



Electron beam test of the MPD electromagnetic calorimeter on the «Pakhra» synchrotron

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- ✓ Initial beam NICA collider parameters :

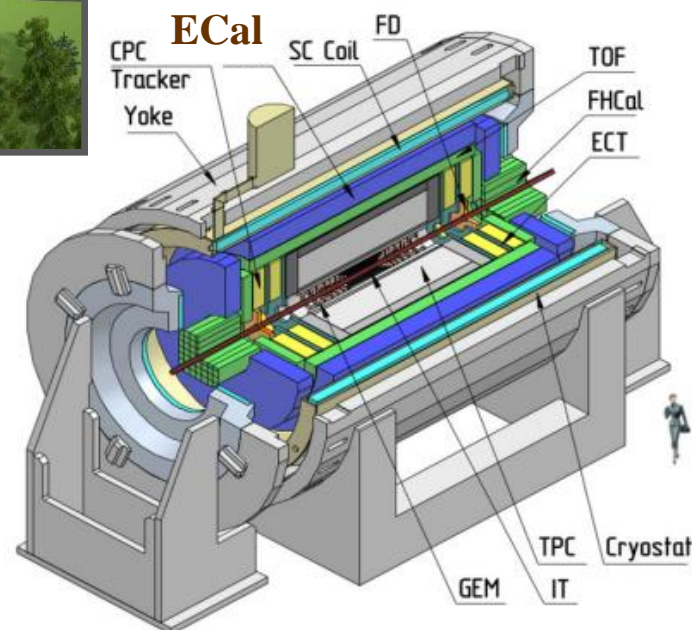
$$\sqrt{S_{NN}} = 7 \div 11 \text{ GeV}$$

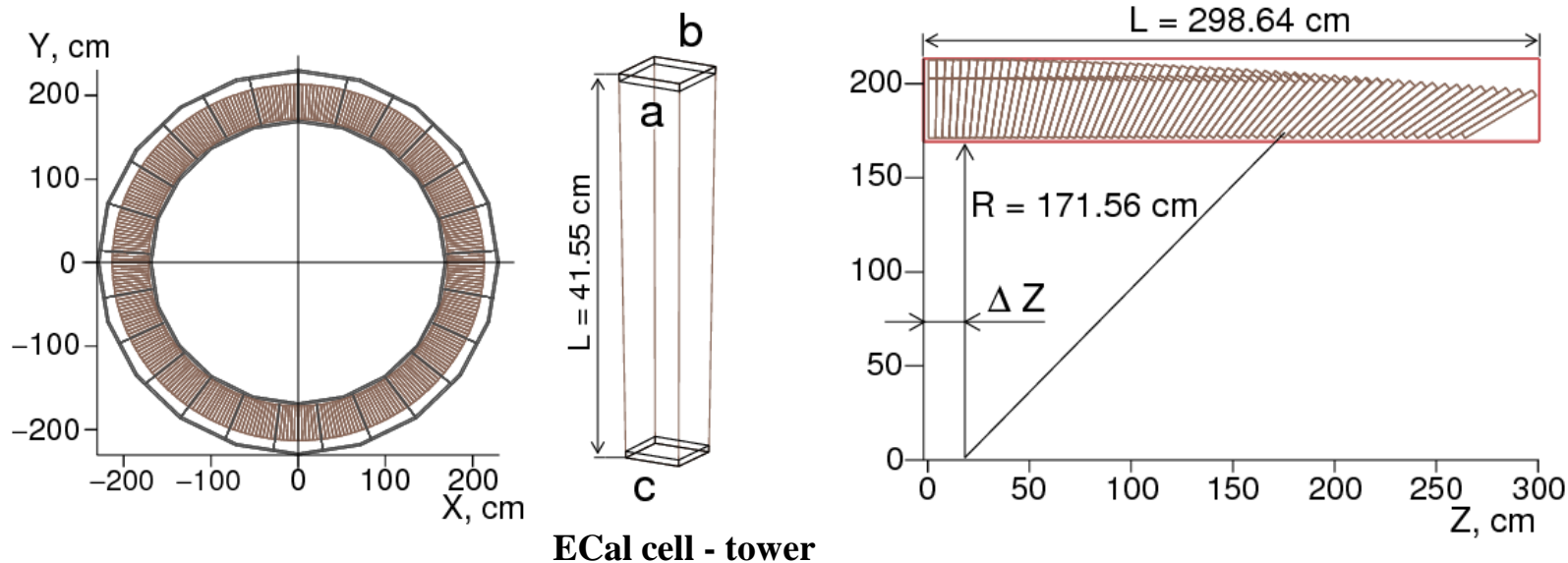
Ions : $\text{Au}^{197}/\text{Bi}^{209} + \text{Au}^{197}/\text{Bi}^{209}$

$$L \sim 10^{27} \text{ cm}^{-2} \text{ c}^{-1}$$

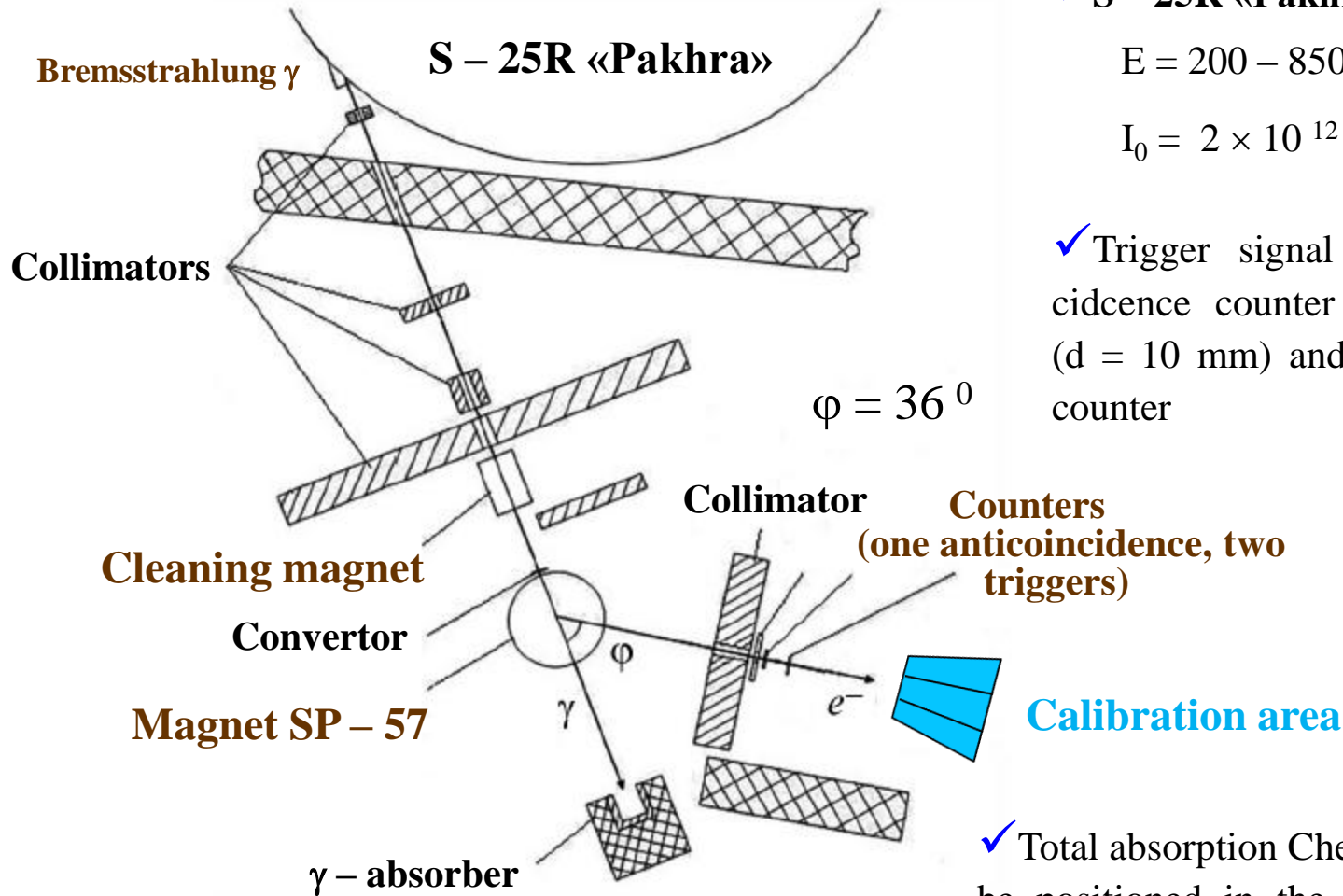
- ✓ Now collider building and MPD hall are ready

- ✓ MPD first stage : construction and installation of the superconducting magnet, ECal, TOF, TPC, ZDC
- ✓ All components of the MPD first stage detector are in production stage, commissioning expected for next year





- ✓ ECal geometry is close to projective, that means all cells are oriented to the NICA collider beams crossing area. ECal consists from 2400 modules of different types
- ✓ Each module consists from 16th cells (towers) fixed by the special gule (N towers = 38 400)
- ✓ Tower calorimeter has a sampling structure consisting of active medium (scintillator, width = 1.5 mm) and absorber (lead, width = 0.3 mm). Total number of such layers in cell is equal to 210.
- ✓ The shapes of the towers are approximately truncated pyramids with base of $4 \times 4 \text{ cm}^2$, more accurately described by a set of trapezoids (milling angle = 0.9° along Z and 1.2° in XY – plane)



✓ S – 25R «Pakhra» parameters :

$$E = 200 - 850 \text{ MeV}$$

$$I_0 = 2 \times 10^{12} e^- / \text{second}$$

✓ Trigger signal : one anticoincidence counter with inner hole ($d = 10 \text{ mm}$) and two scintillation counter

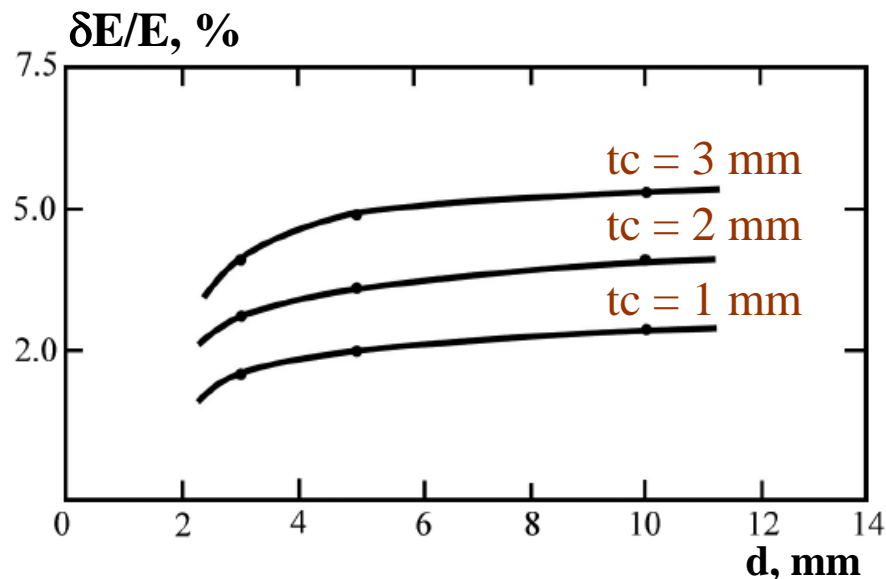
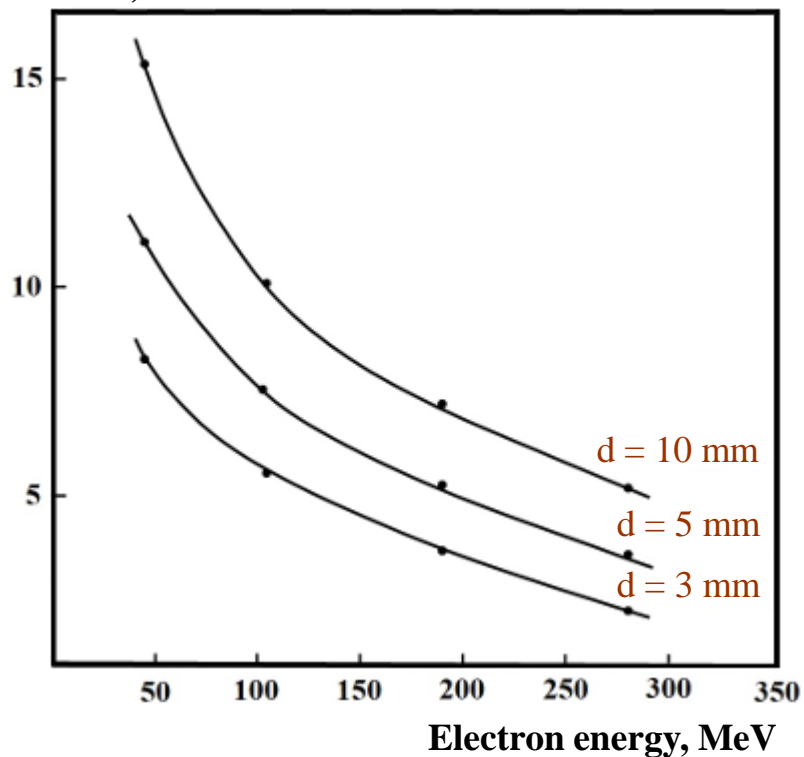
✓ Total absorption Cherenkov counter can be positioned in the calibration area to define electron energy

✓ Beam energy resolution depends from collimator diameter (d) and convertor thickness (tc)

✓ Electron beam was optimized at $d = 5\text{mm}$ and $tc = 1\text{ mm}$ with intensity :

$$I_0 \sim 20 \div 50 \text{ e}^- / \text{second}$$

$\delta E/E, \%$



Calibration beam energies

| Beam energy, MeV | Energy resolution , % |
|------------------|-----------------------|
| 30 | 13.8 |
| 54 | 11.0 |
| 98 | 7.5 |
| 150 | 6.3 |
| 196 | 5.4 |
| 293 | 3.4 |

- ✓ Parameterization of the cosmic ray spectra (atmospheric muons) obtained from experimental data and covers wide energy range
- ✓ Cosmic ray flux at sea level has $\cos^{N-1} \Theta$ dependence, where N – power of the energy distribution [P. Shukla, *Int.J.Mod.Phys. A33* (2018) no.30, 1850175]

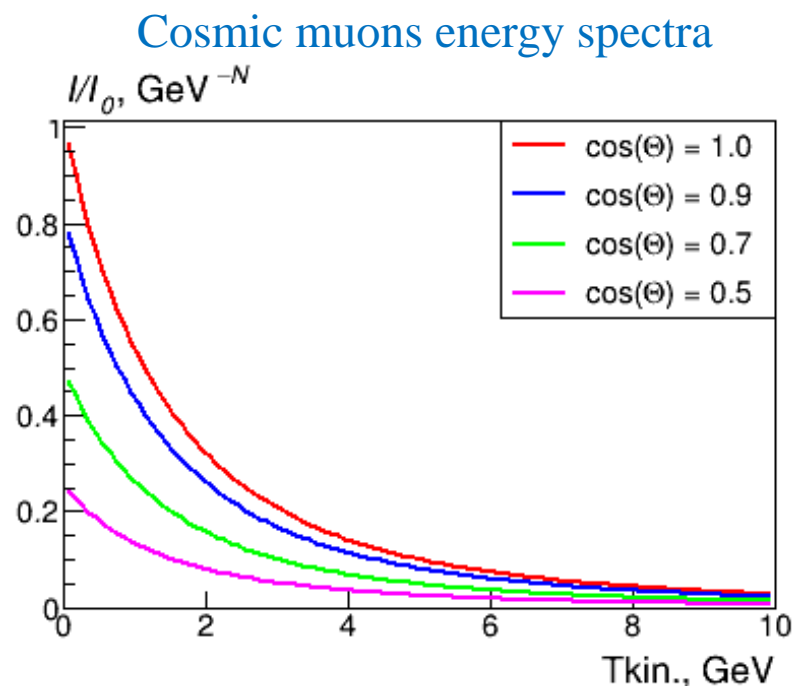
$$I / I_0 = (E_0 + E)^{-N} \times \cos^{N-1} \Theta / (1 + E / 854);$$

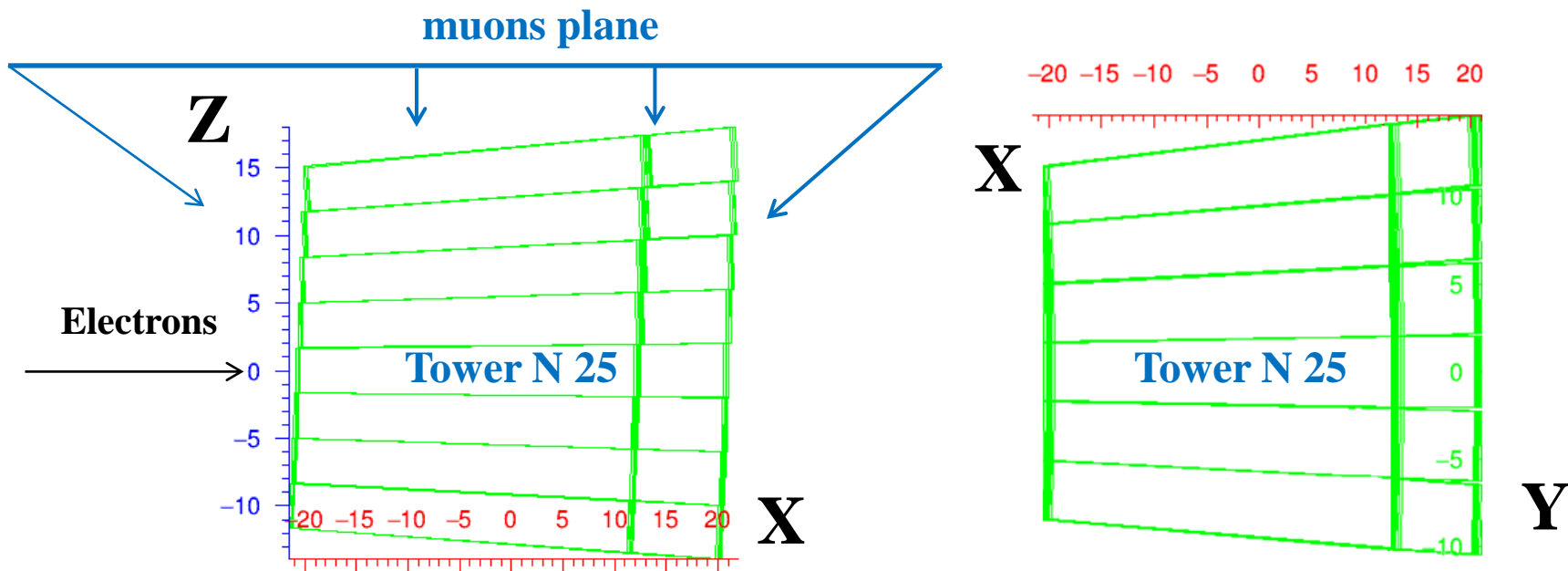
$$E_0 = 4.29, N = 3.01$$

- ✓ $\cos \Theta$ – zenith angle, E – energy;

this formula is valid for muons, not for proton part

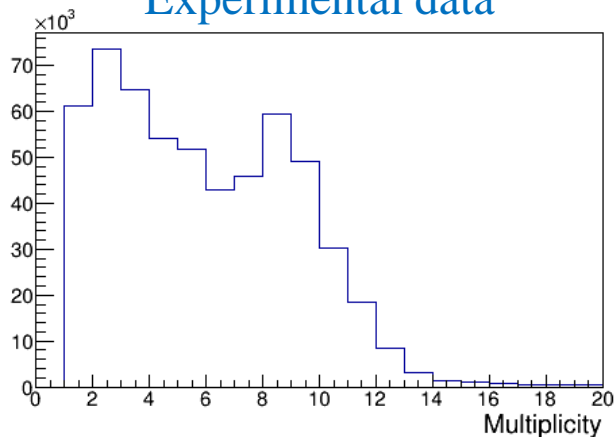
- ✓ Generator was developed in the framework of MpdRoot, where muons are produced as function of energy (0 – 10 GeV) and zenith angle



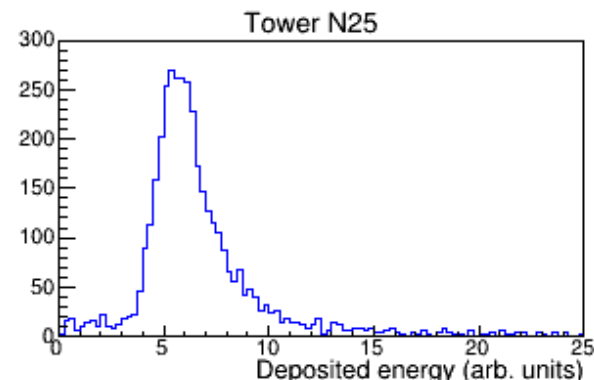
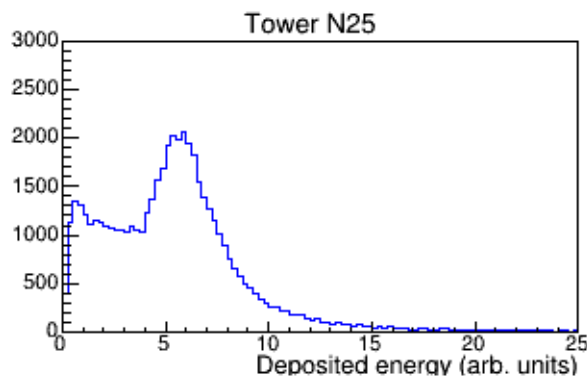
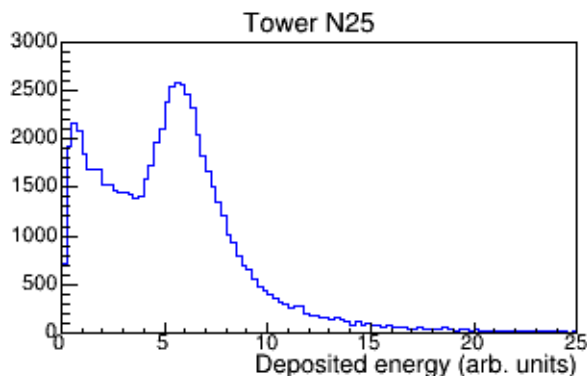


- ✓ Geometry of one module of the first type (emcModule0) is extracted from file emc_v3.root directly to add in the mpdroot frame (only this detector)
- ✓ Three module were joined to the assembly (48 towers, 6 in YZ) \times 8 in XZ), as for experimental data
- ✓ Cosmic muons are generated from large plane at $Z_0 = 20$ cm in all directions, selected only particles passing detector area

Experimental data



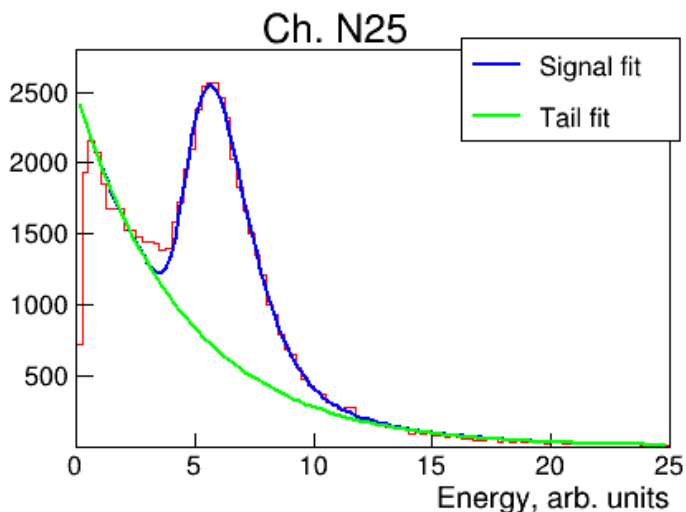
- ✓ Multiplicity of the module assembly is large number (until 15 tower / one muon). It's a good way to collect a large statistics during short time
- ✓ For better presentation, it's useful to add scale coefficient = 10^{-3}



✓ Method 1: all events, low energy part can simply describe by exponential function; use this method for further analysis

✓ Method 2: multiplicity > 7 , less statistics, but tail still exists

✓ Method 3: two (upper and lower towers in line give a trigger), tail is eliminated, but statistics – 7 times less

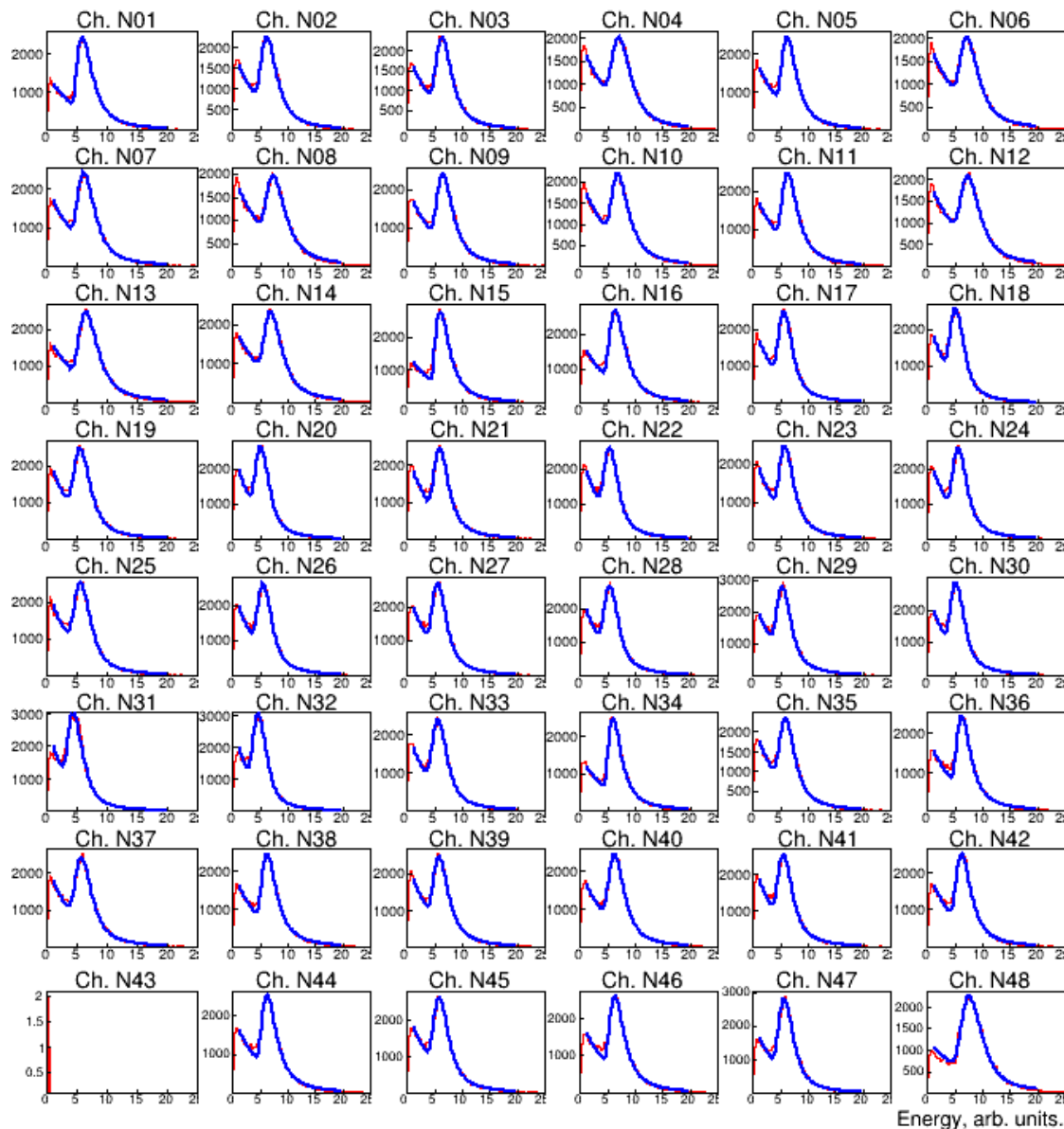


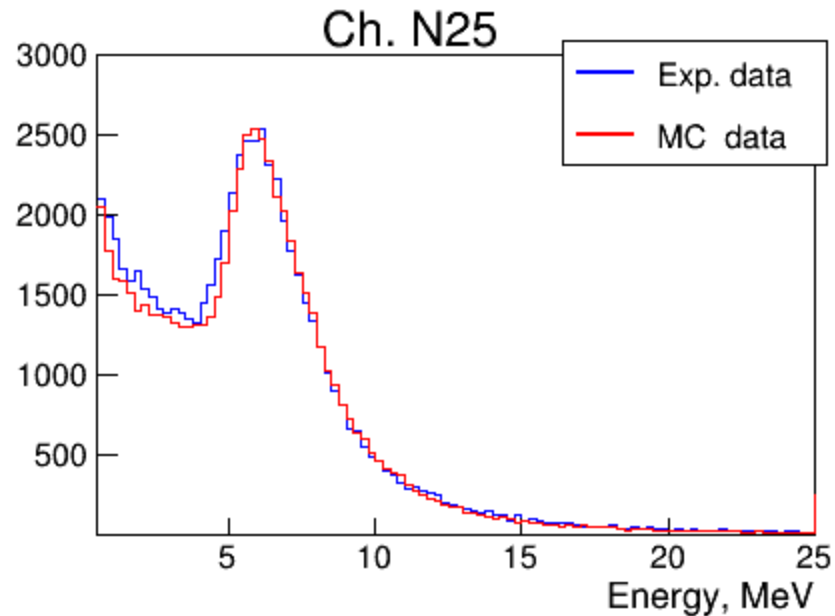
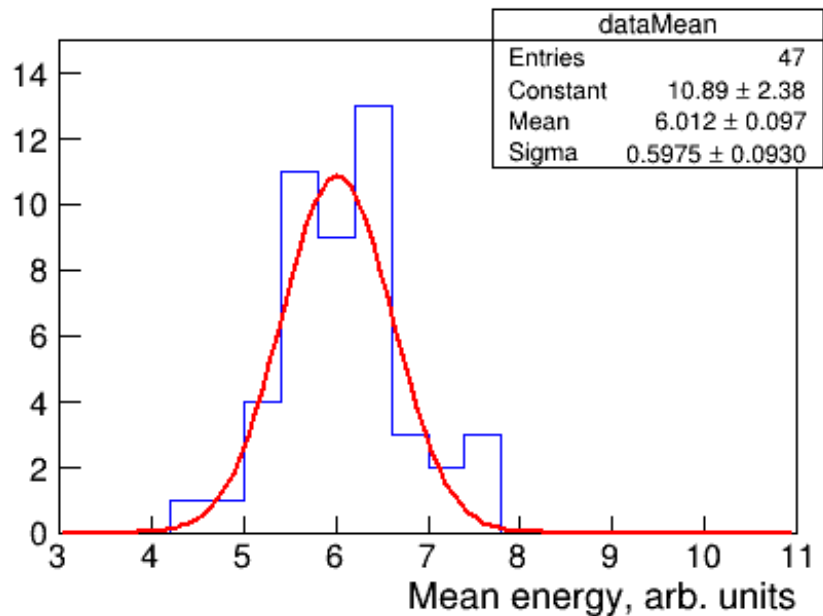
✓ Clear peak from muons can be a reliable tool for calibration

✓ Fit is a sum of two functions (signal and exponential tail)

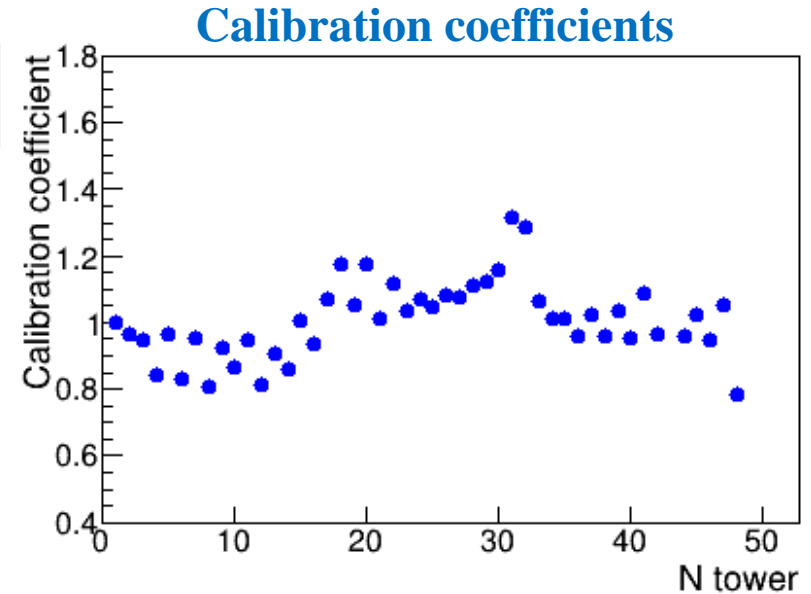
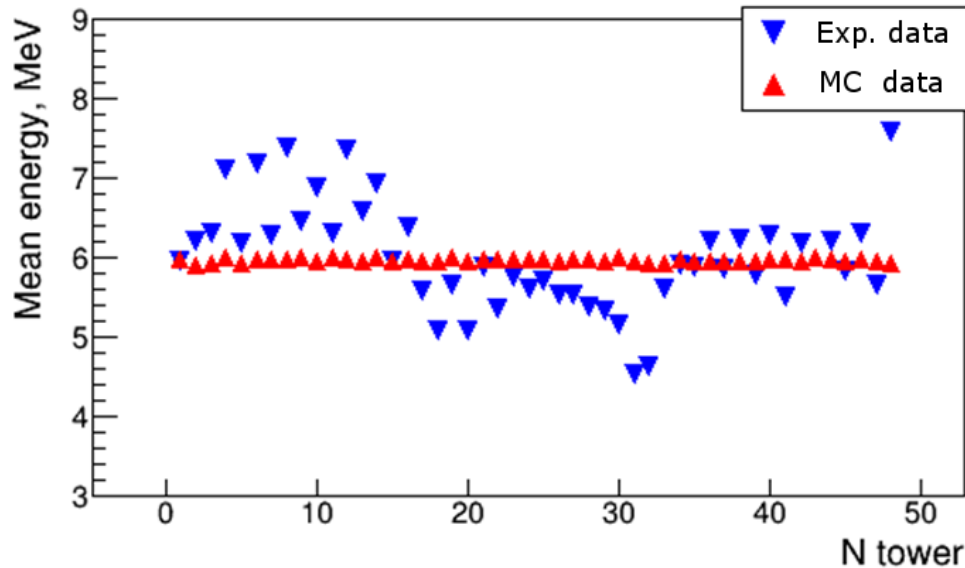
✓ Signal fitting function :

$$\sim (x - x_0)^\alpha / x^\beta$$

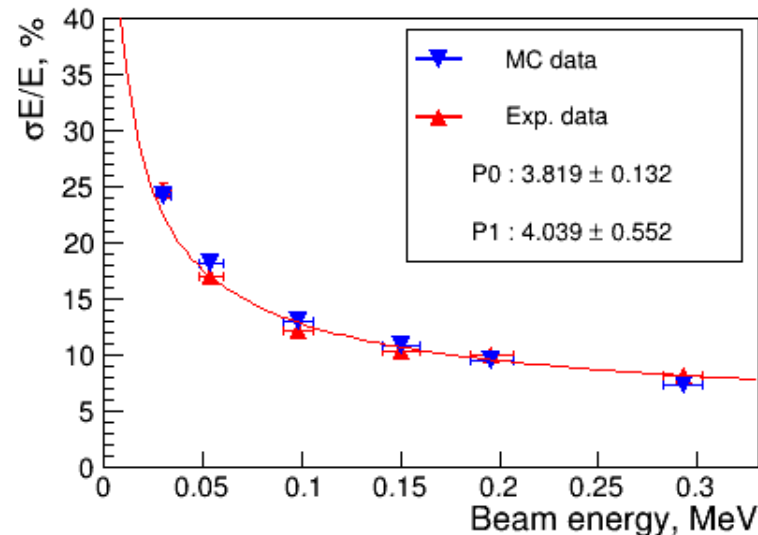
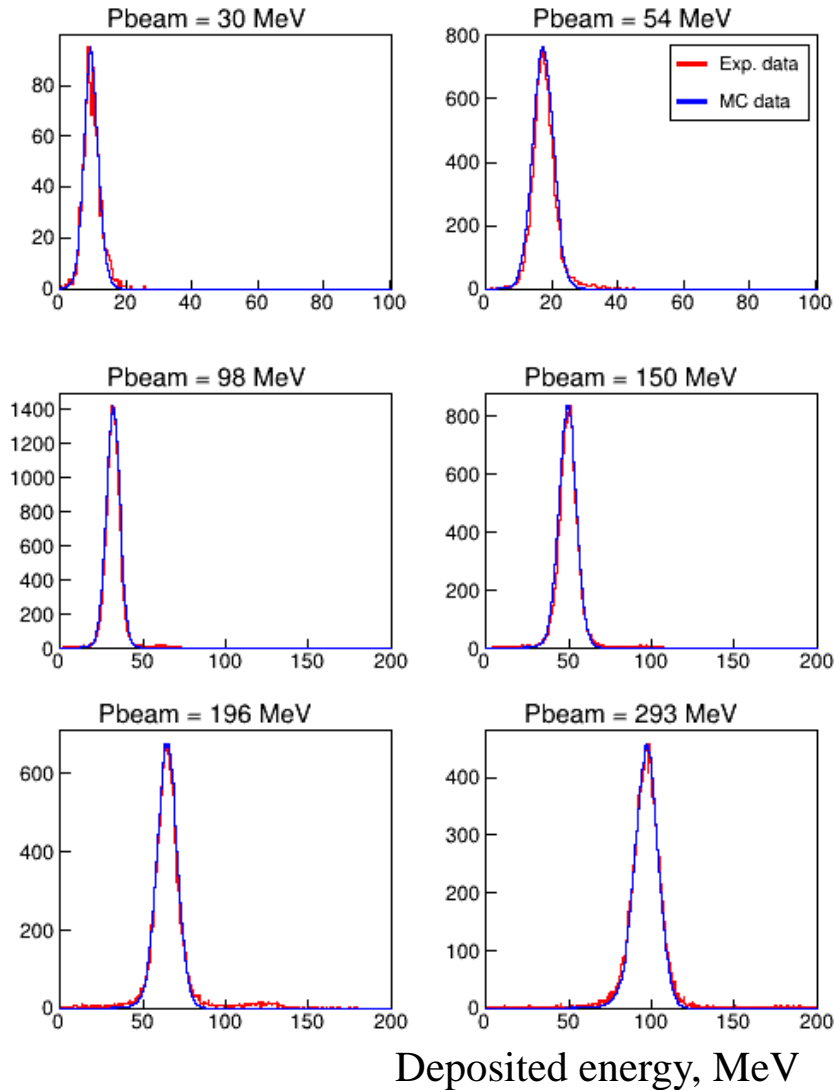




- ✓ Experimental and MC data are in a good agreement
- ✓ MC peak is stable for all 48 towers and equal to 5.95 MeV
- ✓ Experimental data are varied for different modules (each 16 towers)
- ✓ Experimental mean value (peak position) is 6.01 (first tower, arb. units) and comparison with MC gives conversion factor ~ 0.99 MeV

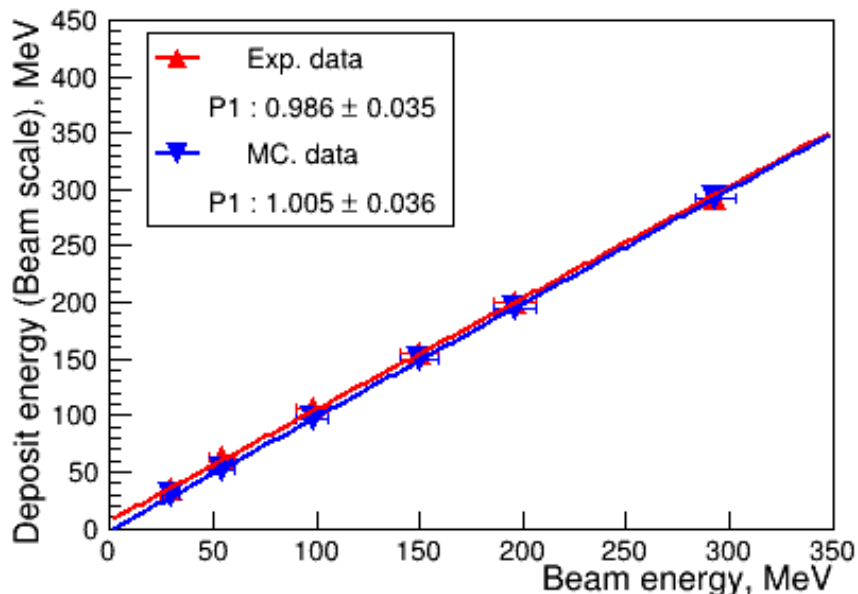


- ✓ Good method to perform a quality control of each produced tower/module
- ✓ Independence of the MC peak position from its number demonstrates a reliability of this calibration method
- ✓ It doesn't take a significant time to be done (one day/night is enough to test module)
- ✓ Calibration coefficients are defined concerning mean value, the fluctuation of this calibration coefficient $\sim 10\%$
- ✓ This calibration is used for further analysis to obtain electron energy

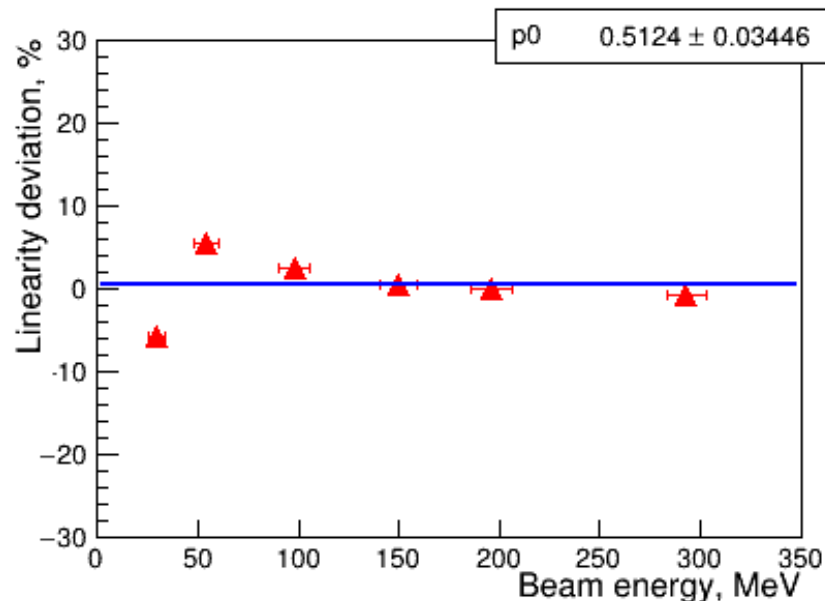


- ✓ Module assembly was measured at 6 electron beam energies
- ✓ Deposit energy is defined by all towers
- ✓ All towers are calibrating by cosmic data
- ✓ Energy resolution is calculated with electron beam spread
- ✓ MC and experimental data are in a good agreement within $\sim 5 \div 6 \%$

ECal linearity



Deviation from linearity



- ✓ Calorimeter linearity is estimated with respect to the largest beam energy (293 MeV)
- ✓ Linearity is observed well for MC data, but for experimental data there is a deviation and scale coefficient is less than unity
- ✓ Some reasons of that are clear :
 - 1) Uncertainty of the beam energy ($\sim 10\%$ of beam energy at 54 MeV)
 - 2) In general, light loss in tower fibers, this effect is larger for low energy electromagnetic shower

- ✓ A significant number of the modules (~ 250) are in the process of production on the factory «Tensor» at Dubna (Russia)
- ✓ The extracted electron beam at S – 25R «Pakhra» (LPI, Troitsk division) with energies until 500 – 600 MeV will be launched to study characteristics and calibration of the produced calorimeter modules. Addition calibration area is prepared for shifters during the calibration of equipment and detectors.
- ✓ Calibration of the new produced modules will be carried out on cosmics to perform test quality procedure, and limited number of modules – on electron LPI beam to measure a preliminary calibration characteristics
- ✓ MC simulation and test strategy are ready. MC and experimental data demonstrate a good agreement, but need some adjustments
- ✓ This work is supported by the *RFBR Grant №18-02-40079*