

ICPPA-2016

# **The big geomagnetic storm of 17 March 2015: Measurements from EPT / PROBA-V as well as their interpretation**

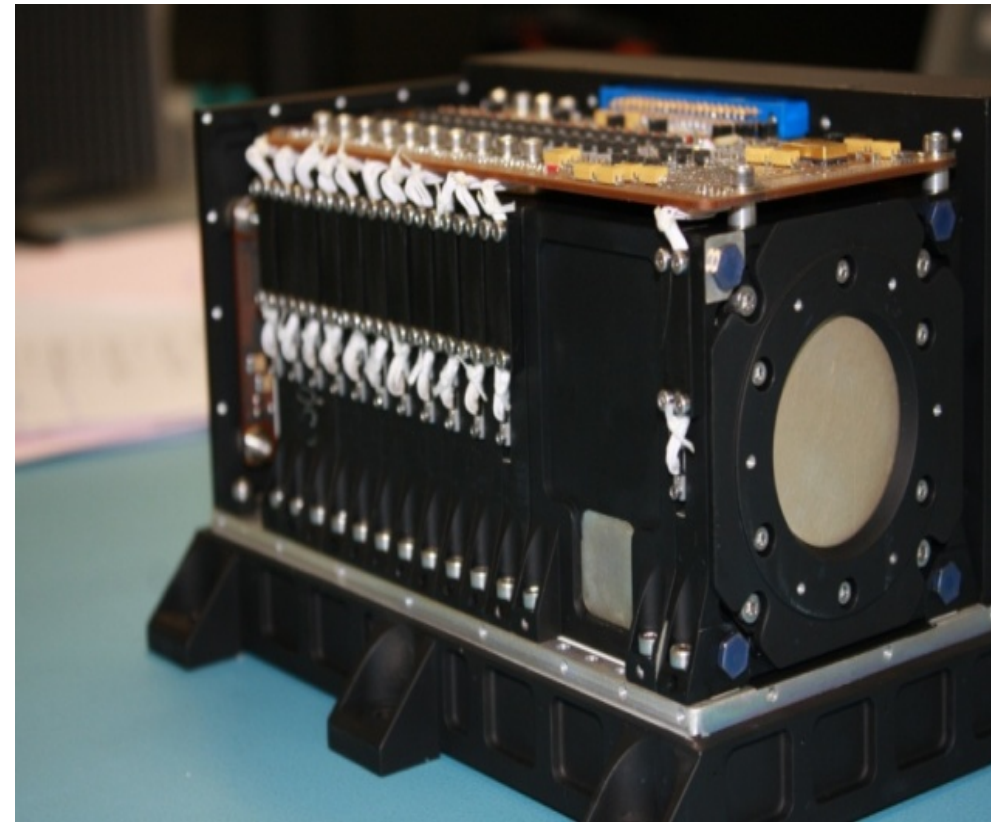
Viviane Pierrard <sup>1, 2</sup>  
Graciella Lopez-Rosson <sup>1</sup>  
Joseph Lemaire <sup>1, 2</sup>

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<sup>2</sup> *CSR/UCL, Louvain-La-Neuve*

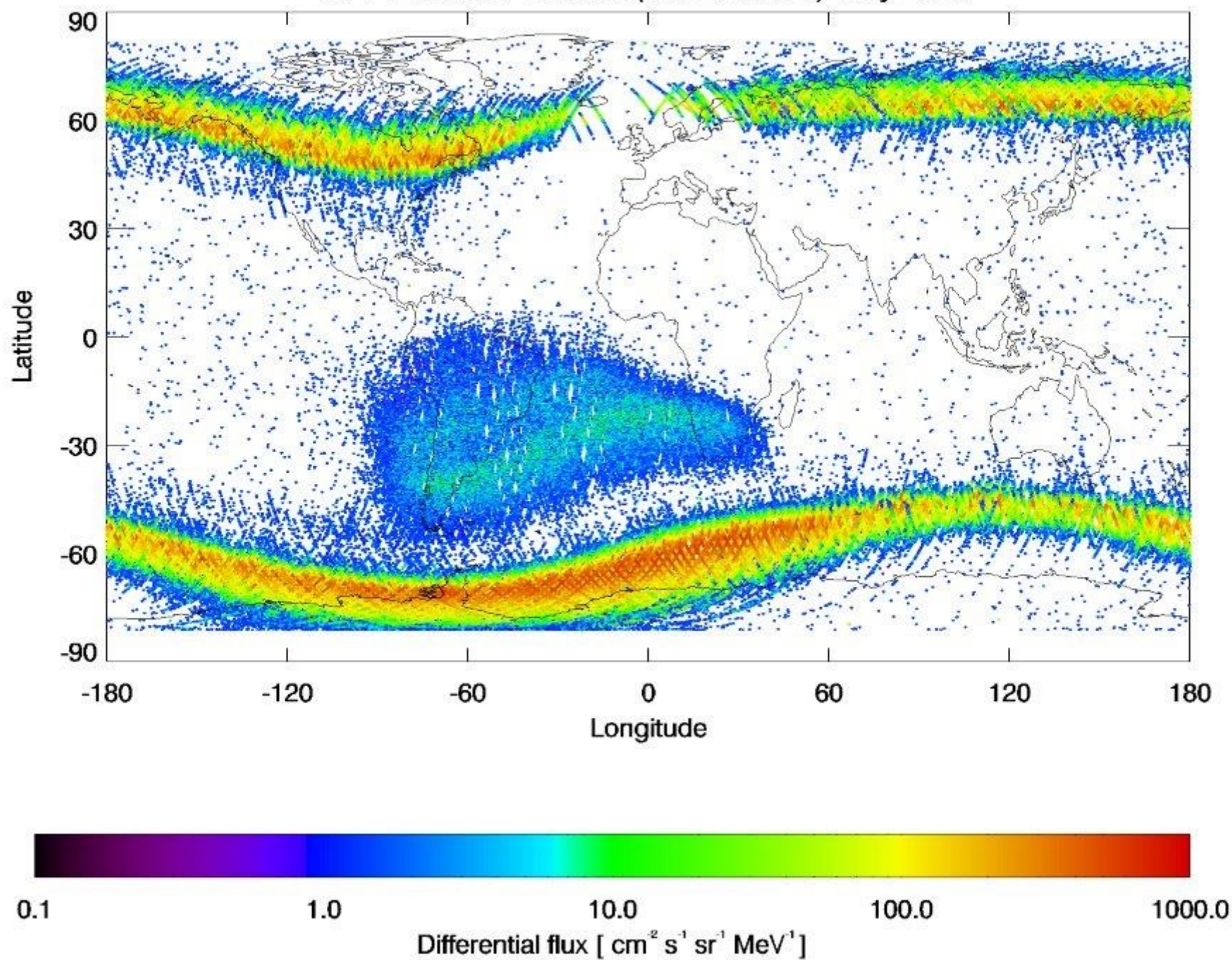
*Moscow, 11 Oct. 2016*

# Energetic Particle Detector: EPT

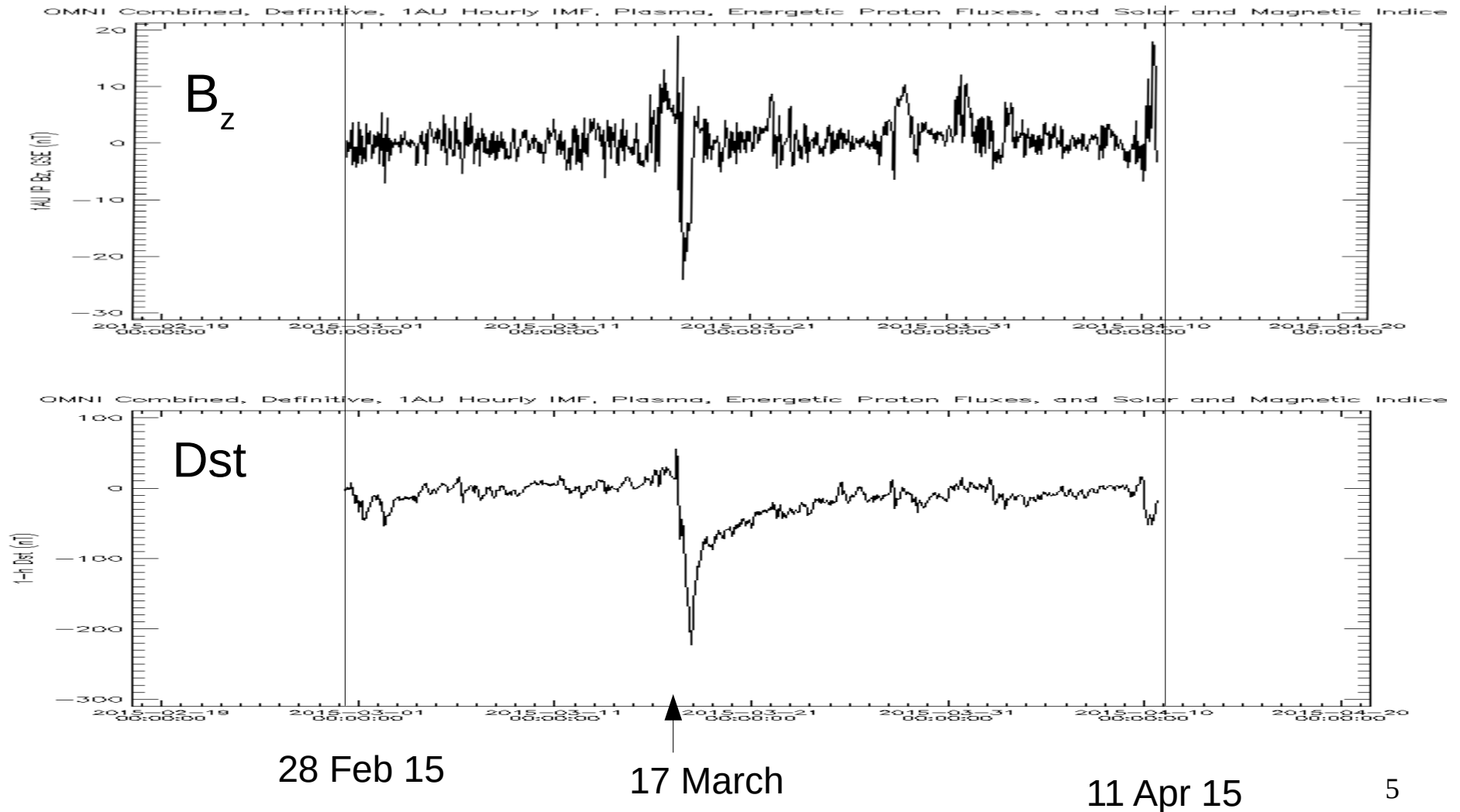
- Design & development:  
Centre Space Radiation / [UCL](#) / [LLN](#)  
Belgian Institute Space Aeronomy / [Brussels](#)  
QinetiQ Space
- Digital PIPS detector/spectrometer/ FOV : 52°  
LES : 3 sensors  $\Delta E - E$   
HES: stack of 10 absorber/detector
- Discrimination particle species :  
 $e^-$  (0.5 - 20 MeV ; 7 ch )  
 $p^+$  (9.5 - 248 MeV ; 11 ch)  
Heavy ions
- Efficiency Calculation with GEANT4
- [Cyamukungu et al.](#), *IEEE Trans. Nucl. Sci.* 61, 3667, 2014,
- [Pierrard et al.](#), *Space Sc. Rev*, 184, 87, 2014



EPT Electron flux ch5(1.0-2.4 MeV) July 2013

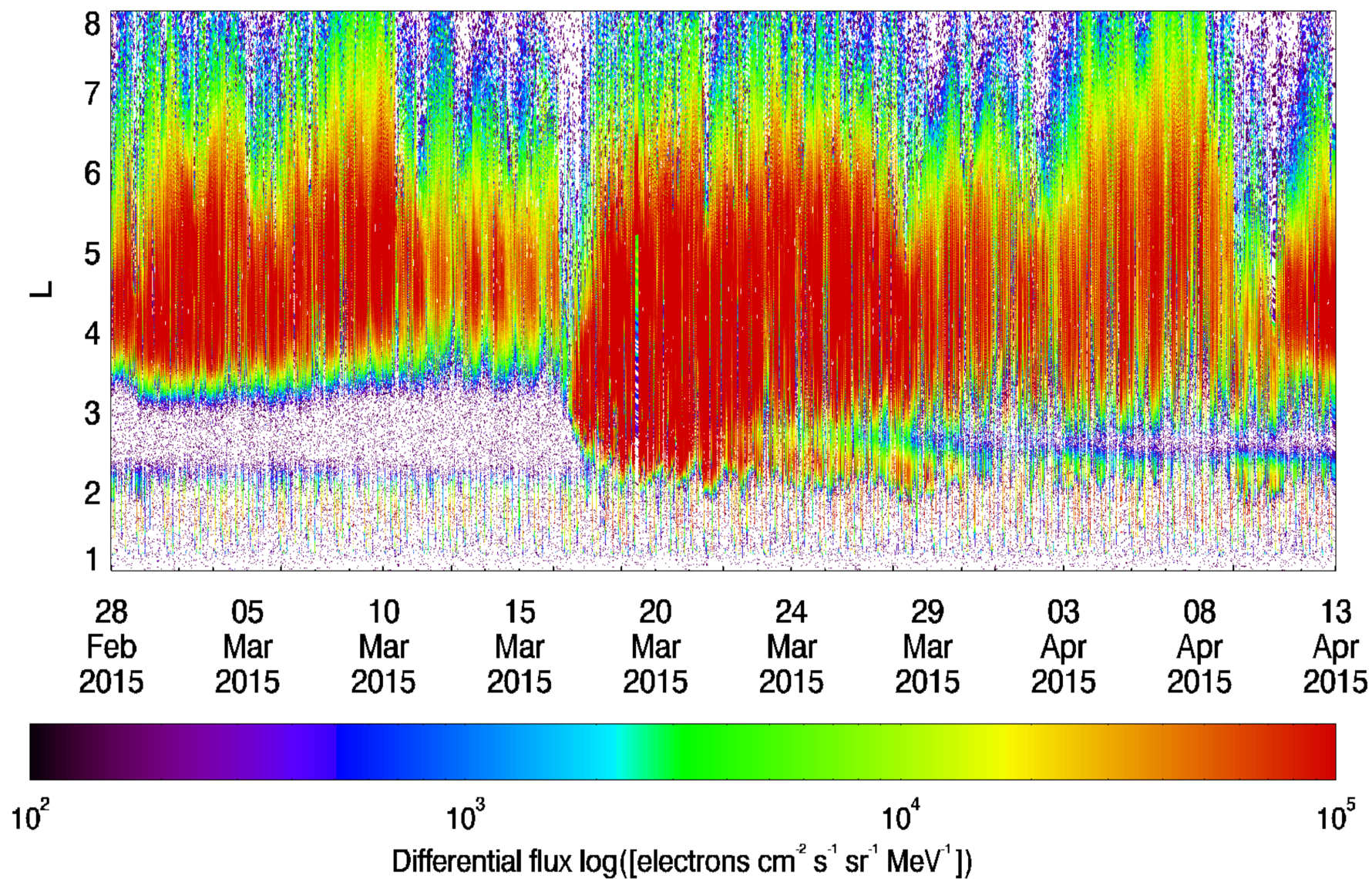


## Geomagnetic Storm of 17 March 2015 (Dst\_min = -223 nT)

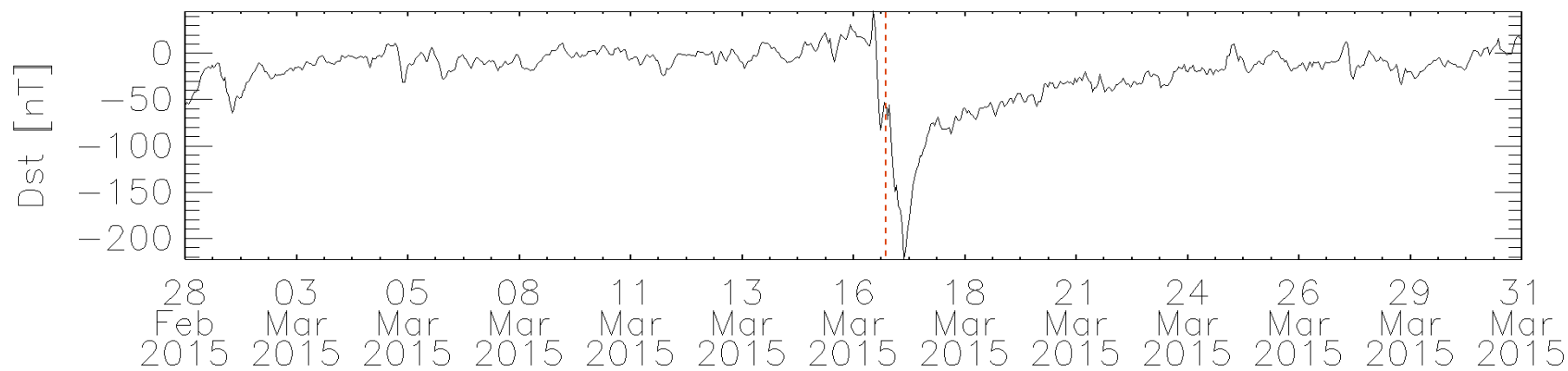
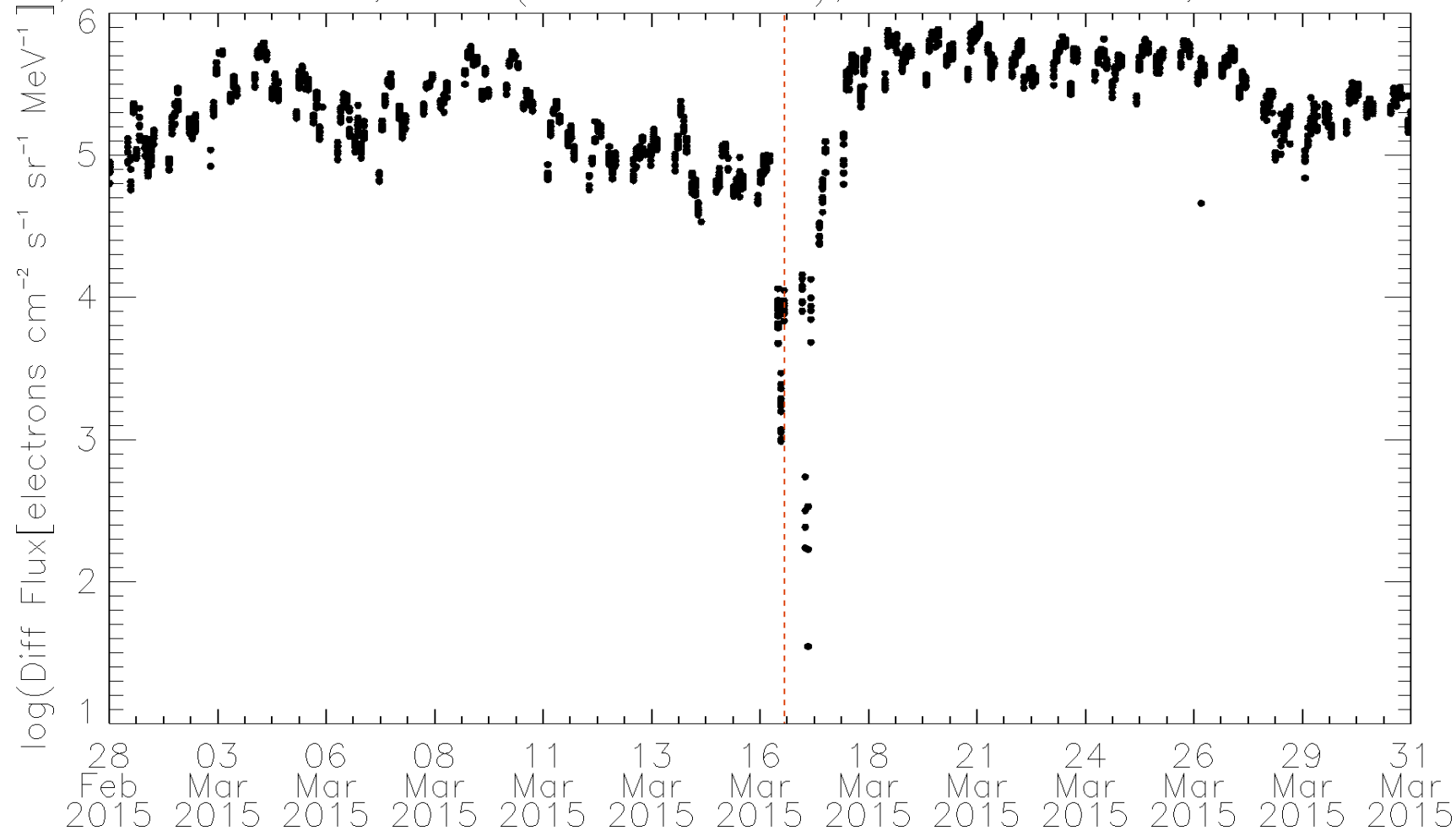




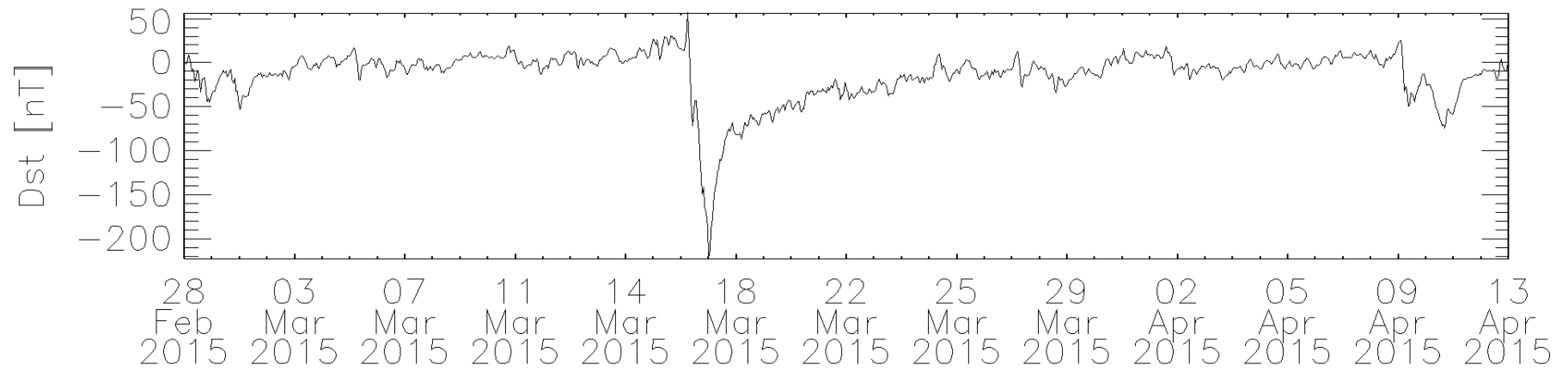
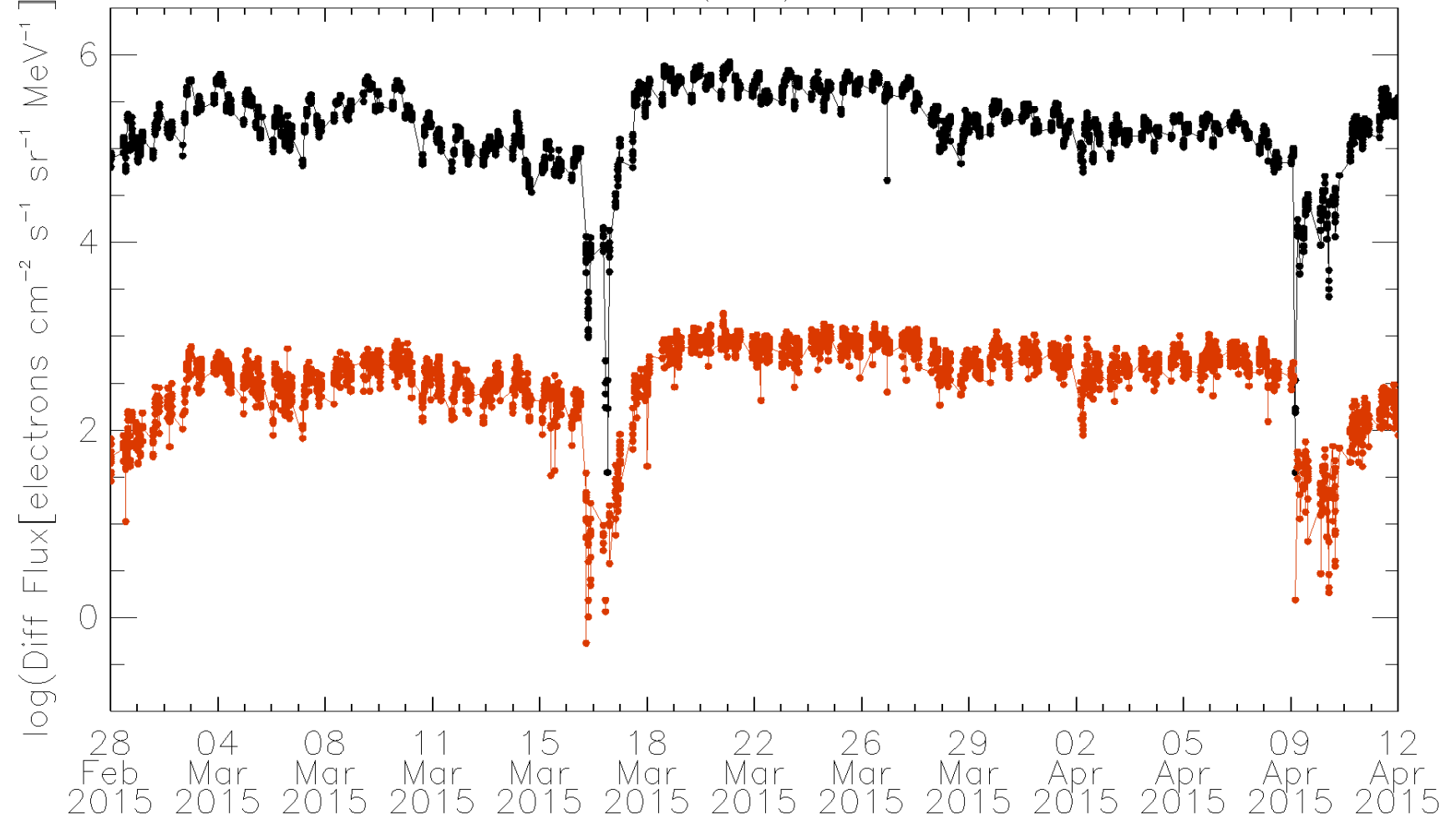
# EPT Lu (ch1 0.5-0.6 MeV) March 2015



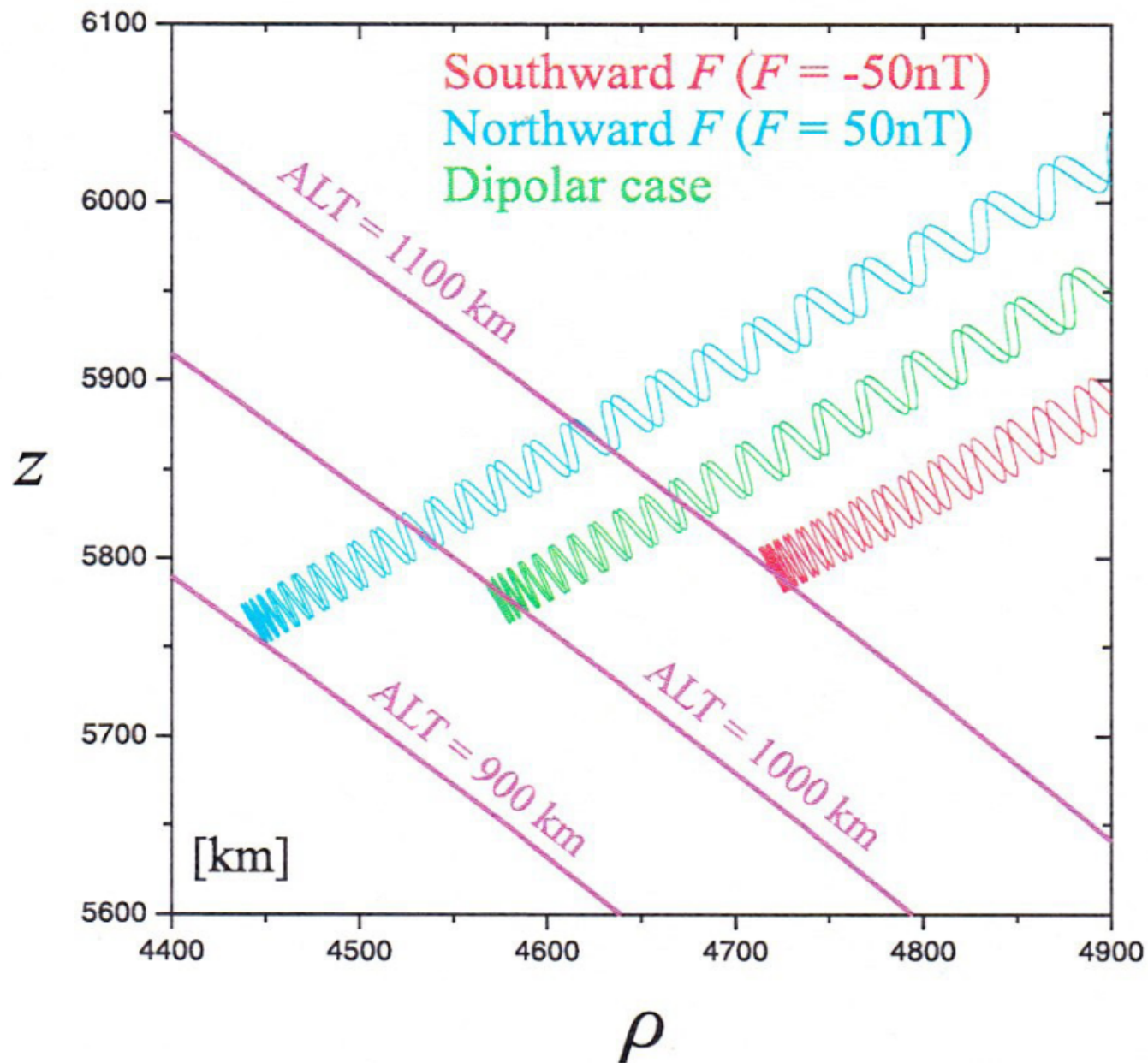
EP, Time vs Flux, ch1 (0.5–0.6 MeV), L=4.6–4.8 Re, B=0.25–0.3 G



EPT time vs flux ch1 and ch5 (red), L=4.6–4.8 Re, B=0.25–0.3 G

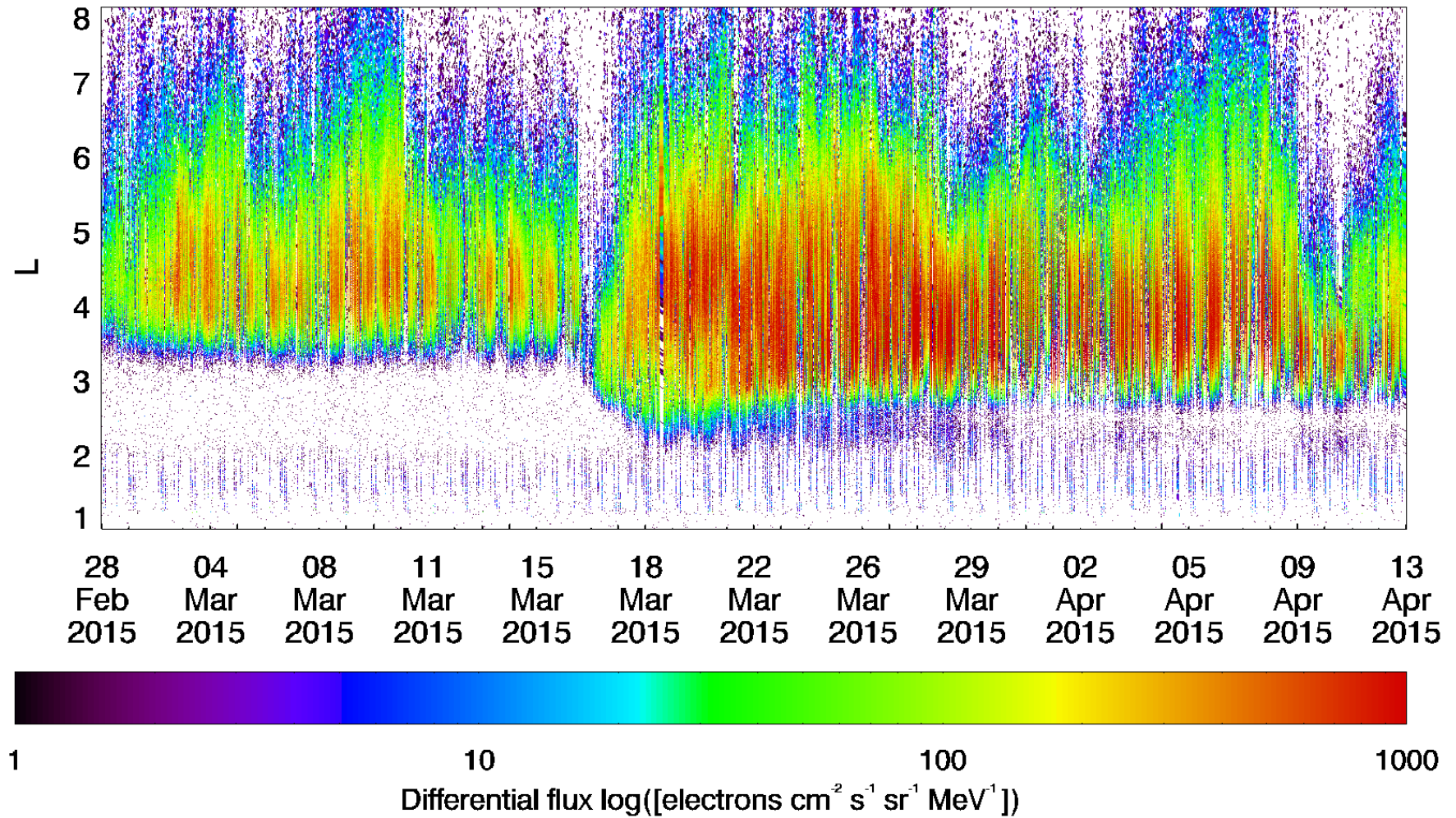


## Effect of a Bz-field (Dst) on trajectory of charged particle

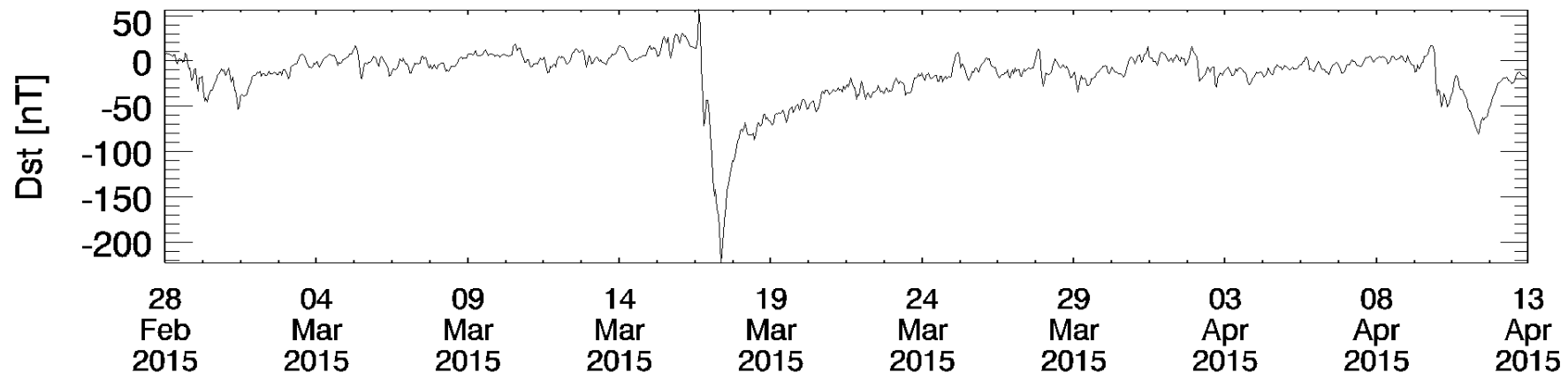
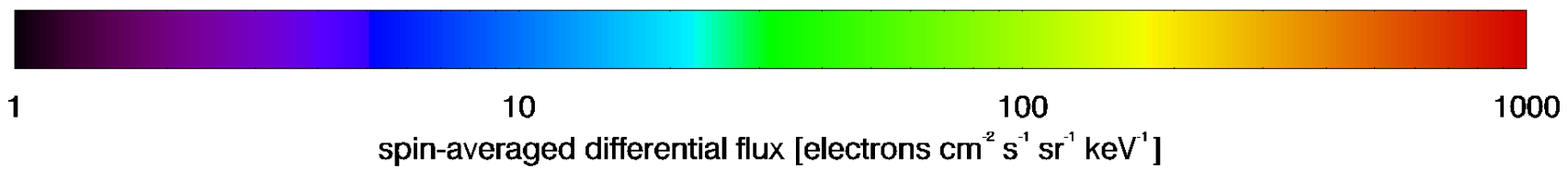
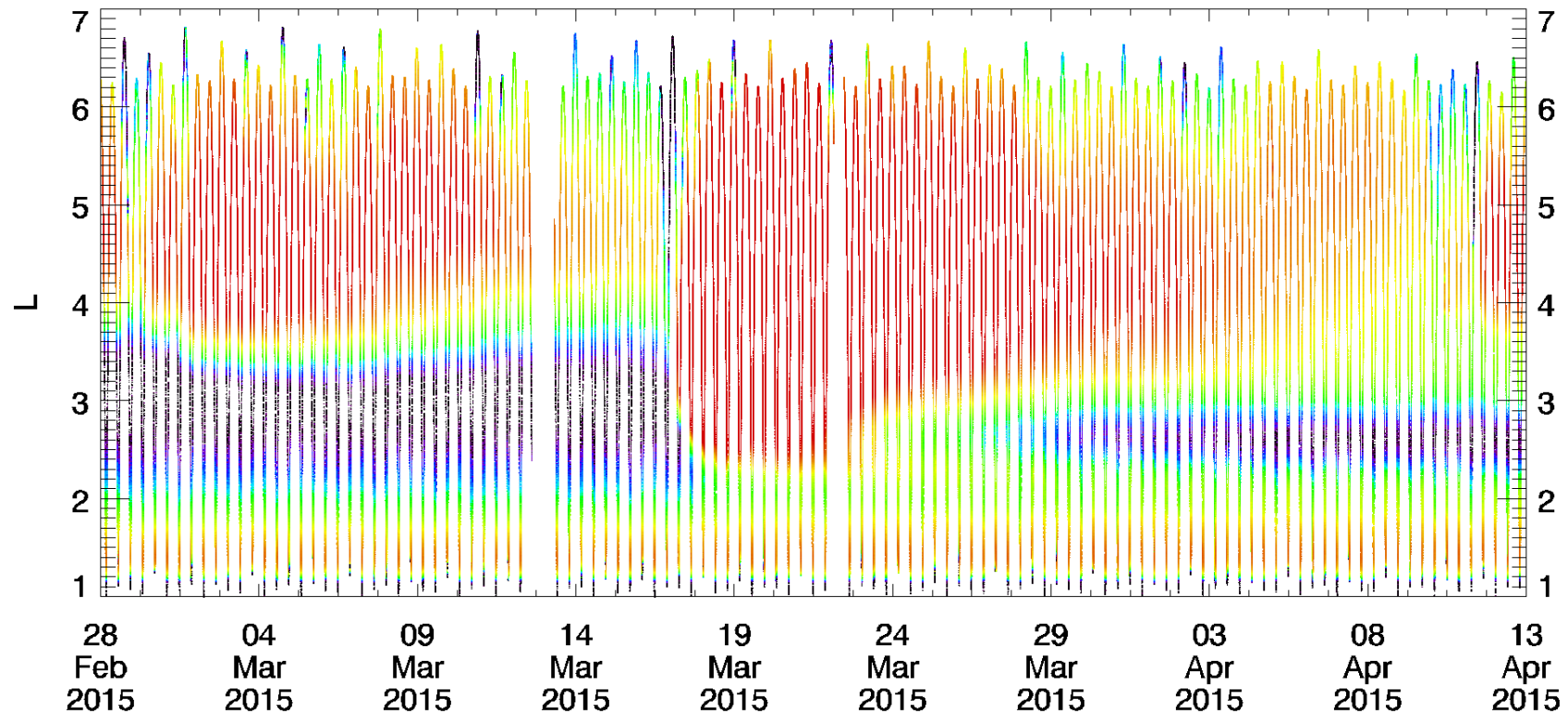




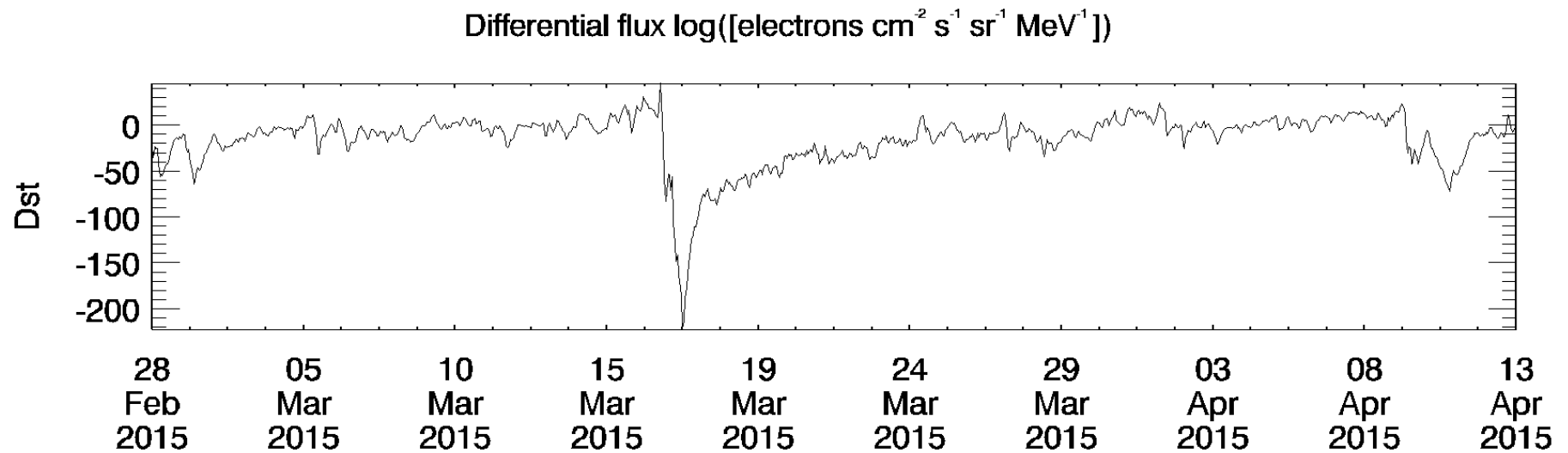
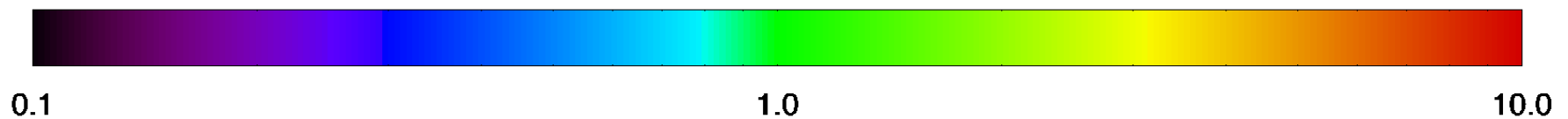
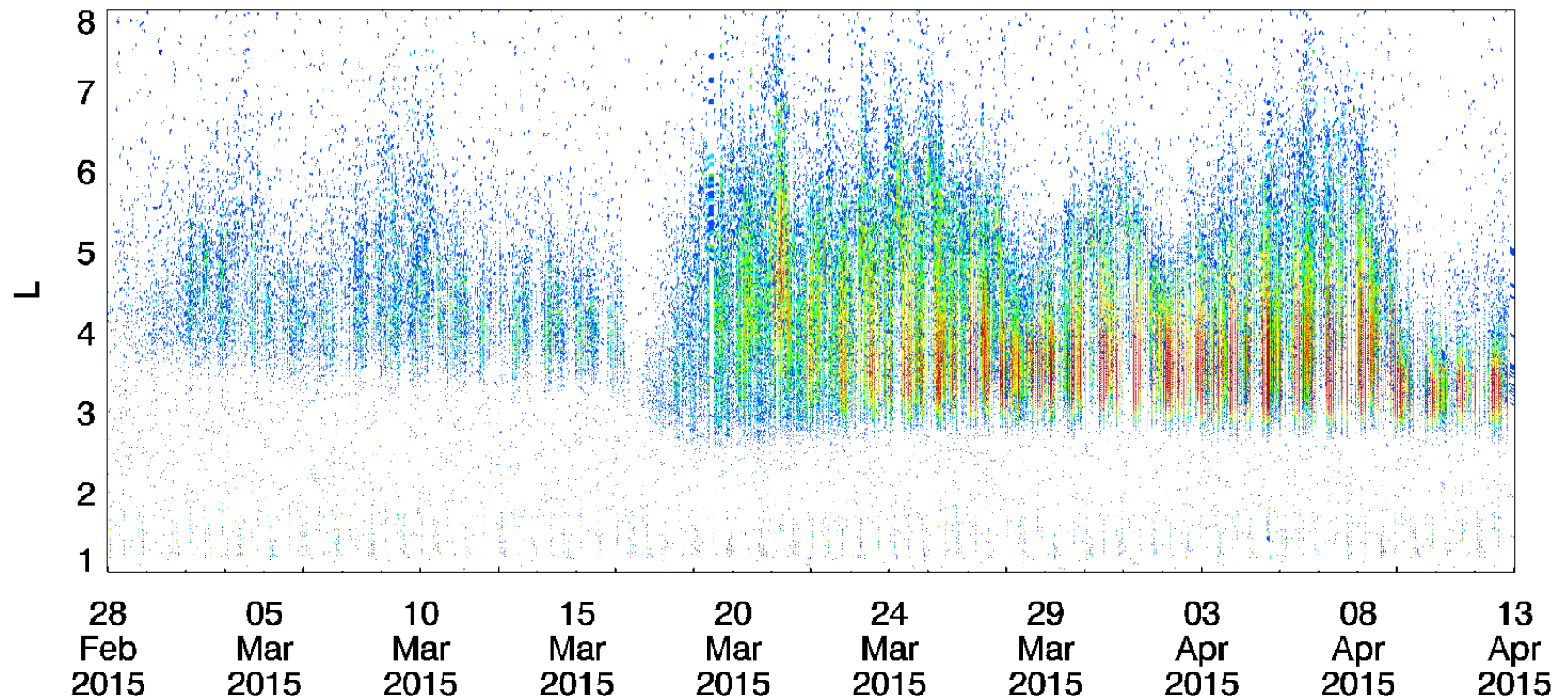
# EPT Lu (ch5 1.0-2.4 MeV) March - 13th April 2015



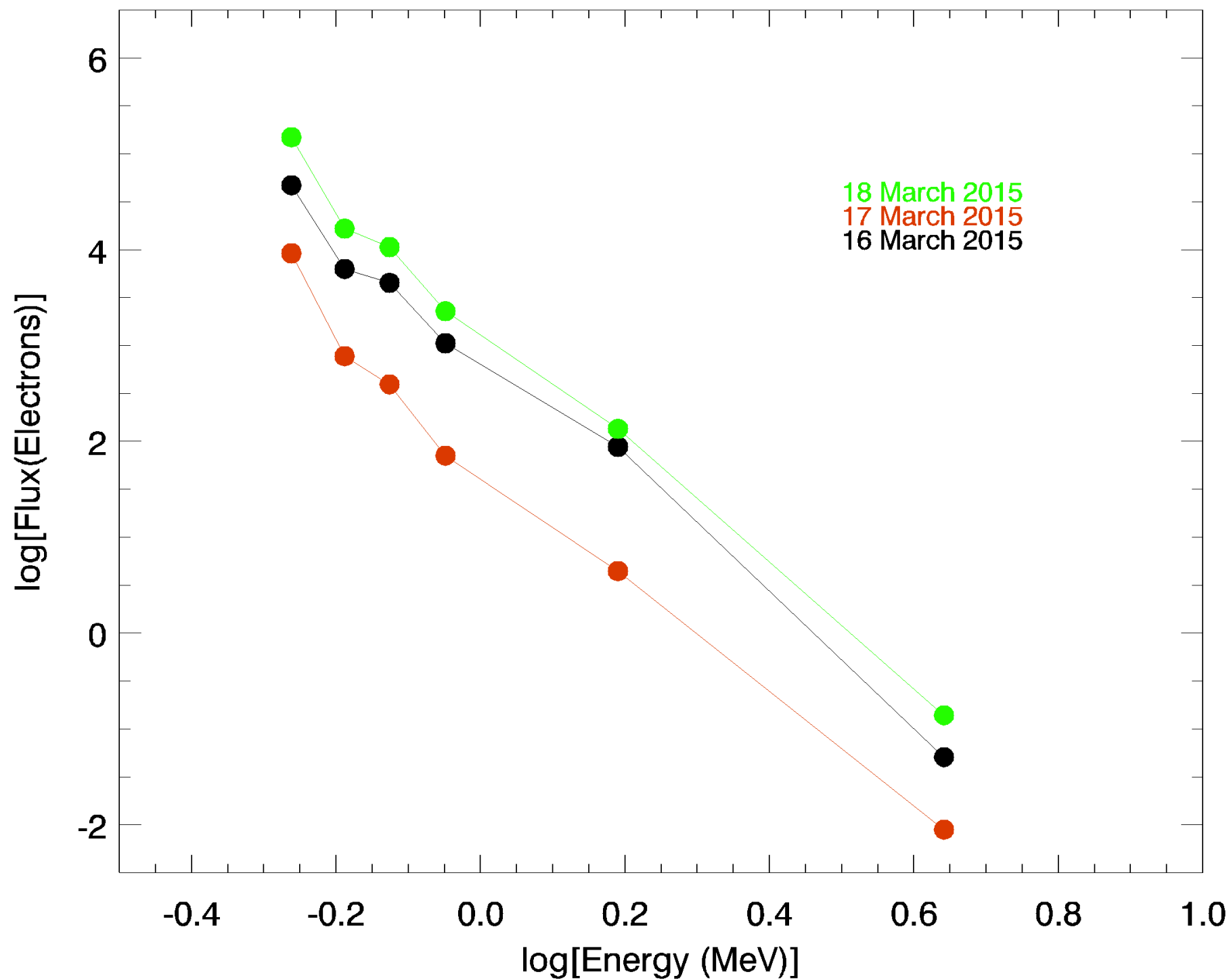
# RBSPB MagEIS channel index 15 energy 742.500000 keV



# EPT Lu (ch6: 2.4-8.0 MeV) March - 13th April 2015

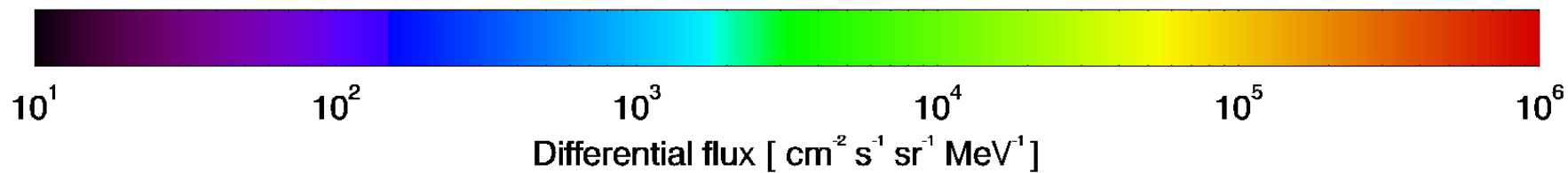
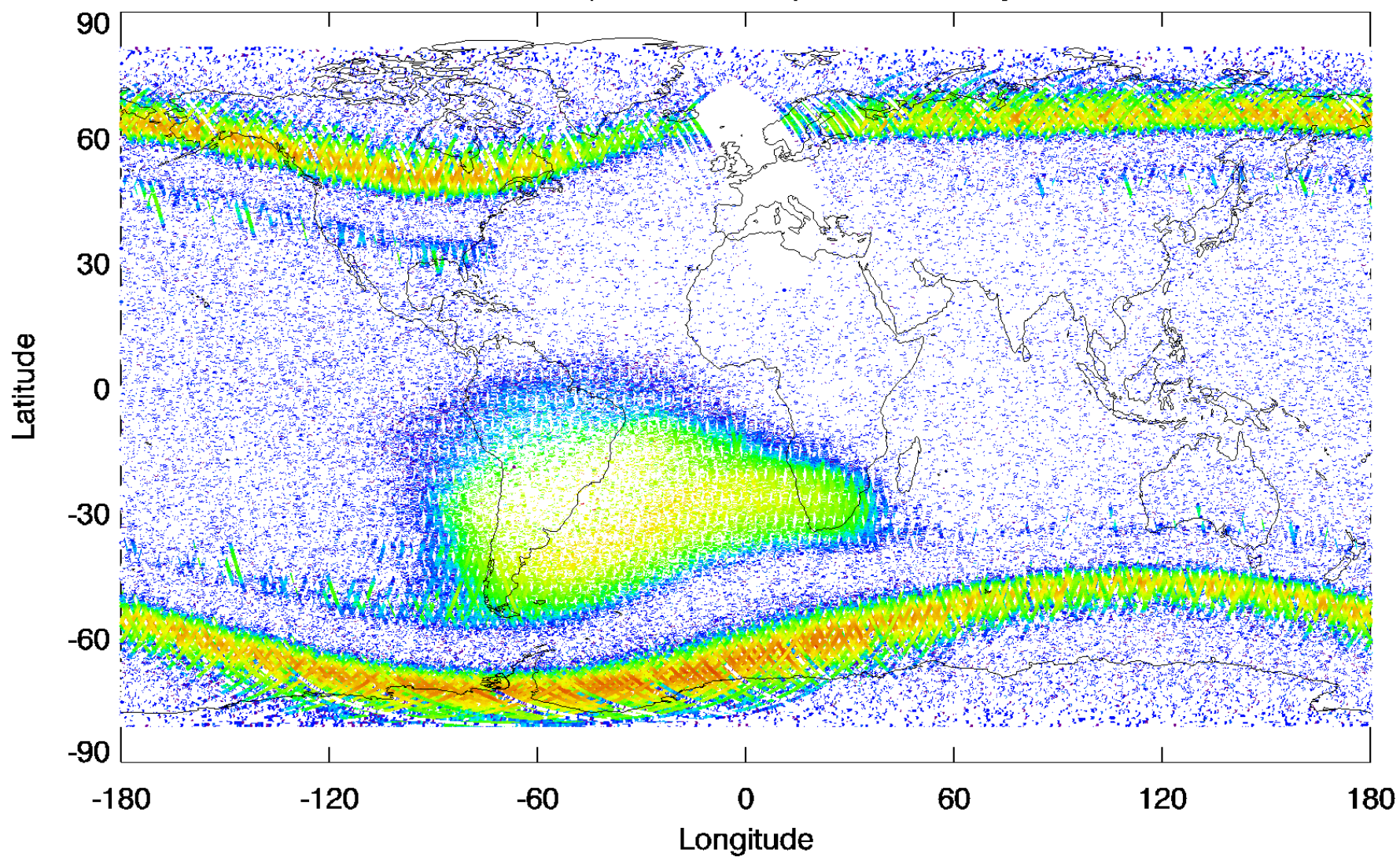


# Electrons Spectra, 17 March event, L=5.0-5.2, B=0.2-0.25

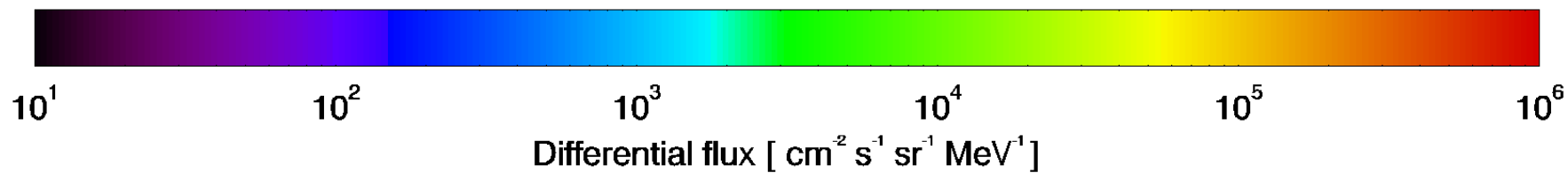
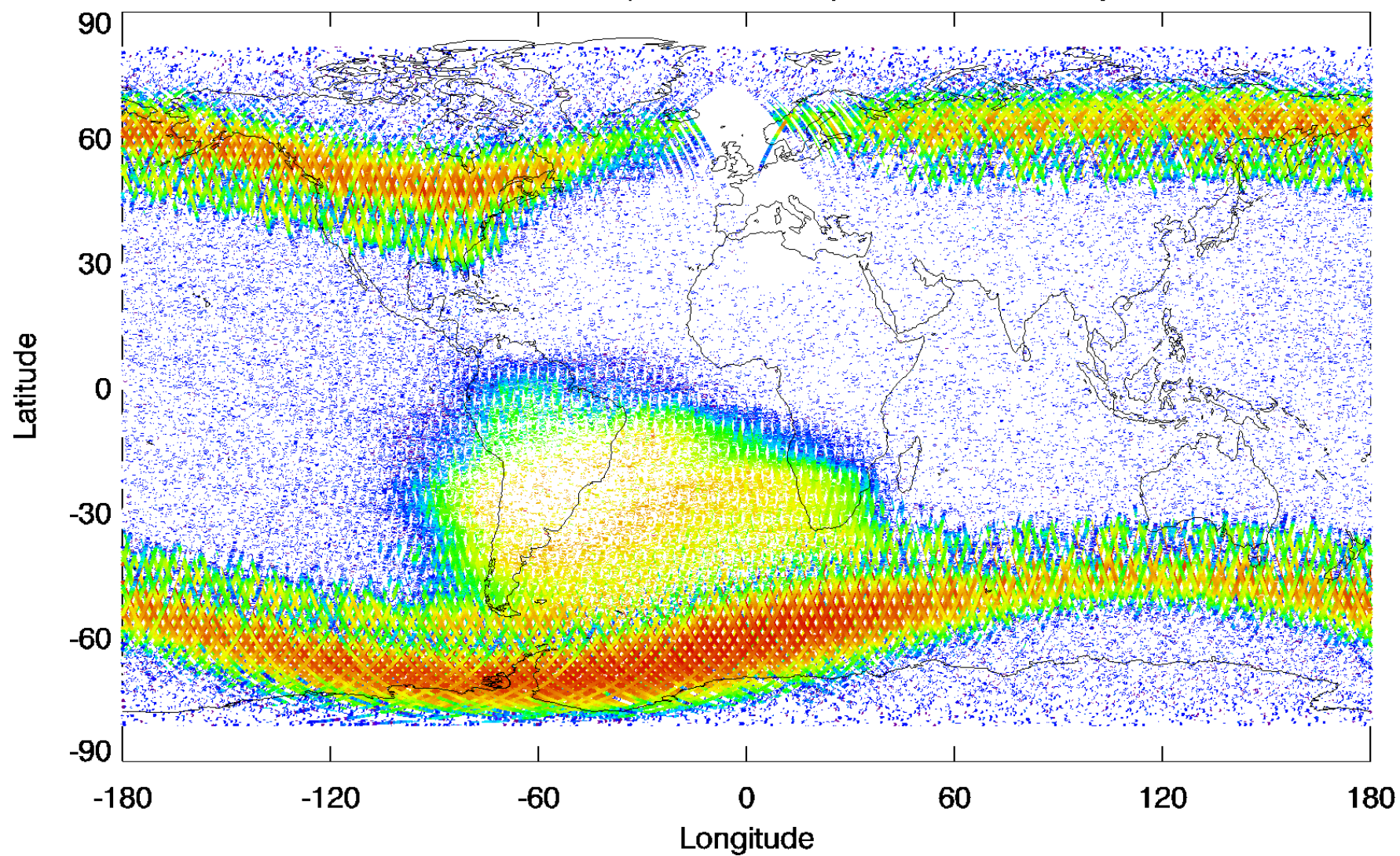




EPT Electron flux ch1 (0.5-0.6 MeV) 15 February-15 March 2015

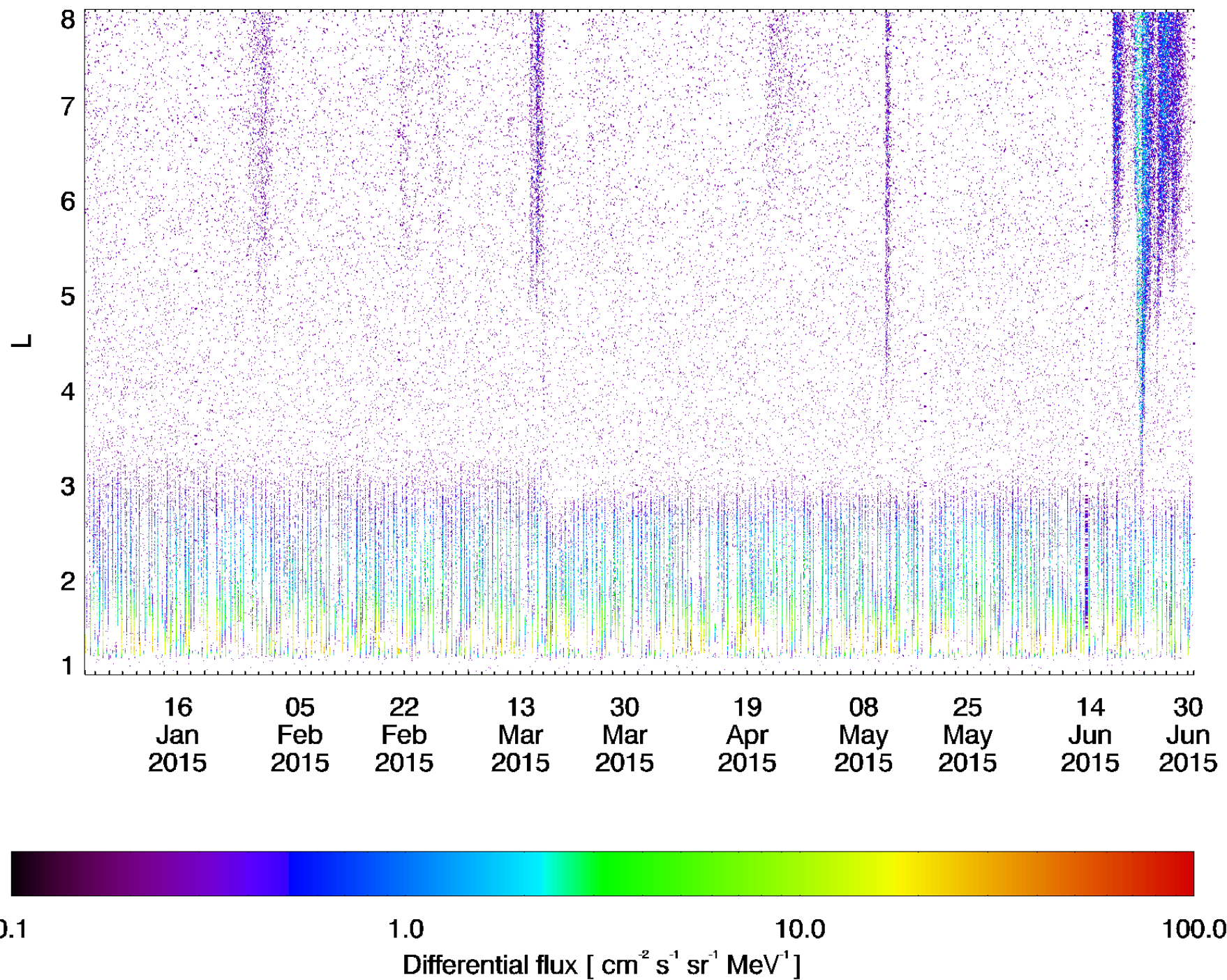


EPT Electron flux ch1 (0.5-0.6 MeV) 16 March-13 April 2015





EPT Lu Protons ch1 (9.5-13.0 MeV) January-June 2015



# Conclusions

The Energetic Particle Detector developed at CSR and BISA is Science-Instrument to study the RB

Uncontaminated spectra of Electrons, Protons, Alphas, ...  
High spectral time resolution flux measurements  
with  $< 5W$  and  $< 5\text{ kg}$

It is still operating nominally on polar LEO at 820 km alt.

Data are available from CSR, UCL, Louvain-La-neuve  
and BISA, Brussels

EPT data have been used for cross-calibration other RB instruments  
in orbit : e.g. VAP/MAGEIS



# Conclusions

In this presentation we have shown

how the fluxes and spectras of trapped electrons of different energies varies during a large Geomagnetic Storm (17 March 2015)  
(World Maps, L-vs-t flux variations, Energy spectras, ...).

To study Betatron effects on RB particles  
as well as the uplift of mirror points,  
acceleration/deceleration of charged particle during GS  
Field-aligned distributions of charged particles of all energies.

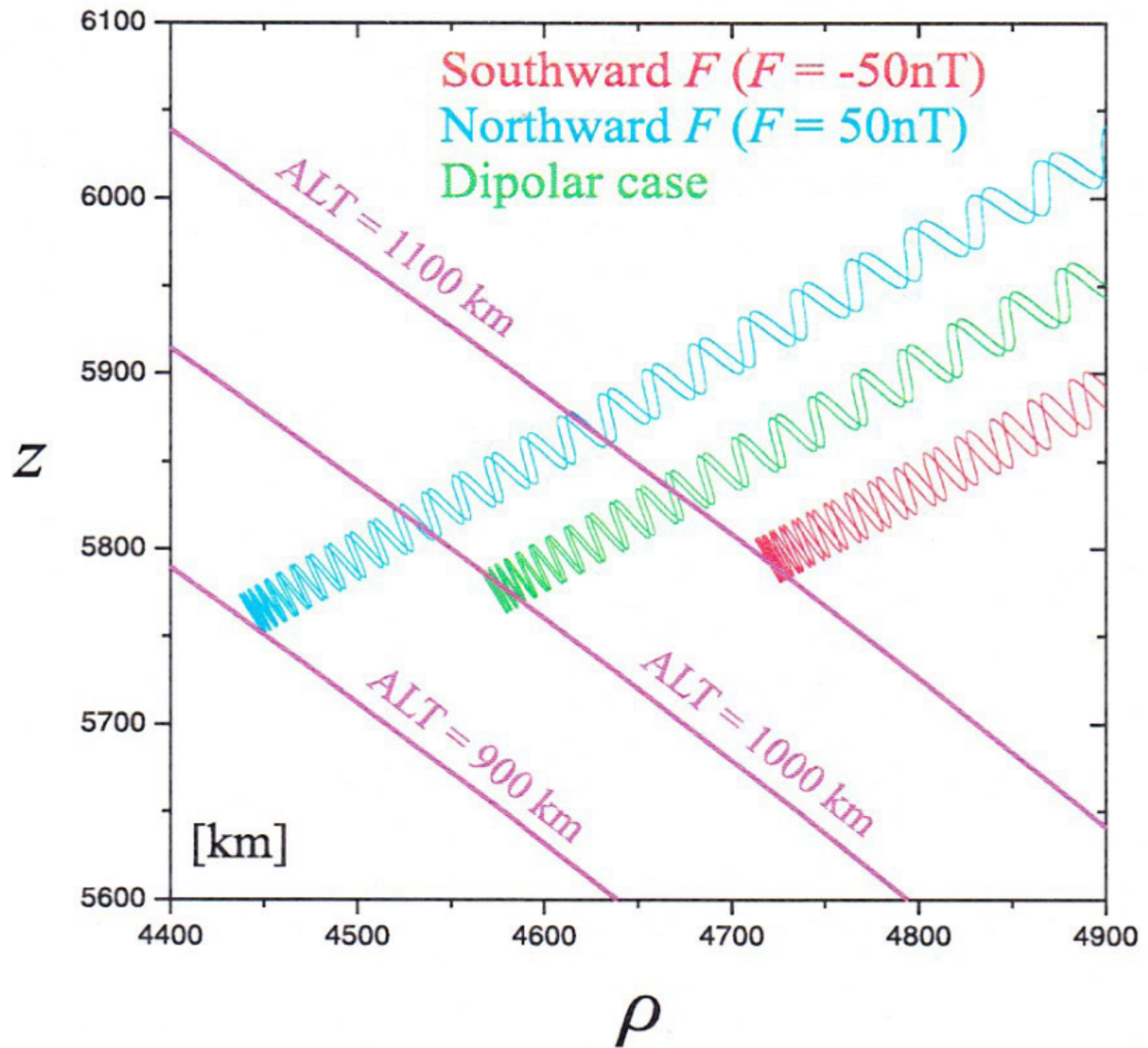
To study the new inhomogeneity discovered in SAA particle fluxes

Drop out of RB fluxes during the main phase of storms

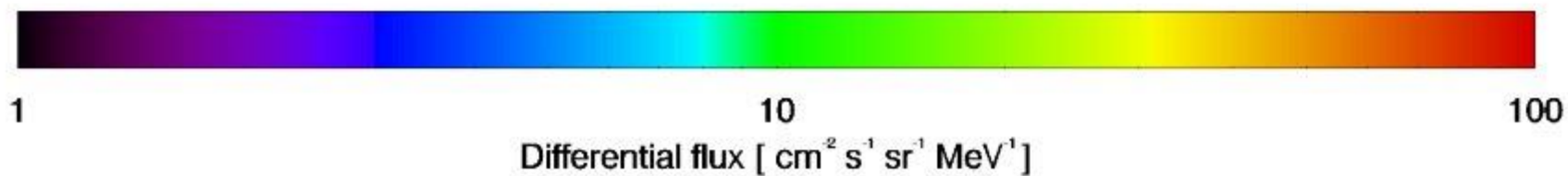
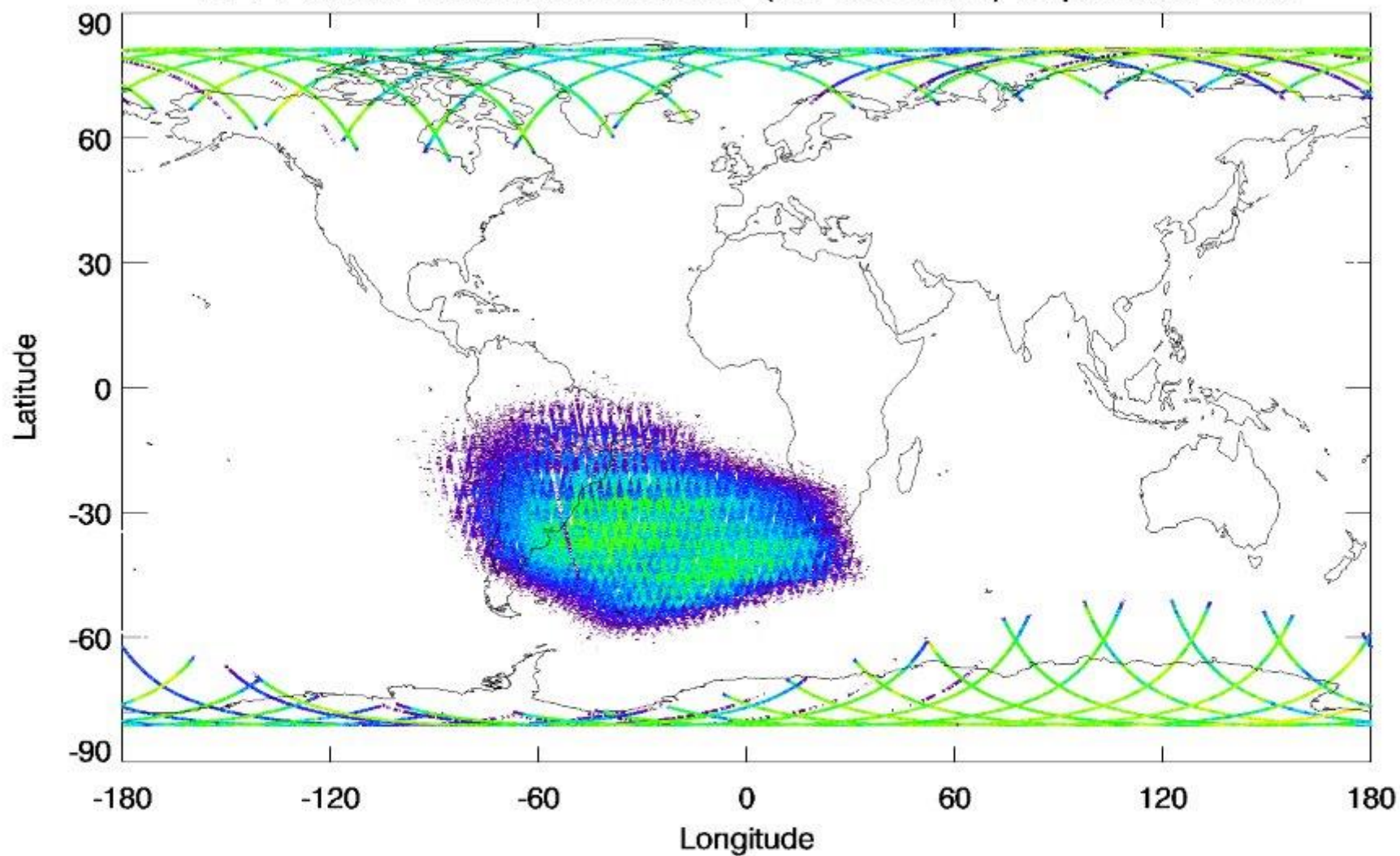
Penetration of SEP into magnetosphere.

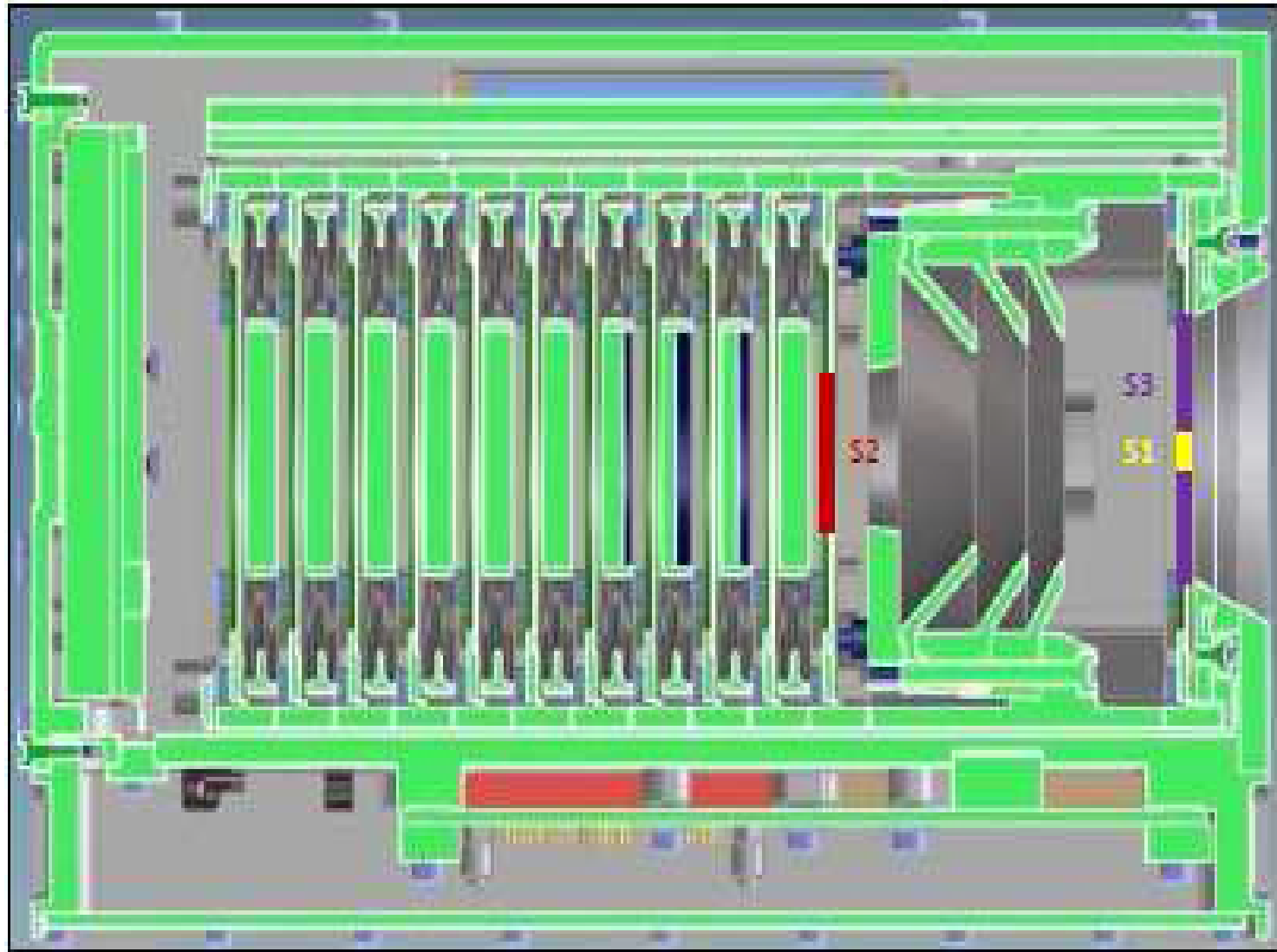
To study the flux enhancement during the Recovery Phases

The post-storm enhancement of the electron fluxes in the outer belt  
that may be explained by the combined effect of  
non-resonant pitch angle scattering and betatron acceleration  
during the main phase and recovery phase of storms



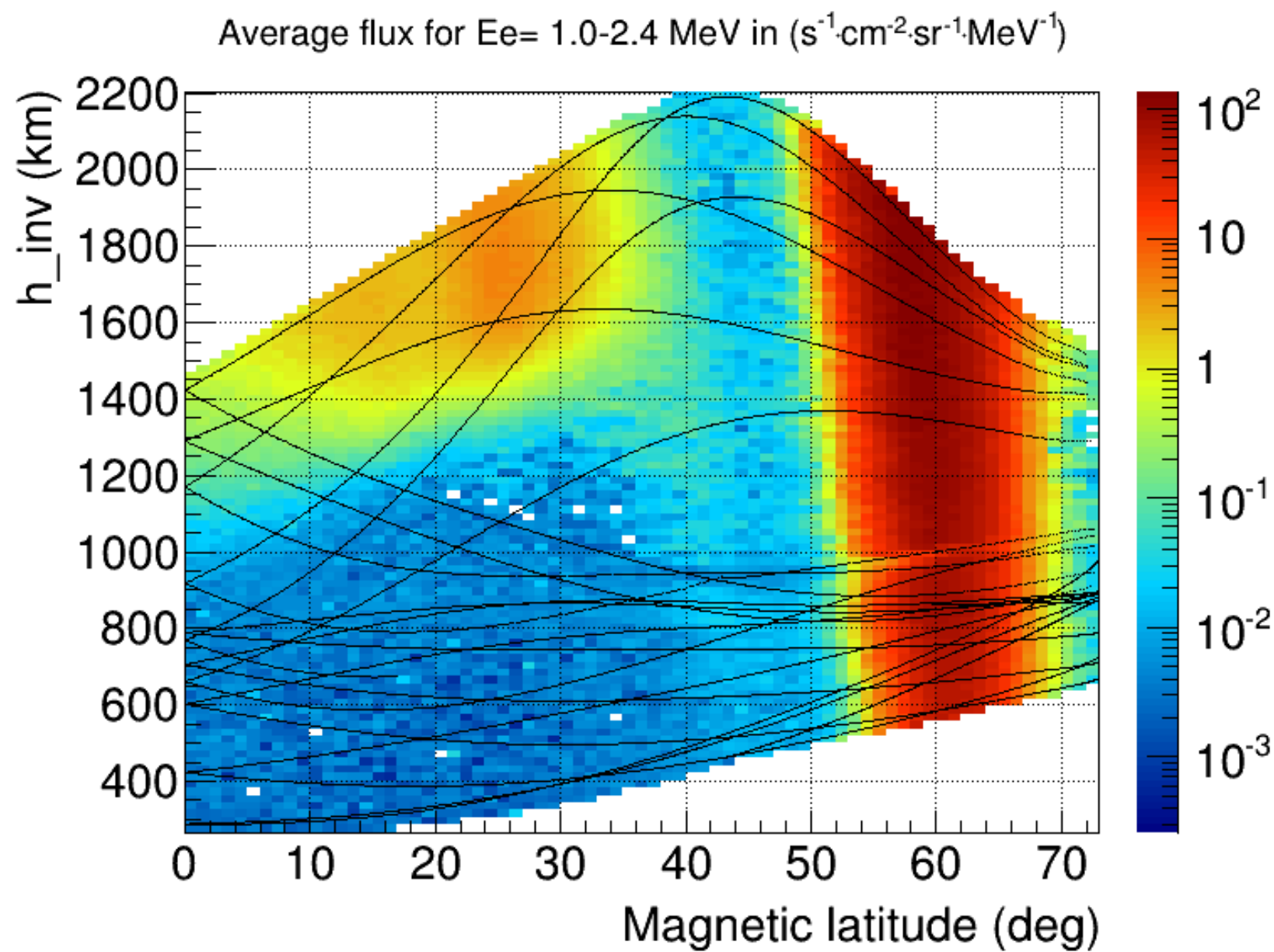
EPT Protons differential flux, Ch1 (9.5-13.0 MeV) September 2013



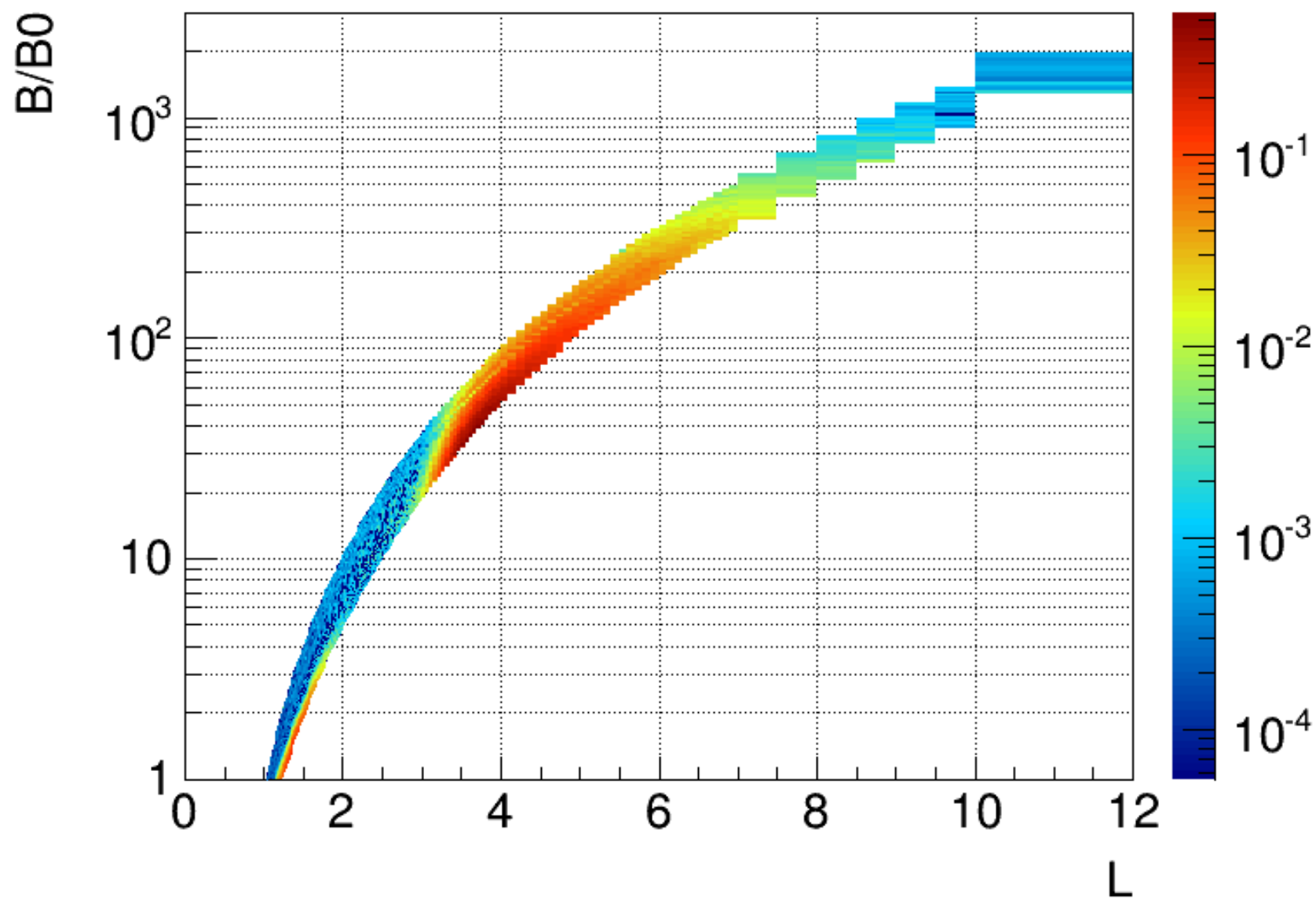




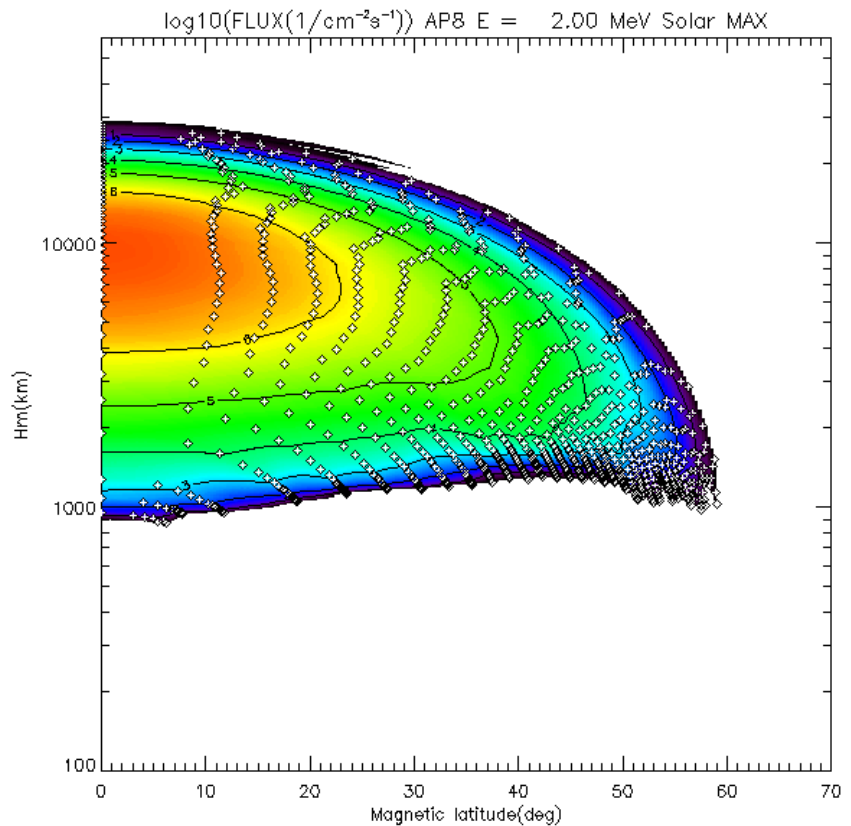
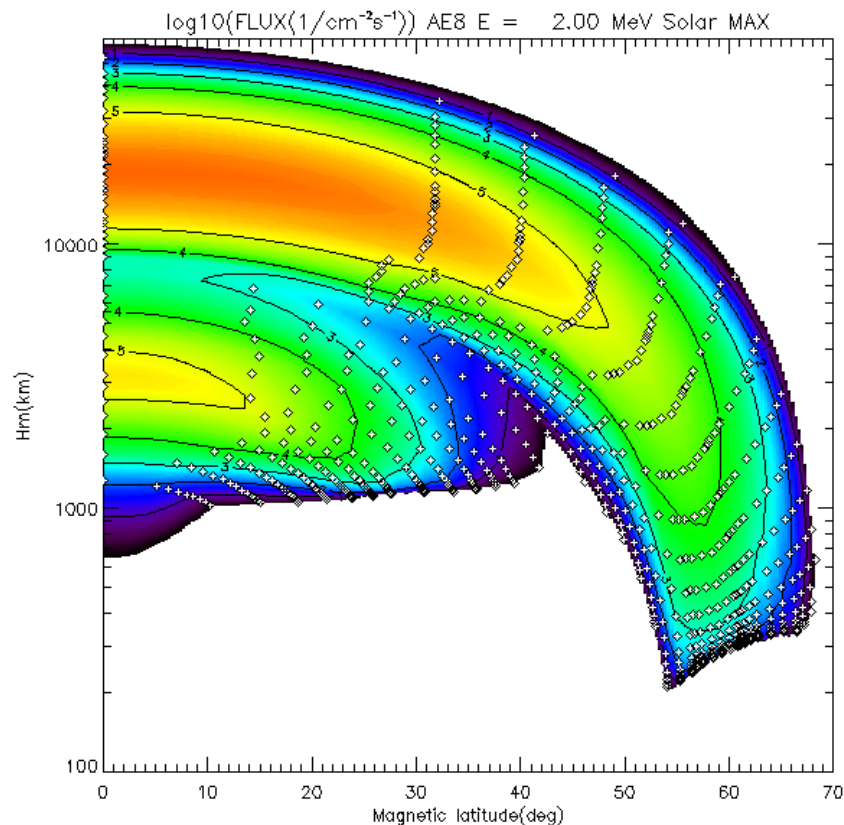
	Channel	Energy (MeV)	Channel	Energy (MeV)	Channel	Energy (MeV)	Channel	Energy (MeV)
LES	1	0.5 – 0.6	20	8.8 – **	39	35.5 – **	58	
	2	0.6 – 0.7	21		40		59	
	3	0.7 – 0.8	22		41		60	
	4	0.8 – 1.0	23	** – 9.5	42		61	
	5		24	9.5 - 13	43	** – 38	62	
	6		25		44	38 – 51	63	
	7		26		45		64	
	8		27		46		65	
	9		28		47		66	
HES	10	1.0 – 3.0	29	13 – 29	48	51 – 118	67	
	11	3.0 - .	30	29 – 61	49	118 – 245	68	
	12	4.0 _.	31	61 – 92	50	245 – 365	69	
	13		32	92 – 126	51	365 – 500	70	
	14		33	126 – 155	52	500 – 615	71	
	15		34	155 – 182	53	615 – 720	72	
	16		35	182 – 205	54	720 – 815	73	
	17		36	205 – 227	55	815 – 900	74	
	18		37	227 – 248	56	900 – 980	75	
	19		38	>248	57	>980	76	



Average flux for  $E_e = 2.4\text{--}8.0$  MeV in  $(\text{s}^{-1}\cdot\text{cm}^{-2}\cdot\text{sr}^{-1}\cdot\text{MeV}^{-1})$



# Log(Invariant Altitude) vs Magnetic Latitude MAX & E = 2 MeV







## Extended Störmer Potential : $V(r_{eq})$

for a set of b-values [ # IMF]

