Copernican Principle Beyond FLRW



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Standard Model of Cosmology

ACDM model has gloriously succeeded in explaining most of the cosmological observations, Throughout all the redshifts





Cosmological Tensions

For the last couple of decades there is a tension ($\sim 5\sigma$) regarding early and late time cosmologies H₀=67.3±0.6 km/s/Mpc

H₀=74.3±1.4 km/s/Mpc



Non Kinematic Dipole

There are increasingly numerous hints that have emerged for a non-Kinematic component in the CMB Dipole.

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If the Dipole in the CMB is due to "Our motion" w.r.t the "CMB



frame", then we should see similar Dipole in the distribution of the distant sources.



Recent Observation

In one of the recent works by N.J. Secrest, S. von Hausegger, M. Rameez, R. Mohamed, S. Sarkar and J. Colin showed a ~ 4.9σ conflict between the expected and measured amplitude of Dipole in the angular distribution of 1.4 million quasar catalogue.





Dipole Cosmological Paradigm



New Directions

4 Conclusion

Anisotropies in radio-source number counts can be used to determine a cosmological standard of rest. Current observations determine it to about $\pm 500 \text{ km s}^{-1}$, but accurate counts of fainter sources will reduce the error to a level comparable to that set by observations of the microwave background radiation. If the standards of rest determined by the MBR and the number counts were to be in serious disagreement, one would have to abandon either

(a) the idea that the radio sources are at cosmological distances, or(b) the interpretation of the cosmic microwave radiation as relic radiation from the big

(b) the interpretation of the cosmic microwav bang, or

(c) the standard FRW Universe models.

Thus comparison of these standards of rest provides a powerful consistency test of our understanding of the Universe.

Quasars in the CATWISE catalogue are typically at redshifts ~ 1. Thus the most logical question that follows is, "what is the simplest generalization of FLRW paradigm ?"

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Copernícan príncíple In Our Context

Given a prior, what is the most symmetric universe?

No príor $\xrightarrow{c.p}$ Any Maximally symmetric spacetime

Expanding universe $\xrightarrow{c.p}$ Isotropy and Homogenety in the spatial slice (FLRW)

What is our prior?



The universe has a cosmic flow or

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Cosmic flow + Expansion $\xrightarrow{c.p}$ Locally rotationally symmetric Bianchi V We call this 'Dipole Cosmology'



Cosmological Principle

If one assumes Copernican principle and Observe something (prior). The symmetries inferred are the principles for a cosmological paradigm.

The Standard Cosmologícal Príncíple

If the universe appears isotropic from a vantage point on the earth, it infers that the universe is also homogeneous. These two conditions together are stated as 'The Cosmological Principle'

The Dipole Cosmological Principle 1) It is homogeneous in the spatial slices or simply transitive under a group of dimension three. 2) It has a cosmic flow along a spatial direction



The Metric $ds^{2} = -dt^{2} + X^{2}dz^{2} + Y^{2}\exp(-2z)(dx^{2} + dy^{2})$



Homogeneity

we can move between any two points in the spatial slice by an action of a group of dimension 3. The three killing vectors for this spacetime are.



- 1) $\xi_1 = \partial_x$
- 2) $\xi_2 = \partial_y$
- 3) $\xi_3 = x\partial_x + y\partial_y + \partial_z$



 $T_{ab} = \begin{pmatrix} \rho + (\rho + p)\sinh^2\beta & -(\rho + p)\sinh\beta\cosh\beta & 0 & 0\\ -(\rho + p)\sinh\beta\cosh\beta & (p + (\rho + p)\sinh^2\beta) & 0 & 0\\ 0 & 0 & p\exp(-2z)Y^2 & 0\\ 0 & 0 & 0 & p\exp(-2z)Y^2 \end{pmatrix}$





Standard Cosmology (Open)

First Order



Conservation

 $\dot{\rho} + 3H(p+\rho) = 0$

Einstein's Equations Dípole Cosmology $\left(\frac{\dot{a}}{a}\right)^2 - \frac{k}{a^2} = \frac{\rho}{3} \qquad 2\frac{\dot{X}\ddot{Y}}{X} + \left(\frac{\dot{Y}}{Y}\right)^2 - \frac{3}{X^2} = \rho\cosh^2\beta + p\sinh^2\beta$ $\frac{2}{V}\left(\frac{\dot{X}}{V} - \frac{\dot{Y}}{V}\right) = (\rho + p)\sinh\beta\cosh\beta$ $\dot{\rho} + (\rho + p) \left(\frac{\dot{X}}{V} + \frac{2\dot{Y}}{V} + \tanh\beta\dot{\beta} - \frac{2\tanh\beta}{V}\right) = 0$ 17 $\dot{p} + (\rho + p) \left(\frac{X}{X} + \coth\beta\dot{\beta}\right) = 0$







- $p = w(t)\rho$
- w(t) = a function of time.
- Tilt variation $\dot{\beta}(\cosh\beta w(t)\tanh\beta)$
- Main message: We can have in acceleration.

• The consistency is restored by introducing equations of state (EoS),

• We will mainly highlight the evolution of the tilt $(\beta(t))$ & shear $(\sigma(t))$. The two main categories of our results are for 1) w(t) = Constant 2)

$$h\beta) = -\frac{\dot{w}}{1+w} + (3w-1)H - \frac{2}{3}\sigma - \frac{2w \tanh\beta}{X}$$

ncreasing tilt for an universe with late tin





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Plots for the 'Tilt' and 'Shear' for an effective EoS $w(t) = -1 + a \exp(-bt)$

arXiv: 2209.14918[astro-ph.CO]

Evolution of 'Tilt' & 'Shear' for $w(t) = w_{LCDM} = -\tanh^2 \left(1.5 \sqrt{\Omega_{\Lambda} H_0^2 t} \right)$

• We look for an Copernican ansatz that allows a cosmic flow.

- The spacetime turns out to have Spatially Homogeneous slice, with an axial symmetry along a spatial direction.
- Owing to the SH, the field equations are ODEs & the solutions are easily tractable.
- As tilt is a function of time, we are expected to find different angular distributions for observations at different redshifts.
- This system has a instability in tilt. It argues for a tilt growth at late times even for a minimal initial tilt perturbation (nearly isotropic CMB).

