

Search for low mass WIMP dark matter with DarkSide

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Dark Matter from Astrophysical Observations





Temperature spectrum of Cosmic Microwave Background



Rotation curves of stars and gas in galaxies

Direct WIMP Dark Matter Search with Argon, DarkSide Program



DarkSide-10 2010–2012 10 kg-scale prototype



Darkside-50 2013–2019 46 kg active mass



DarkSide-20k 2026 -20 tonne fiducial mass Approaching 'v-floor'

The goal: explore dark matter to the neutrino floor and beyond with extremely low instrumental background



DS-LowMass



ARGO ~2030 3k tonne-year

Direct Dark Matter Detection

One of the well-known candidates for dark matter - WIMP, Weakly Interacting Massive Particle, M ~ [GeV -TeV], e.g. from supersymmetry: neutralino, lightest superpartner with R-parity conservation

Dual-phase time-projection chamber (TPC)



Low Radioactivity Underground Argon



DarkSide Collaboration. "Separating 39Ar from 40Ar by cryogenic distillation with Aria for dark matter searches". Eur.Phys.J.C 81 (2021) 4, 359

DarkSide Collaboration. "Design and construction of a new detector to measure ultra-low radioactive-isotope contamination of argon". JINST 15 P02024 (2020)

DArT detector

Small-batch ³⁹Ar assay facility in Canfranc Lab, Spain **Sensitivity:** depletion factor U.L. of 6×10⁴ at 90% C.L. in 1 week of counting time



Two types of analyses for WIMP search

S1+S2

Search for 2 signals from nuclear recoil WIMP mass range > 20 GeV/c² Pulse Shape Discrimination Background fræ analysis (< 0.1 events) Energy diapason 45-200 keVnr

S2-only

Search for low energy nuclear recoil signals WIMP mass range 1.8-20 GeV/c² Full signal acceptance at 0.1 keV_{ee} Energy threshold 0.6 keV_{nr}

Best DM exclusion limit for 1.8–6 GeV/c² WIMP



Light WIMP search in DS-50



Renewed analyses for light WIMP search in DS-50

* Underground Argon campaign

- First 9 months of data is not used to wait for the ³⁷Ar (cosmogenic) to decay
- Extended exposure 650 live-days of data (12 ton-day) ~x 1.8 exposure than previous analysis
- Improved data selection. Detector
 was quite stable for the whole period
 of 26 months
 - $\delta T = \pm 0.02 K$, $\delta P < \pm 0.005 psi$,
 - $\delta(S1) \sim 0.4\%, \delta(S2) < 1\%, \tau e > 10 ms$ (more than x20 of the full drift time)
- New LAr response calibration
- New background modeling, including atomic effects of β-decays



Renewed Analyses for Light WIMP Search in DS-50 Data Selection

Fiducialization - events in the volume under 7 inner PMTs are selected to shield radiation from materials

Pileup pulses - pulse rise time and width are used to remove multiple low-S2 pulses misidentified as one pulse

Surface-α veto -S2/S1 band from calibration datasets used to reject surface-α induced event



Data Selection – Spurious Electron Veto



See also Astropart. Phys. 140 102704 (2022)

DS-50 observes the SE rate exponentially decaying with ~5 ms and ~50 ms time-constants A few delayed electrons signal is found after normal event : "Spurious Electron (SE)".

20 ms veto is applied for all events to maximize S/B ratio



Data Sample



Signal efficiency is >95% for the region of interest ($N_{\rho} = [4, 170]$)

Liquid Argon Response Calibration

Response to nuclear recoils

The model is based on Ar & erecombination process and energy losses due to nuclear stopping power

Parameters are tuned using DS-50 neutron calibration data (AmC and AmBe neutron sources) and external datasets from SCENE and ARIS experiments

The most "conservative" prediction among several theoretical models used



20

Energy [keVnr]

50

100

250

Liquid Argon Response Calibration

Response to electrons

Parameters are tuned for ³⁷Ar (L1shell) peak from the UAr campaign and ³⁹Ar sample from AAr campaign

Compared to the previous analysis, ^{83m}Kr and ³⁷Ar (K-shell) peaks are removed from the calibration to avoid distortion from their complex decay schemes

This allows for energy Region of Interst to be extended up to 170 e-(21 keV)



Detector Background

The activities are constrained by fitting higher energy spectrum, cross-checked by $\beta+\gamma$ fast coincidence events

 $^{39}Ar: 0.7 \pm 0.1 \text{ mBq/kg}$

 85 Kr : 1.8 ± 0.1 mBq/kg



The spectrum shape is accounted for recent calculations of atomic exchange and screening effects with associated uncertainties



Background Only Fit

Observed data is consistent to the background model prediction within the systematic uncertainties.

Background model and uncertainty (red line and shaded area) from the data fit in the [4, 170] N_e range, and the individual contributions from the internal (³⁹Ar and ⁸⁵Kr) and external components (cryostat and PMTs).

An excess of events with respect to the background model is observed below $4 N_{e}$.

All nuisance parameters fall within 1σ , without any strong correlation



Light WIMP New Exclusion Limits in DS-50



DarkSide-LowMass Detector for Light WIMP Search

Materials r/a purity studied in DarkSide-20k and DEAP-3600

Dominant γ sources (cryostat & photoelectronics) deparated from TPC by active vetoes

SiPM photosensors developed for DS-20k Much lower γ activity than PMTs

Underground argon low in ³⁹Ar and further depleted, using DS-20k's Urania & Aria plants



Low-background SiPMs developed for DarkSide-20k

Developed with Fondazione Bruno Kessler (FBK) Photodetection efficiency: > 40% at 77K Dark count rate: <0.01 Hz/mm2 at 77 K (7 VoV)







SiPMs [8×12 mm²]



PhotoDetection Unit (PDU) (16 tiles, arranged into 4 channels) [20×20 cm2]

Optical Plane (264 PDUs)

DarkSide Collaboration. "Cryogenic Characterization of FBK RGB-HD SiPMs". JINST 12 P09030 (2017)

A. Gola et al. "NUV-Sensitive Silicon Photomultiplier Technologies Developed at Fondazione Bruno Kessler". Sensors19(2), 308 (2019)

Projected and current 90% CL upper limits on spinindependent DM-nucleon scattering for Darkside-LowMass



arXiv:2209.01177

Conclusion

- Search for low mass WIMPs was carried out for the extended dataset from 2-years of DarkSide-50 experiment with low-radioactivity Underground Argon target.
- Liquid Argon Time Projection Chamber achieved high accuracies in measuring the extremely low energy depositions in liquid argon targets.
- The new analysis benefits from efficient data selection based on better understanding of the detector, more accurate calibration of the detector response, improved background model.
- The most stringent exclusion limit achieved on spin-independent DM-nucleon scattering for low-mass WIMPs at Mχ = [1.2, 3.6] GeV/c². arXiv:2207.11966-11968
- New experiment DarkSide-LowMass has been designed for low-mass WIMP search down to the "solar-neutrino fog". arXiv:2209.01177
- DarkSide-20k, DarkSide-LowMass, and ARGO experiments will explore WIMP down to the "solar-neutrino fog". DS-20k data taking will start from 2026.

DARKSIDE

Thank you for your attention

