

Contribution ID: 22 Type: Oral talk

Alpha-particle states in relativistic nuclear fragmentation

Wednesday, 23 October 2024 09:30 (15)

Ensembles of He and H isotopes can be studied with unique completeness and resolution in nuclear emulsion layers longitudinally exposed to relativistic nuclei [1,2]. Determination of the invariant mass of their pairs or triplets by emission angles in the velocity conservation approximation is sufficient to identify a number of unstable states $-{}^{8}Be(0^{+})$, ${}^{8}Be(2^{+})$, ${}^{9}B$, ${}^{12}C(0_{2}^{+})$, ${}^{12}C(3^{-})$, ${}^{6}Be$.

The BECQUEREL experiment [3,4], using this approach, is aimed at searching for the α -particle Bose-Einstein condensate (α BEC), an unstable of S-wave α -particle state. $^8Be(0^+)$ is associated with 2α BEC, and $^{12}C(0_2^+)$ or the Hoyle state with 3α BEC. In the relativistic fragmentation of heavy nuclei, an enhancement of 8 Be, 9 B and $^{12}C(0_2^+)$ is detected, suggesting their synthesis in the fusion of associated α -particles. The focus of the search is the 4α BEC state of $^{16}O(0_6^+)$ at 660 keV above the 4α threshold, decaying into $\alpha^{12}C(0_2^+)$ or 2^8 Be. In this context, the status of the analysis of α -particle fragmentation in a nuclear emulsion exposed to 84 Kr nuclei at 950 MeV per nucleon is presented. Secondary stars produced by relativistic neutrons are observed in the nucleus fragmentation cone [4]. The neutron average energy in the parent nucleus system is estimated to be 1.3 MeV [5].

The α BEC search leads to the study of nuclear matter in the region of temperature and density from red giants to supernova. It is characterized by the ratios of 1,2,3 H and 3,4 He. Nuclear emulsion layers exposed to heavy nuclei of several GeV per nucleon at the NICA accelerator complex are optimal for identifying H and He isotopes by multiple scattering, searching for unstable states, and assessing neutron accompaniment. An exposure to 124 Xe nuclei of 3.8 GeV per nucleon, performed at the NICA/Nuclotron accelerator complex, allows the use of proven approaches. Parameters of the beam are determined using the CR-39 track detector by direct crater counting on the Olympus BX63 microscope.

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Track Classification: Nuclear physics