

Recent results on kaon physics from OKA experiment

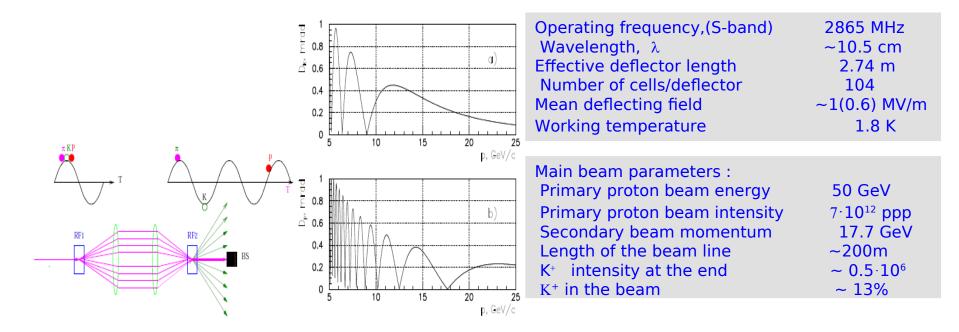
V. Obraztsov, SRC "Kurchatov Institute" - IHEP, Protvino On behalf of the «OKA» collaboration (IHEP-INR-JINR) "The 7-th international conference on particle physics and astrophysics", Moscow, 22.10-25.10 2024

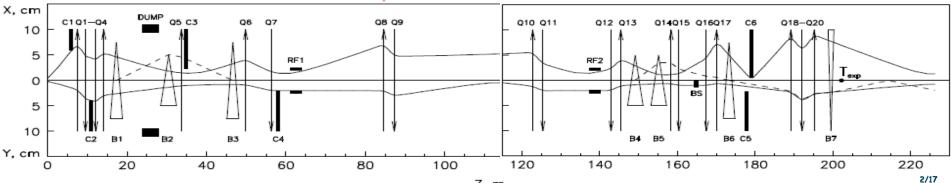
The talk layout

- OKA beam, detector
- Search for the ALP in the $K^+ \rightarrow \pi^+ \pi^0$ a decay
- $K^+ \rightarrow \mu^+ \nu \gamma$, measurement of $F_v F_A$
- $K^+ \rightarrow e^+ \nu \pi^0 \gamma$, $K^+ \rightarrow \mu^+ \nu \pi^0 \gamma$, $\chi PT O(p^4)$, search for T(CP) violation
- $K^+ \rightarrow \pi^+ \pi^+ \pi^- \gamma$, $\pi^+ \pi^0 \pi^0 \gamma$ test of χPT

OKA: The experiment with RF-separated K[±] beam @U-70

RF separation with Panofsky scheme is realised. It uses two SC RF deflectors. Sophisticated cryogenic system, built at IHEP provides superfluid He for cavities cooling.

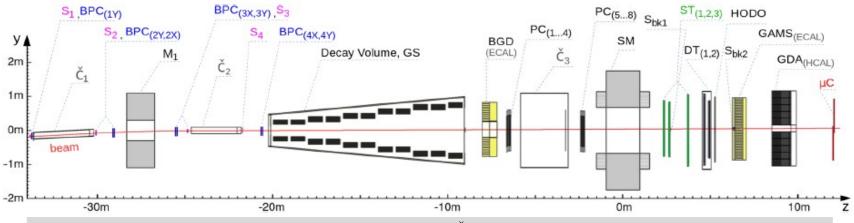




Z, m

° DKA

OKA detector



- 1. Beam spectrometer: 1mm pitch BPC ~1500 channels; Sc and \check{C} counters
- 2. Decay volume with Veto system:

L=11m; Veto: 670 Lead-Scintillator sandwiches 20* (5mm Sc+1.5 mmPb), WLS readout

- 3. PC's, ST's and DT's for magnetic spectrometer:
- ~5000 ch. PC (2 mm pitch) + 1300 DT (1 and 3 cm)
- 4. Pad(Matrix) Hodoscope ~300 ch. WLS+SiPM readout
- 5. Magnet: aperture 200*140 cm²
- 6. Gamma detectors: GAMS2000, BGD EM cal. ~ 4000 LG.
- 7. Muon identification: GDA-100 HCAL + 4 muon counters (μ C) behind
- 8. For some runs Cu target inside decay volume was used: Ø=8 cm, t=2mm and C3 big Cerenkov counter

Main triggers	$S_1 \cdot S_2 \cdot S_3 \cdot \overline{C_1} \cdot C_2 \cdot \overline{S_{bk}} \cdot \left(\Sigma_{GAMS} > 2.5 GeV \right) \cup \left(2 \leq MH \leq 4 \right)$					
Prescaled triggers	$S_1 \cdot S_2 \cdot S_3 \cdot C_1 \cdot C_2 \cdot S_{bk} / 10 \qquad S_1 \cdot S_2 \cdot S_3 \cdot C_1 \cdot C_2 \cdot S_{bk} \cdot \mu C / 4$					
	Run's in 2010-2013, 2016, 2018 $N_{K} \sim 5 \ge 10^{10}$					
	Main directions of the data analysis:					

 $K^{+} \rightarrow e^{+} \nu \pi^{0}, K^{+} \rightarrow \mu^{+} \nu_{s}, K^{+} Cu \rightarrow K^{+} \pi^{0} Cu, K^{+} \rightarrow \pi^{+} \pi^{0} a, K^{+} \rightarrow \mu^{+} \nu \gamma, K^{+} \rightarrow e^{+} \nu \pi^{0} \gamma, K^{+} \rightarrow \mu^{+} \nu \pi^{0} \gamma, K^{+} \rightarrow \pi^{+} \pi^{-} \pi^{0} \gamma, K^{+} \rightarrow \pi^{+} \pi^{0} \pi^{0} \gamma, K^{+} \rightarrow \mu^{+} \nu \gamma, K^{+} \rightarrow \mu^{+} \mu^{+}$





General view of the OKA setup

«OKA» setup



ST, DT chambers, Matrix Hodoscope, ECAL



Decay volume Veto System



RF deflector in the beamline

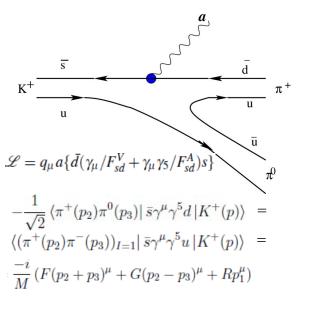


Liquid He lines



Tail of the beam line

Search for the ALP in $K^+ \rightarrow \pi^+\pi^0$ a decay



F,G,R from $K^+ \rightarrow \pi^+\pi^- l\nu$ (Kl4)

The QCD Axion is a hypothetical pseudoscalar particle, invented to solve the strong CP problem. It's properties are described by the decay constant f_a , related to Peccei-Quinn symmetry braking scale Λ_{PQ} : $f_a = \Lambda_{PQ}/4\pi$. The QCD axion mass $m_a = m_{\pi} f_{\pi}/f_a$.

 $a \rightarrow \gamma \gamma$; $\tau_a = 2^8 \pi^3 f_a^2 / (\alpha m_a^3)$. If axion is dark matter $\rightarrow \tau_a \ge 13.8 \text{ Gyr} \rightarrow m_a \le 10 \text{ eV}$. For axion-like particles (ALP) m_{ALP} is not set by QCD only \rightarrow two free parameters: m_{ALP} , $f_{ALP} = m_{ALP} < 1 \text{ GeV}$.

Axion may have vector and/or axial couplings to quark currents, in particular to sd FCNC P-conservation \rightarrow vector $K^+ \rightarrow \pi^+ a$ axial $K^+ \rightarrow \pi^+ \pi^0 a$ $\mathscr{L} = q_\mu a \{ \overline{d}(\gamma_\mu / F_{sd}^V + \gamma_\mu \gamma_5 / F_{sd}^A) s \}$

We assume that the axion decays outside the setup.

Start 1	<u>from 3.65 10⁹ events</u> Common cuts for $K^+ \to \pi^+ \pi^0$ a and $K^+ \to \pi^+ \pi^0$					
•	1 beam track , 1 secondary track $\theta > 4$ mrad , vertex matching CDA < 1.25 cm.					
•	no extra track segments behind the SM magnet					
•	vertex inside the DV.					
•	17. $0 < p_{beam} < 18.6 \text{ GeV}$					
•	number of showers in GAMS or BGD not associated with $track = 2$					
•	π^0 identification $ m_{\gamma\gamma} - m_{\pi0} < 15 \text{ MeV}$					

After selections 44.5 $10^6 \text{ K}^+ \rightarrow \pi^+ \pi^0$

Search for the axion in $K^+ \rightarrow \pi^+\pi^0$ a decay

In order to disentangle $K^+ \rightarrow \pi^+\pi^0 a$ from $K^+ \rightarrow \pi^+\pi^0(\gamma)$, $K^+ \rightarrow \pi^+\pi^0\pi^0$, $K^+ \rightarrow e^+\nu\pi^0$, $K^+ \rightarrow \mu^+\nu\pi^0$

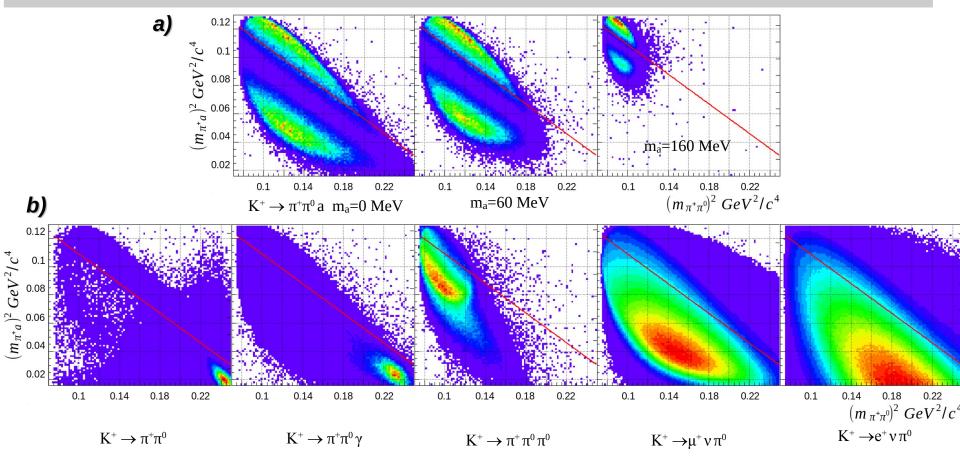
- $E_{mis} = E_{K+} E_{\pi+} E_{\pi0} > 2.8 \text{ GeV}$ Cut on missing energy
 - $P_{\pi^+}^* < 150 \text{ MeV}$, $P_{\pi^0}^* < 189 \text{ MeV}$ Cuts on the momenta of pions in the K⁺ rest frame againsts K⁺ $\rightarrow \pi^+\pi^0$
 - No signal in muon counters μC to suppress $K^+ \rightarrow \mu^+ \nu \pi^0$

E- the energy of the shower, assosiated with the track, againsts $K^+ \rightarrow e^+ v \pi^0$

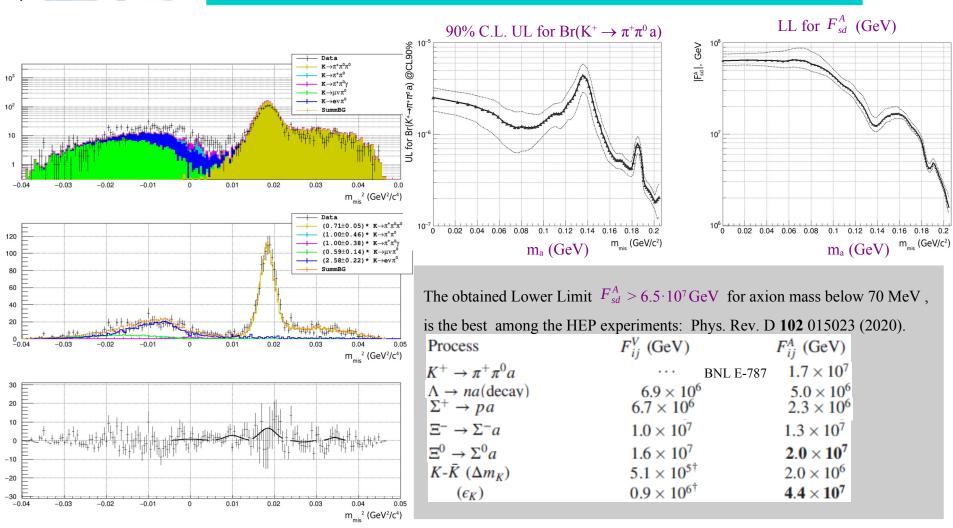
- track is identified as π^+ in GAMS or in GDA-100
- E_{GS} < 100 MeV

 $E/p \le 0.83$

the cut on the energy in the guard system - against $K^{\scriptscriptstyle +} \to \pi^{\scriptscriptstyle +} \pi^{\scriptscriptstyle 0} \pi^{\scriptscriptstyle 0}$



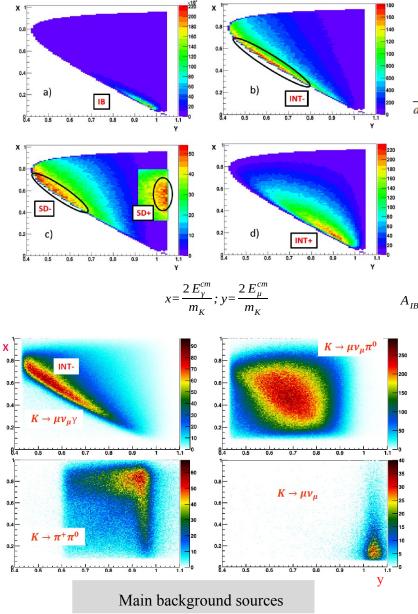
Search for the axion in $K^+ \rightarrow \pi^+ \pi^0 a$ decay

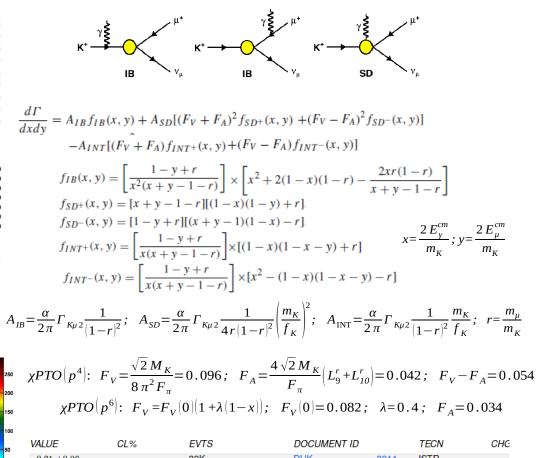


Supernova bound: In the neutron stars(NS) n, p, e, Λ coexist. $\Lambda \rightarrow$ n a new cooling mechanism of NS. Maximum during few seconds after SN explosion, when protoneutron star reaches T~ 0.1 MeV SN1987A $F_{sd}^{A}, F_{sd}^{V} > 10^{9}$ GeV Model dependent !



Study of the $K^+ \rightarrow \mu^+ \nu \gamma$ decay





VALUE	C	L%	EVIS	DOCUMENT ID		TECN	CHG	
-0.21 ± 0.06			22K	DUK	2011	ISTR	-	
••• We do not use the following data for averages, fits, limits, etc. •••								
-0.24 to 0.04	1 9	0	2588	ADLER	2000B	B787	+	
-2.2 to 0.6	9	0		DEMIDOV	1990	XEBC		
-2.5 to 0.3	9	0		AKIBA	1985	SPEC		
References:								
DUK	2011	PL B695 59	Extraction of Kaon	Formfactors from H	$K^- o \mu u \gamma$ C	ecay at ISTRA+	Setup	
ADLER	2000B	PRL 85 2256	Measurement of S	tructure-Dependent	$K^+ ightarrow \mu^+ u$	$\gamma_{\mu}\gamma$ Decay		
DEMIDOV	1990	SJNP 52 1006	Measurement of the	he $K^+ o \mu^+ u \gamma$ Deca	y Probabili	ty		
AKIBA	1985	PR D32 2911	A Study of the Rad	liative Decay $K^+ ightarrow$	$\mu^+ u_\mu\gamma$			



$K^+ \rightarrow \mu^+ \nu \gamma$ selection and analysis

I beam K⁺ track

Q

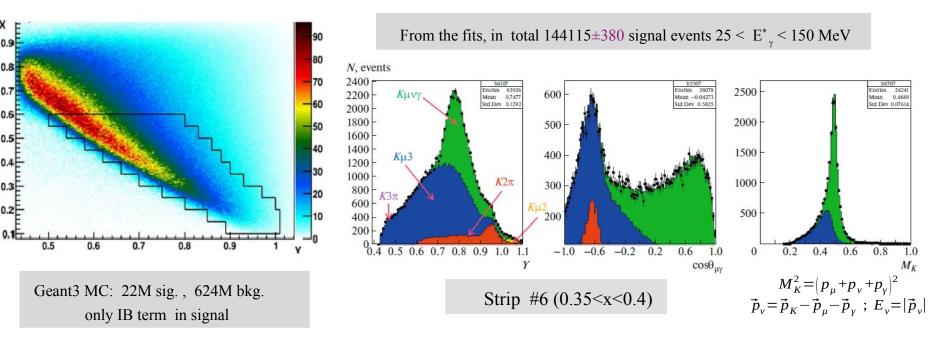
- 1 secondary track identified as μ in GAMS, GDA-100 and MC
- Decay vertex inside DV
- 1 e.m. shower in GAMS with E > 1GeV not associated with charged track
- $E_{GS} < 10 \text{ MeV}$; $E_{EGS} < 100 \text{ MeV}$

Fit procedure

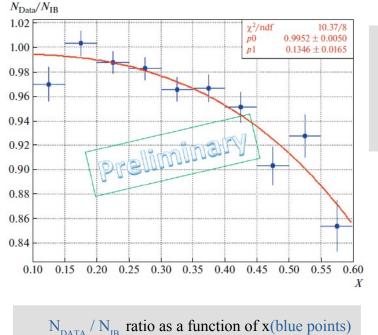
• x,y region is devided into strips $\Delta x=0.05$ (~12 MeV)

Plot y-disribution; select cuts $\{y_{\min}, y_{\max}\}$; plot ; select $\cos_{\min} \operatorname{cut}$; Plot M_K
Simultaneous fit of the 3 histograms, parameters- N ... N.

- Simultaneous fit of the 3 histograms, parameters- N_{sig} , N_{bkg} COS $\theta_{\mu\gamma}$ both signal(IB) and background shapes are taken from MC
- to correctly estimate errors, fit only M_{K}^{-} plot with initial parameters of the simultaneous fit







Red line is the result of the fit with $p_{sig}(x)=p0(1+p1\cdot\phi_{INT}(x)/\phi_{IB}(x))$ p0 is the normalization $p0=0.9952 \pm 0.005$; $p1=Fv-Fa=0.135\pm0.017$ $\phi_{INT}(x)$ - x-distribution of reconstructed MC-signal weighted events $w_{INT}=(M_K/F_K) f_{INT}(x_{true},y_{true}); \phi_{IB}(x)$ - the same with $w_{IB}=f_{IB}(x_{true},y_{true})$

Systematics

• Non-ideal description of signal and background by MC: $1.3 < \chi^2/NDF < 1.7$

Stat. errors in each bin of N_{DATA}/N_{IB} -plot scaled with $\sqrt{(\chi^2/NDF)}$. New value Fv-Fa=0.138±0.026 (nominal 0.134±0.021) $\rightarrow \sigma_{shape} = 0.012$ Width of -x- strips: Fv-Fa calculation repeated for 2 different values of width $\Delta x=0.035$, $\Delta x=0.07$ (nominal 0.05) $\rightarrow \sigma_{\Delta x} = 0.008$ The fit range in x (number of -x- strips): remove one or two bins on the left(right) edge. $\rightarrow \sigma_{x} = 0.005$ -y- limit in the strips: instead of maximizing S/ $\sqrt{(S+B)}$ use FWHM from the signal MC $\rightarrow \sigma_{y} = 0.005$ Effect of INT+: INT+ term is added to N_{DATA}/N_{IB} fit. The BNL E787 value |Fv+Fa|= 0.165±0.013 is used (±0.178) $\rightarrow \sigma_{SYS} = 0.024$



 $K^+ \rightarrow \mu^+ \nu \gamma$ decay Fv-Fa

"OKA" $F_{V}-F_{A} = 0.135 \pm 0.017_{stat} \pm 0.024_{syst}$ $\chi PT O(p^{4})$ $F_{V} = \frac{\sqrt{2}M_{K}}{8\pi^{2}F_{\pi}} = 0.096 ; F_{A} = \frac{4\sqrt{2}M_{K}}{F_{\pi}} (L_{9}^{r} + L_{10}^{r}) = 0.042$ $F_{V}-F_{A} = 0.054$ 2.8 σ difference

Lattice calcullations: $F_v - F_A = (0.083 \pm 0.013) - (0.019 \pm 0.012) \cdot x_\gamma$ Phys. Rev. D 103, 014502 (2021) (1.8 σ)

The measured value is in a reasonable agreement with ISTRA+ result: And with a (model dependent) result of BNL E865 ($K^+ \rightarrow \mu^+ \nu \ e^+ e^- + \ \mu^+ \nu \ e^+ e^-$) $F_V - F_A = 0.077 \pm 0.026$ (1.47 σ) (1.47 σ)

Expect doubling of the statistics by the end of 2024



$K^+ \rightarrow \pi^0 \mu \nu \gamma (K \mu 3 \gamma)$

This decay complements Ke3γ much better studied by OKA and NA62. OKA publications: JETP Lett. v.116 No 9 (2022), EPJ C(2021) 81. $K^+\mu 3\gamma$ was first seen by ISTRA+ and KEK K470 in 2006 and later by BNL E787 in 2010. For K⁰ was discovered by NA48 in 1998 and later improved by KTeV in 2005

There are calculations of Branching and T-odd asymmetry:

 $\int \frac{2005}{\xi = \vec{p}_{\gamma} \cdot (\vec{p}_{l} \times \vec{p}_{\pi}) / m_{K}^{3}} \qquad A_{\xi} = \frac{N_{\xi > 0} - N_{\xi < 0}}{N_{\xi > 0} + N_{\xi < 0}}$

Q.

 Γ_{19}/Γ

$\Gamma(K^+ ightarrow \pi^0 \mu^+ u_\mu \gamma) / \Gamma_{ m total}$

VALUE (10^{-5})	CL%	EVTS	DOCUMENT ID		TECN	CHG	COMMENT
$1.25\pm0.25\qquad \text{OUR AVERAGE}$							
$1.10 \pm 0.32 \pm 0.05$		23	¹ ADLER	2010	B787		$30 < E_\gamma < 60$ MeV
$1.46 \pm 0.22 \pm 0.32$		153	² TCHIKILEV	2007	ISTR	-	$30 < E_\gamma < 60$ MeV
		• • We do no	ot use the following	data for a	verages, fits, lin	nits, etc. • •	
$2.4 \pm 0.5 \pm 0.6$		125	SHIMIZU	2006	K470	+	$E_\gamma >$ 30 MeV; $\Theta_{\mu\gamma} > 20^\circ$
<6.1	90	0	LIUNG	1973	HLBC	+	$E(\gamma$) $>$ 30 MeV

 1 Value obtained from B($K^{+} \rightarrow \pi^{0} \mu^{+} \nu_{\mu} \gamma$) = (2.51 ±0.74 ±0.12) × 10⁻⁵ obtained in the kinematic region $E_{\gamma} >$ 20 MeV, and then theoretical $K_{\mu3\gamma}$ spectrum has been used. Also B($K^{+} \rightarrow \pi^{0} \mu^{+} \nu_{\mu} \gamma$) = (2.51 ±0.74 ±0.12) × 10⁻⁵ obtained in the kinematic region $E_{\gamma} >$ 20 MeV, and then theoretical $K_{\mu3\gamma}$ spectrum has been used. Also B($K^{+} \rightarrow \pi^{0} \mu^{+} \nu_{\mu} \gamma$) = (2.51 ±0.74 ±0.12) × 10⁻⁵ obtained in the kinematic region $E_{\gamma} >$ 20 MeV, and then theoretical $K_{\mu3\gamma}$ spectrum has been used. Also B($K^{+} \rightarrow \pi^{0} \mu^{+} \nu_{\mu} \gamma$) = (2.51 ±0.74 ±0.12) × 10⁻⁵ obtained in the kinematic region $E_{\gamma} >$ 20 MeV, and then theoretical $K_{\mu3\gamma}$ spectrum has been used. $\pi^0 \mu^+ \nu_\mu \gamma$) = (1.58 ±0.46 ±0.08) ×10⁻⁵, for E_γ > 30 MeV and $\theta_{\mu\gamma}$ > 20°, was determined.

² Obtained from measuring $B(K_{u33}) / B(K_{u3})$ and using PDG 2002 value $B(K_{u3}) = 3.27\%$. $B(K_{u33}) = (8.82 \pm 0.94 \pm 0.86) \times 10^{-5}$ is obtained for 5 MeV < E_{γ} < 30 MeV.

Kl3γ theory, experiment $E_{\gamma}^*>30 MeV \theta_{\mu\gamma}>20^\circ$	Br Kμ3γ x10 ⁵	Br Ke3γ x10 ⁴	A _ξ Kµ3γ QED FSI	A _ξ Ke3γ QED FSI
Bijnens et al. (1993) χPT O(p ⁶)	1.9	2.8		
Braguta et al. (2002) $\chi PT O(p^4)$	2.15	3.18	1.14x10-4	-0.59 x10 ⁻⁴
Khriplovich, Rudenko (2011)	1.81 ± 0.2	2.72 ± 0.1	2.38 x 10 ⁻⁴	-0.30 x 10 ⁻⁴
NA62		3.087 ± 0.037		< 8.6 x 10 ⁻³
OKA	2.0±0.1	2.98± 0.094	< 0.1	< 13 x 10 ⁻³

From Braguta et al. (2003) for NP $K\mu 3\gamma$:

$$A_{\xi} = -(3.6 \cdot 10^{-3} Im(g_s) + 1.2 \cdot 10^{-2} Im(g_p) + 1.0 \cdot 10^{-2} Im(g_v + g_a))$$

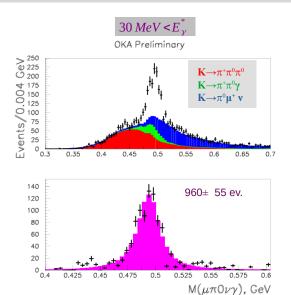
Sensitivity of Ke3 γ to NP is not as good :

 $A_{\xi} = -(2.9 \cdot 10^{-6} \,\text{Im}(g_s) + 3.7 \cdot 10^{-5} \,\text{Im}(g_p) + 3.0 \cdot 10^{-3} \,\text{Im}(g_v + g_a))$

Event selection

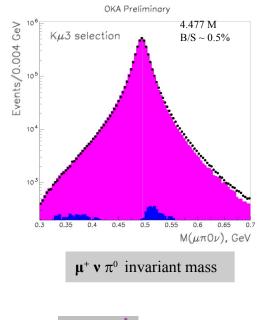
- 1 beam K⁺ track 1 secondary track identified as μ in GAMS, GDA-100 and μ C
- Decay vertex inside DV
- 3 e.m. shower in GAMS with E > 0.6GeV not ass. with track
- π^0 identification $|m_{\gamma\gamma} m_{\pi 0}| < 15$ MeV (best combination)
- $E_{miss} > 0.5 \text{ GeV}$
- The position of radiative photon at GAMS surface is Q. not near beam hole nor at the boudary
- Q. $E_{GS} < 10 \text{ MeV}$; $E_{EGS} < 100 \text{ MeV}$
- Number of additional track segments after spectrometer Q. magnet is zero

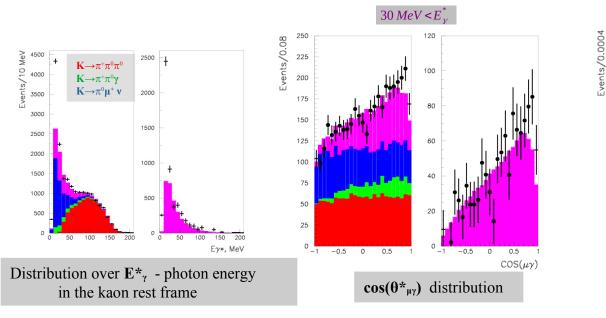
Given Miss-mass $(P_K - P_{\pi^+} - P_{\pi^0})^2 < 0.014 \text{ GeV}^2$ (against $K \rightarrow \pi^+ \pi^0 \pi^0$ bkg)

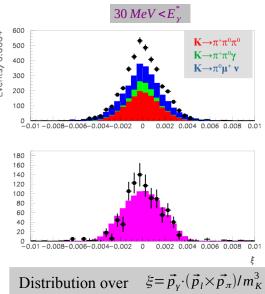


 $\mathbf{K} \rightarrow \pi \mu \nu \gamma u c a y study$ (110mmai

Branching : The decay $\mathbf{K} \rightarrow \mu^+ \mathbf{v} \pi^0$ is used for the normalisation $\mathbf{Br}(\mathbf{K}\mu\mathbf{3}\gamma)/\mathbf{Br}(\mathbf{K}\mu\mathbf{3}) = (4.5 \pm 0.25 \text{ (stat)}) \cdot 10^4, 30 \text{ MeV} < E_{\gamma}^* < 60 \text{ MeV}$ Using PDG value: $\mathrm{Br}(\mathbf{K}\mu\mathbf{3}) = 3.352\%$: $\mathbf{Br}(\mathbf{K}\mu\mathbf{3}\gamma) = (1.49 \pm 0.085 \text{ (stat)}) \cdot 10^5, 30 \text{ MeV} < E_{\gamma}^* < 60 \text{ MeV}$ in agreement with ISTRA+ measurement, but statistical error is 3 times smaller. For the comparison with theory : $\mathbf{Br}(\mathbf{K}\mu\mathbf{3}\gamma) = (2.0 \pm 0.1 \text{ (stat)}) \cdot 10^5, E_{\gamma}^* > 30 \text{ MeV}, \theta_{\mu\gamma} > 20^\circ$ Bijnens et al. χ PT O(p⁶) 1.9 x 10⁵, Braguta et al. χ PT O(p⁴) 2.15 x 10⁵, Khriplovich et al. 1.8 x 10⁵ For the T-odd asymmetry the result is: $\mathbf{A}_{\xi} = -0.006 \pm 0.069$ ($\mathbf{A}_{\xi} < 0.1 90\%$ C.L.)





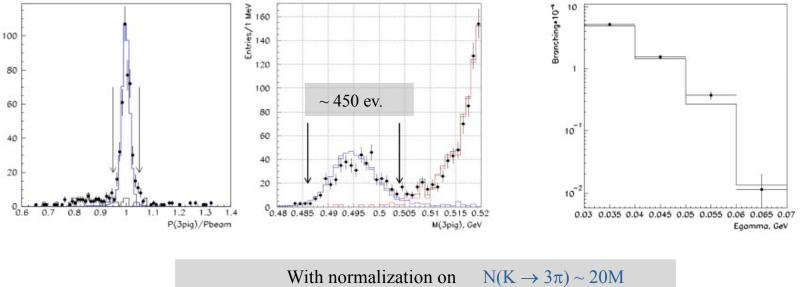




Study of the decays $K^+ \rightarrow \pi \pi \pi \gamma$

G. D'Ambrosio, G. Ecker, G. Isidori, H. Neufeld "The present experimental status of K3pig decays is rather meager" χ PT O(p⁴) "generalized bremsstrahlung"

K⁺ → $\pi^+\pi^-\gamma$ was first found in ITEP Xenon BC В.В. Бармин и др., ЯФ 50(1989)679-682 7 ev. E^{*}_γ ~ 5-10 MeV, Br. measured



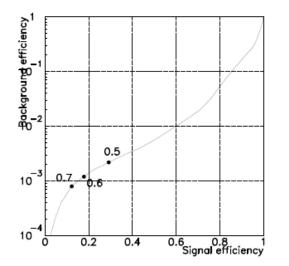
 $Br(K^+ \to \pi^+ \pi^- \gamma) = (7.1 \pm 0.4_{stat} \pm 0.3_{syst}) \cdot 10^{-6} \quad E^*_{\gamma} > 30 \text{ MeV}$ $\chi PT O(p^4) \qquad 6.65 \cdot 10^{-6}$

Study of this decay is being continued on ~20 times higher statistics of NA62

$K^+ \rightarrow \pi^+ \pi^0 \pi^0 \gamma$ decay Eur.Phys.J.C 84 (2024) 4, 345

Complementary to $K^+ \rightarrow \pi^+ \pi^- \gamma$, less charged particles in the final state

 K^- → $\pi^-\pi^0\pi^0\gamma$ searched for at ISTRA В.Н. Болотов и др., Письма в ЖЭТФ, 1985 т.42, вып.9, с.390-392 5 ev. E^*_{γ} >10 МэВ, Br. estimated



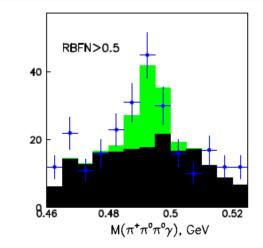
NN performance : with RFBN>0.6 the background is suppressed by ~1000 with a signal efficiency of 0.2

 53.8 ± 13 events in the peak. p-value of the no-signal hypothesis is $9 \cdot 10^{-5}$ Events selection (started from 3.6.109 events)

- 1 beam K⁺ track
- 1 secondary, θ >2 mrad, # (µC), = (E/p GAMS)
- Decay vertex inside DV, CDA < 1cm.</p>
- 5 e.m. showers in GAMS E > 0.5 GeV not ass. with a track Selection of $\pi^0 \pi^0$: $(m_{\gamma 1 \gamma 2} - m_{\pi 0})^2 + (m_{\gamma 3 \gamma 4} - m_{\pi 0})^2$ (best combination) **230K events, main background - K** $\rightarrow \pi^+ \pi^0 \pi^0$ (x 5000) Then NN (Radial Basis Function Network RFBN) is used

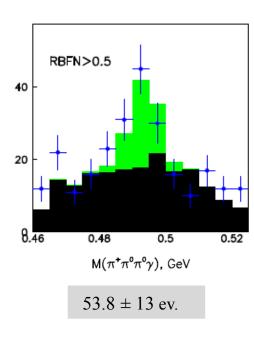
Input parameters for NN:

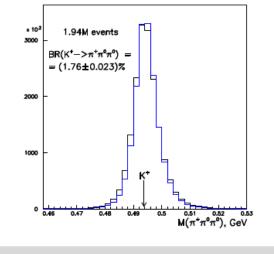
- $\Delta E = E_{\pi+} + \Sigma E_{\gamma i} E_{beam}$
- $E_{\gamma 5}$ energy of free (5-th) gamma
- $d_{\gamma 5}$ distance from 5-th gamma to the track on the GAMS plane
- χ^2 of the shower shape fit for the 5-th gamma χ^2 of the K⁺ $\rightarrow \pi^+ \pi^0 \pi^0$ 3C- fit
 - χ^2 of the K⁺ $\rightarrow \pi^+ \pi^0 \pi^0 \gamma$ 3C- fit
- $\mathbf{M}(\pi^{+}\pi^{0}\pi^{0})$



° and

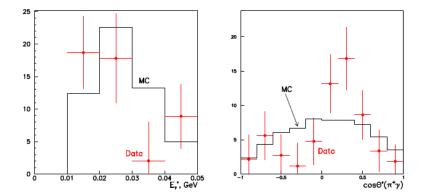
Распад $K^+ \to \pi^+ \pi^0 \pi^0 \gamma$ Eur.Phys.J.C 84 (2024) 4, 345





With normalization on $N(K^+ \rightarrow \pi^+ \pi^0 \pi^0) \sim 2M$ Br $(K^+ \rightarrow \pi^+ \pi^0 \pi^0 \gamma) = (3.7 \pm 0.9_{stat} \pm 0.3_{syst}) \cdot 10^{-6} E^*_{\gamma} > 10 \text{ MeV}$ Br $(\chi \text{PT O}(p^4)) = 3.76 \cdot 10^{-6}$

The differential spectra over $E^*_{\ \gamma}$ and $cos\theta_{\pi\gamma}$ in the K⁺ rest frame are also obtained



The following results have been obtained by the OKA collaboration:

✓ Search for the ALP in the decay $\mathbf{K}^+ \rightarrow \pi^+ \pi^0 \mathbf{a}$ is performed. No signal found, 90% C.L upper limit $\mathbf{Br} < 2.5 \cdot 10^{-6} \div 2 \cdot 10^{-7}$ for the ALP mass from 0 to 200 MeV. A lower limit for the F_{sd}^A - coupling constant of the axion to the axial sd current is $F_{sd}^A > 6.5 \cdot 10^7 \,\text{GeV}$ for the ALP mass below 70 MeV.

✓ The radiative decay $\mathbf{K}^+ \rightarrow \boldsymbol{\mu}^+ \mathbf{v} \gamma$ is studied on statistics of ~144K events for 25 MeV < \mathbf{E}^*_{γ} < 150 MeV. A destructive interference between IB and SD- is clearly seen. The difference of the vector and axial vector constants Fv-Fa is measured:

 $F_{V}-F_{A} = 0.135 \pm 0.017_{stat} \pm 0.024_{syst}$

which is 2.8 σ from χPT $O(p^4)$ and 1.8 σ from Lattice .

✓ The decay $\mathbf{K}^+ \to \boldsymbol{\mu}^+ \mathbf{v} \pi^0 \gamma$ is studied on statistics of ~1K events for $E_{\gamma} > 30$ MeV region. Branching fraction is measured: $\mathbf{Br}(\mathbf{K}\boldsymbol{\mu}\mathbf{3}\gamma) = (\mathbf{1.98} \pm \mathbf{0.1} \text{ (stat)}) \cdot \mathbf{10}^{-5}$

To be compared with $\chi PT O(p^4) = 2.15 \cdot 10^{-5}$; $\chi PT O(p^6) = 1.9 \cdot 10^{-5}$ An upper limit for the CP-odd asymmetry is obtained:

 $A_{\xi} = -0.006 \pm 0.069$ ($A_{\xi} < 0.1$ 90% C.L.)

✓ The decay K⁺ → π⁺π⁰π⁰ γ is seen for the first time ~50 events with E_γ > 10 MeV. Branching fraction is measured:

Br(K⁺ $\rightarrow \pi^{+}\pi^{0}\pi^{0}\gamma) = (3.7 \pm 0.9_{stat} \pm 0.3_{syst}) \cdot 10^{-6} \text{ E}_{\gamma}^{*} > 10 \text{ MeV}$

To be compared with $Br(\chi PT O(p^4)) = 3.76 \cdot 10^{-6}$

✓ The decay $K^+ \rightarrow e^+ v \pi^0 \pi^0 \pi^0$ is searched for . No events found. $Br(K^+ \rightarrow \pi^+ \pi^0 \pi^0 \gamma) < 5.4 \cdot 10^{-8}$ (See talk of A. Kulik on Thursday morning session)