrow 3\alpha$ and $62 \pm 3\%$ for $^{16}O \rightarrow 4\alpha$

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The Hoyle State in Relativistic ¹²C Dissociation



Distribution of 3*a*-triples over the invariant mass Q_{3a} of 316 "white" stars ${}^{12}C \rightarrow 3a$ (solid) and 641 "white" stars ${}^{16}O \rightarrow 4a$ (dashed) at 3.65 A GeV.

A fraction of events containing ⁸Be (HS) decays is $45 \pm 4\%$ (11 ± 3%) for ¹²C and 62 ± 3% (22 ± 2%) for ¹⁶O. The growth of 2a- and 3a-combinations enhances the contribution of ⁸Be and HS.

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HS can occur as a result of α decay of the 0⁺₆ excitation of the ¹⁶O nucleus. The distribution of "white" ¹⁶O \rightarrow 4 α stars over the invariant mass of 4 α -quartets $Q_{4\alpha}$ (Left) is described by the Rayleigh distribution with $\sigma_{Q4\alpha} = (6.1 \pm 0.2)$ MeV. The condition $Q_{3\alpha}$ (HS) < 700 keV shifts the Q4 α distribution and indicates 9 events $Q_{4\alpha} < 1$ MeV with $\langle Q4\alpha \rangle$ (RMS) = 624 ± 84 (252) keV. The contribution of ¹⁶O (0⁺₆) $\rightarrow \alpha$ + HS decays is estimated to be 7 ± 2% for normalization to HS. It can be concluded that the direct dissociation of α + HS dominates in the formation of HS.

33 events were identified ${}^{16}\text{O} \rightarrow 2^8\text{Be}$, which is $5 \pm 1\%$ of the "white" stars ${}^{16}\text{O} \rightarrow 4\alpha$. Then, the statistics of coherent dissociation for the ${}^{16}\text{O} \rightarrow 2^8\text{Be}$ and α HS channels has a ratio of 0.22 ± 0.02 . The distribution over the invariant mass $Q_{4\alpha} < 1.0$ MeV of events ${}^{16}\text{O} \rightarrow 2^8\text{Be}$ (Right) indicates two candidates ${}^{16}\text{O} (0^+_6)$. Thus, the estimate of the probability ratio of the ${}^{16}\text{O} (0^+_6) \rightarrow 2^8\text{Be}$ and α + HS channels is 0.22 ± 0.17 , which is unsatisfactory. At the same time, the increase in the statistics of events ${}^{16}\text{O} \rightarrow 4\alpha$ can be considered exhausted. There remains the possibility of studying (3-4) α -ensembles in the fragmentation of heavier nuclei.



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Light Clusters in Nuclei and Nuclear Matter: Nuclear Structure and Decay, Heavy Ion Collisions, and Astrophysics

David Blaschke, Hisashi Horiuchi, Masaaki Kimura, Gerd Roepke and Peter Schuck

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Unstable states in dissociation of relativistic nuclei

Recent findings and prospects of research

D. A. Artemenkov, V. Bradnova, M. M. Chernyavsky, E. Firu, M. Haiduc, N. K. Kornegrutsa, A. I. Malakhov, E. Mitsova, A. Neagu, N. G. Peresadko, V. V. Rusakova, R. Stanoeva, A. A. Zaitsev, P. I. Zarubin

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Au 10.7 GeV per nucleon









3)



It is possible that the unstable states are part of the nuclear structure that manifests itself in fragmentation. The alternative consists in the formation of ⁸Be in the interaction of α -pairs with the pickup of accompanying α -particles and nucleons by them. Its consequence would be an increase in a ⁸Be yield with a α -particle multiplicity n_{α} , and possibly ⁹B and HS. In the first case, an inverse correlation can be expected.

5)





na	N _{na} (⁸ Be)/N _{na}	N _{na} (⁹ B)	N _{na} (HS)	N _{na} (28Be)
	(% N _{nu})	(% N _{na} (⁸ Be))	(% N _{na} (⁸ Be))	(% N _{na} (⁸ Be))
2	3/133 (2 ± 1)	-	-	-
3	14/162 (9 ± 3)	1 (7)	-	-
4	25/161 (16 ± 4)	7 (28 ± 12)	2 (8 ± 6)	-
5	23/135 (17 ± 4)	5 (22 ± 11)	-	1 (4)
6	31/101 (31 ± 7)	9 (29 ± 11)	2 (6 ± 4)	-
7	31/90 (34 ± 7)	6 (19 ± 9)	2 (6 ± 4)	3 (10 ± 6)
8	32/71 (45 ± 10)	8 (25 ± 10)	2 (6 ± 4)	2 (7 ± 5)
9	29/54 (54 ± 13)	9 (31 ± 12)	3 (10 ± 6)	5(17 ± 8)
10	22/39 (56 ± 15)	4 (18 ± 10)	-	5(23 ± 12)
11	10/15 (67 ± 27)	3 (30 ± 20)	1 (10)	2(20 ± 16)
	19/30 (63 ± 19)	7 (37 ± 16)	2(11 ± 8)	6 (32 ± 15)
12	2/5	1	-	1
13	2/4	1	-	1
14	3/3	1	-	1
15	1/1	-	-	-
16	1/2	1	1	1

Olympus BX63 microscope:

- 1) digital camera DP74,
- 2) motorized object stage,
- 3) joysticks for controlling focus and movement of object stage in XOY plane,
- 4) microscope control unit,
- 5) PC for working with resulting image.

The monitor shows an enlarged image of interaction of a Kr nucleus below 1 GeV per nucleon

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nlpha	4	5	6	7	8	9-13
$N_{n\alpha}$	40(69)	50(54)	21(27)	10(19)	15(12)	7(3)
$N_{nlpha}/N_{ev},\%$	7.9 ± 1.0	6.2 ± 0.9	3.1 ± 0.6	2.2 ± 0.5	1.4 ± 0.4	0.4 ± 0.2
$N_{n\alpha} (\geq 1^8 \text{Be})$	5(15)	16(10)	12(13)	4(10)	11(8)	4(3)
$N_{n\alpha} (\geq 1^8 \text{Be}) / N_{n\alpha}, \%$	19 ± 5	25 ± 6	52 ± 13	48 ± 16	70 ± 21	70 ± 35
$N_{n\alpha}(2^8 \text{Be})$	0	2	2	1	5	2
N_{HS}	1	2	1	1	2	2

SUMMARY

Near threshold α -ensembles are generated in the limiting nuclear fragmentation. Only the nuclear track emulsion method answers the challenges. The ⁸Be $\rightarrow 2\alpha$ and HS \rightarrow ⁸Be α decays determined from the emission angles of the fragments at the very beginning of the spectrum of invariant masses.

This approach has been tested for nuclei from Ne to Au in a range from several to several tens of GeV per nucleon. The ⁸Be probability increases in the α -particle multiplicity. The ⁹B and HS contributions also increase, but their statistics are extremely low. The exotically large sizes and lifetimes of ⁸Be and HS suggest the synthesis of the α BEC states by combining α -particles u $2\alpha \rightarrow {}^{8}Be$, ${}^{8}Be\alpha \rightarrow HS$, $HS\alpha \rightarrow {}^{16}O(0{}^{+}_{6})$, ${}^{28}Be \rightarrow {}^{16}O(0{}^{+}_{6})$. Then α BEC can be regarded as a short-lived state of nuclear matter, not associated with the excitation of the parent nucleus.

In development, the BECQUEREL project aims to measure ⁸⁴Kr fragmentation below 1 A GeV to clarify the relationship ⁸Be and HS and the multiplicity and search for the 4 α BEC state . To track the evolution and universality of the conclusions, the 3.65 A GeV ²⁸Si fragmentation is analyzed.. The search for events in them by transverse scanning on the motorized microscope Olympus BX63 makes it possible to achieve the required statistics. Track measurements in selected α -ensembles are carried out on the most precise KSM microscopes.

Searches for *a*BEC lead to the study of nuclear matter with temperature and density ranging from red giants to supernovae. Future exposures to heavy nuclei at several GeV per nucleon of the NICA accelerator complex will make it possible to study relativistic ensembles of H and He isotopes of unprecedented multiplicity under optimal conditions.