

FACILITIES AND ADVANCED DETECTOR TECHNOLOGIES  
SECTION

**PMT/WLS plate optical modules  
for Cherenkov detectors**

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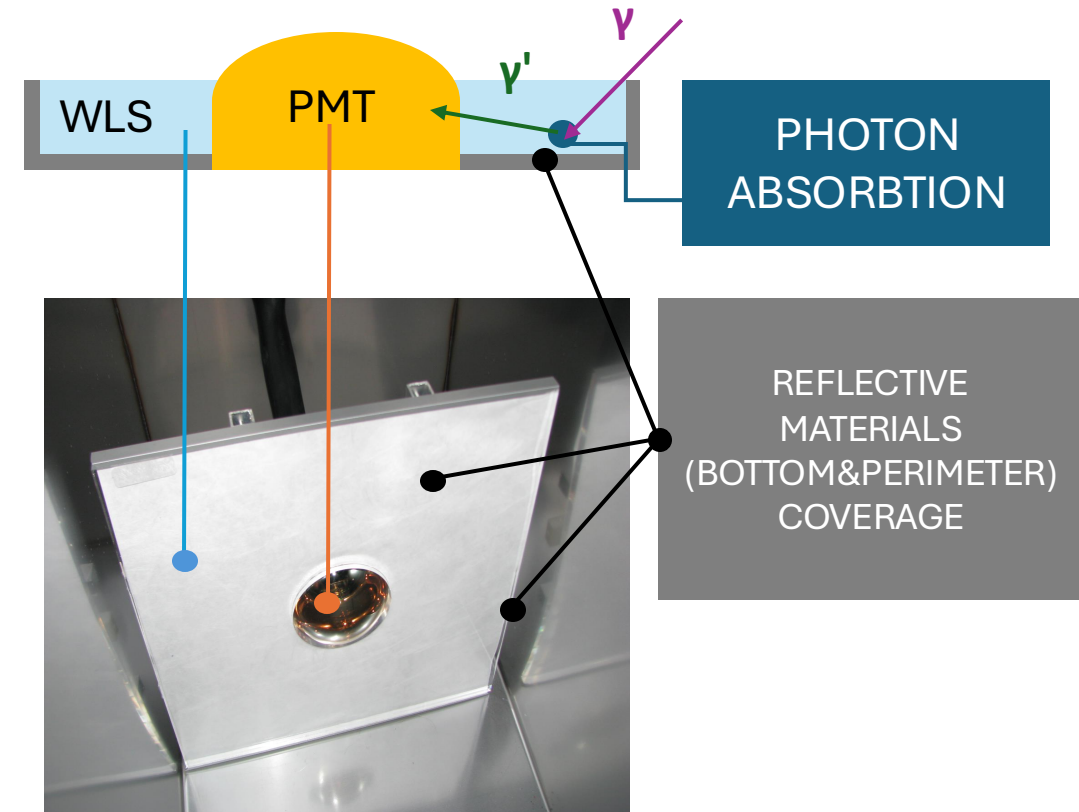


# WLS/PMT OPTICAL MODULES

WLS plate (wavelength shift) with PMT (photomultiplier tube) optical module for Cherenkov detectors: WLS plate **increases useful surface** of detector **for light collection** and **absorbs** Cherenkov's light then **reemits it in sensitivity** part of **PMT spectrum**.

WLS plates in report are made from **PMMA** (polymethylmethacrylate, as known as acrylic glass) with WLS dopants in Dzerzhinsk at "SRI Polymers".

Such optical modules can be used in water Cherenkov detectors like a component of veto-systems. For example, **Outer Detector (OD)** of **Hyper Kamiokande (HK)**. **Design** of WLS plates in report is **developed for HyperK OD**. Plate size is  $300 \times 300 \times 7 \text{ mm}^3$ .



Pic.1. Schematic illustration of the operating principle of the WLS/PMT optical module. Reflective materials are applied in the form of a film around the perimeter and a substrate at the bottom.

# PMT STUDY



Pic.2. Device for cathode sensitivity measuring.

Cathode sensitivity is important characteristic for PMT and WLS plate **optical connection**. There was comparison of two main **3" PMT candidates** for usage in **OD: Hamamatsu R14374** and **NNVT 2031**.

**Both** of them have **identical specification**. The **sensitivity area** of photocathode corresponds for cathode diameter **less than 72  $\varnothing$ mm** due to specification.

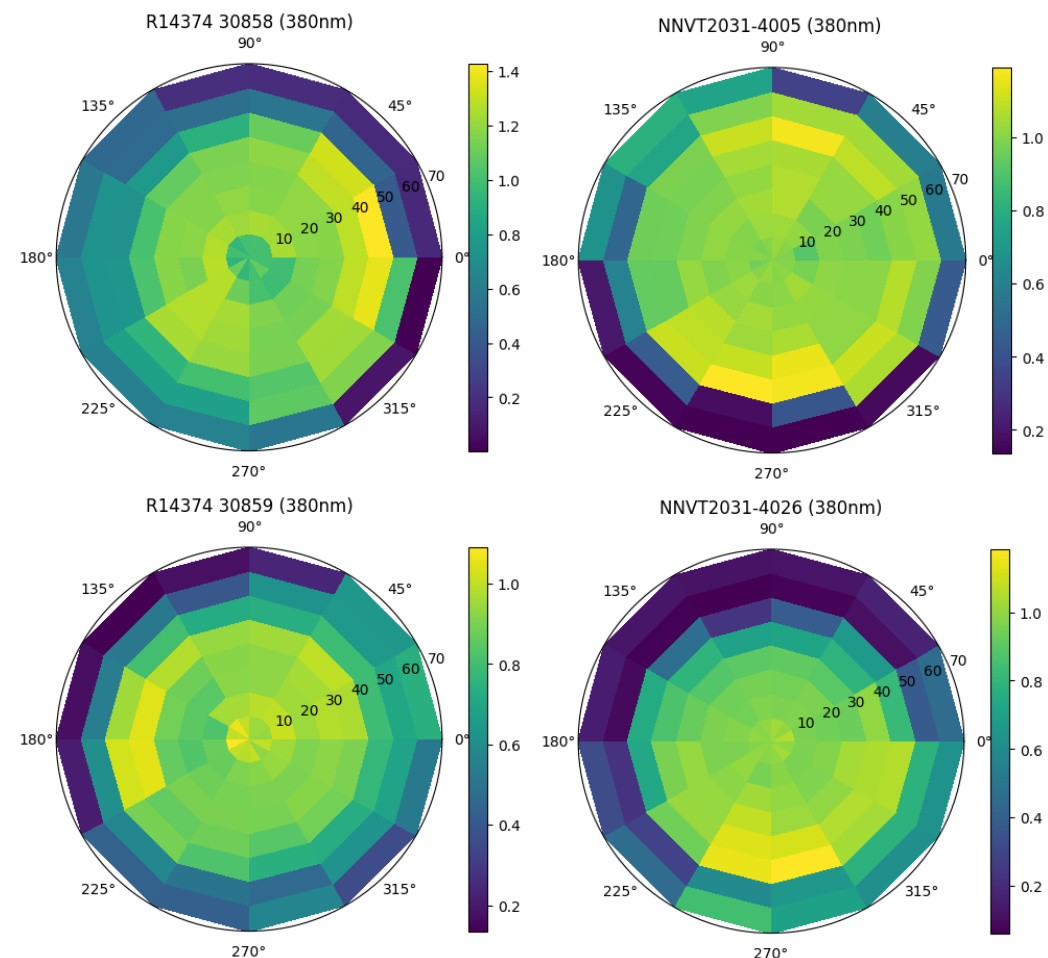
Special device was developed for measurements of the cathode sensitivity. It can provide measurements with azimuthal and polar coordinates variation.

# CATHODE SENSITIVITY PRELIMINARY RESULTS

Varying the polar and azimuthal angles made it possible to construct heatmaps of the tested photomultipliers. The sensitivity of the cathode was estimated based on heatmaps.

Result of the test for LED 380nm:

- All PMTs have **weaker signal** in "non-sensitivity" area
- **R14374** has **more stable** cathode sensitivity, especially in **contact area** with WLS

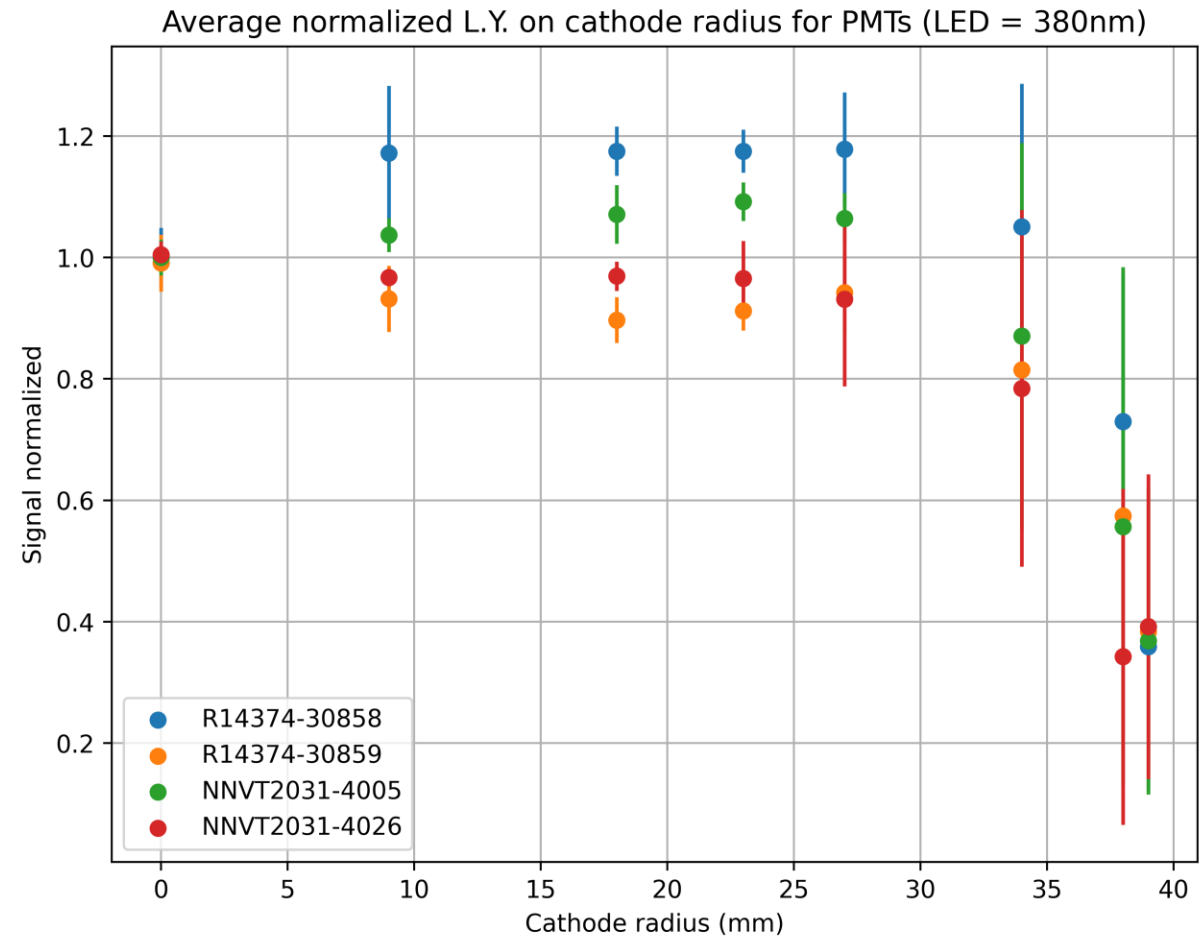


Pic.3. Cathode sensitivity heatmaps:  
leftside – R14374; rightside – NNV2031.

# CATHODE SENSITIVITY PRELIMINARY RESULTS

Results for **the signal averaged over the cathode radius** show that R14374 **does not differ** from NNVT 2031 **in the sensitivity area** (less than 36 mm of the cathode radius) and is quite stable.

The **non-sensitivity** area starts from a cathode radius of **34 mm**, and the reduction level is about **60%**.



Pic.4. Average signal for cathode radius.

# DARK RATE

## PRELIMINARY RESULTS

Dark rate measurement for Hamamatsu R14374 with **Threshold 0.3 p.e** and temperature **20 °C** for fixed **gain level =  $5 \times 10^6$** .

| Nº      | SUPPLY VOLTAGE, V | GAIN X 10 <sup>6</sup> | DARK RATE, Hz |
|---------|-------------------|------------------------|---------------|
| KM32077 | 1110              | 5.0                    | 290           |
| KM30858 | 1095              | 5.0                    | 291           |
| KM60330 | 1075              | 5.0                    | 245           |
| KM60344 | 1140              | 5.0                    | 286           |
| KM30863 | 1320              | 5.0                    | 435           |
| KM30859 | 1225              | 5.0                    | 478           |
| KM60329 | 1135              | 5.0                    | 323           |

# WLS PLATES STUDY

## WLS DOPANT REVIEW

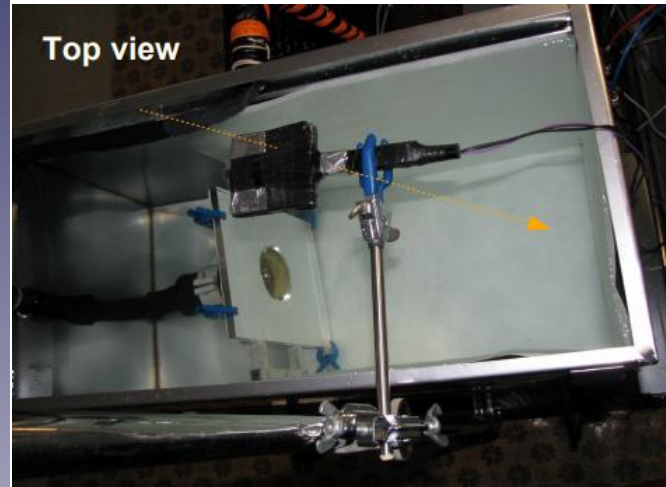
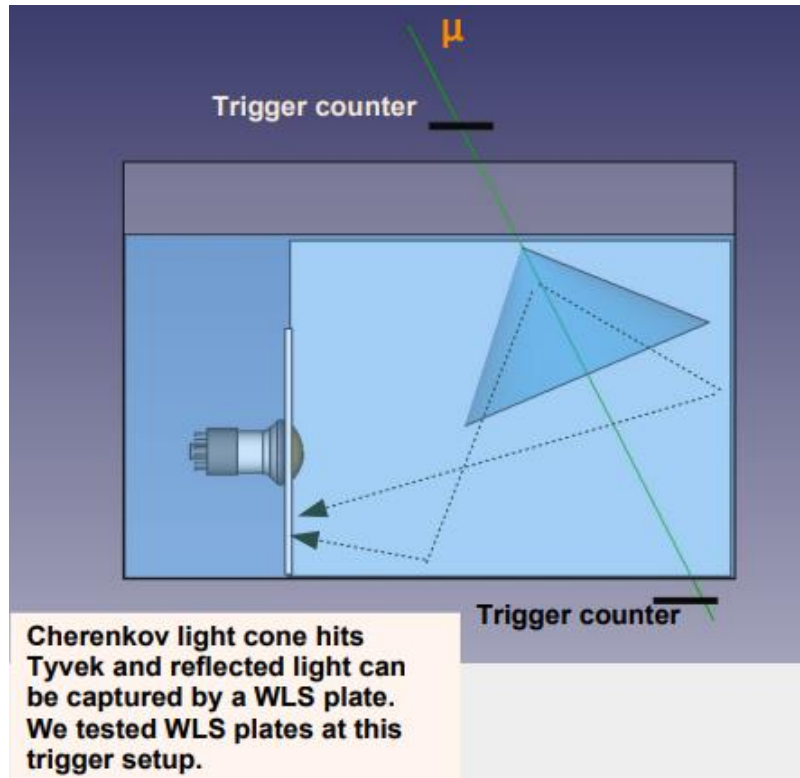
**POPOP** for long UV, **PPO** for short UV

**Marker rule: first place – dopant name**  
**Then – concentration [mg/L]**

**Task:**  
Check all combination of dopants and it's concentration to find most optimal variant.

| WLS                     | POPOP                        | PPO                          |
|-------------------------|------------------------------|------------------------------|
| Absorption spectrum, nm | 250 - 390<br>peak: 360       | 290 - 330<br>peak 300        |
| Emission spectrum, nm   | 380 - 510<br>peak: 390 - 450 | 320 - 410<br>peak: 340 - 380 |

# WLS DOPANT SELECTION



**Infant-K** is special **water Cherenkov detector** that was created for **WLS plates study**.

Plate size: 7 x 300 x 300 mm<sup>3</sup>

Double fluor POPOP50-PPO3000 - 15 pcs

Single fluor POPOP50 - 7 pcs

Single fluor POPOP800 - 3 pcs

Light source:

reflected Cherenkov light from cosmic muons

Reflector:

unknown Tyvek, 190  $\mu\text{m}$  thickness

Water: distilled for industry purposes

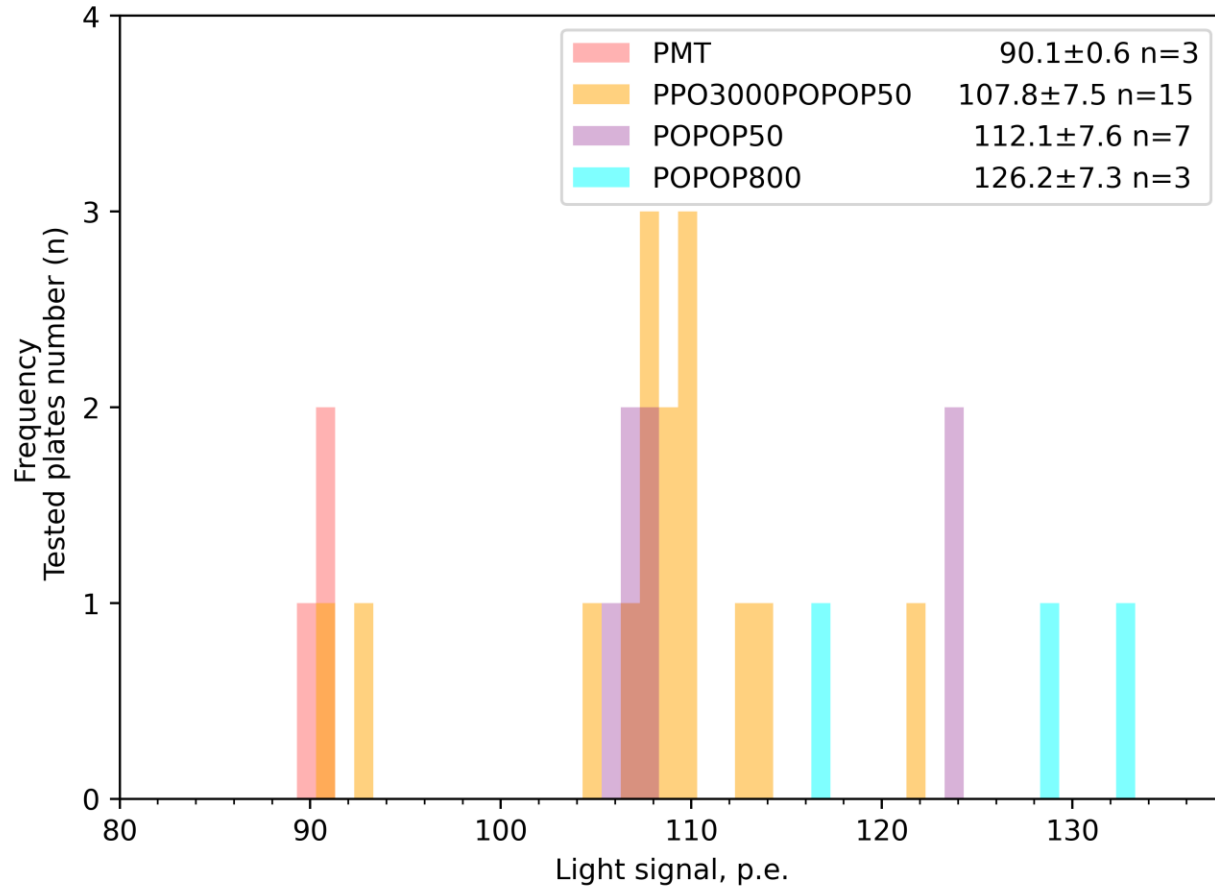
PMT : Hamamatsu R14374

Pic.5. Schematic view of water Cherenkov detector Infant-K.  
Trigger telescope is arranged to select cosmic muon tracks inclined in direction away from a WLS plate.

Measurement cycle has taken over 2 months.



# WLS DOPANT SELECTION



Pic.6. Results of Infant-K tests.

Average light yield from a WLS plate  
(PMT contribution is subtracted):

- POPOP50-PPO3000 : 108 p.e. (15 pcs)
- POPOP50 : 112 p.e. (7 pcs)
- POPOP800 : 126 p.e. (3 pcs)

Spread in light yields:  $\pm 12\%$  from average.

# WLS PLATES DARK RATE

Dark rate is important for triggering systems. Low dark rate provides HK with longer exposition time. In this test dark rate was checked with **temperature 13 °C** like in OD:

| WLS plates                    | Bare PMT | Pure PMMA (no WLS fluor) | POPOP 800 | POPOP 400 | POPOP 200 | POPOP 50 | POPOP 50/PPO3000 |
|-------------------------------|----------|--------------------------|-----------|-----------|-----------|----------|------------------|
| Dark rate at 0.25-0.3 p.e. Hz | ~ 200    | ~ 240                    | ~ 800     | ~ 490     | ~ 390     | ~ 300    | ~ 680            |

**POPOP 50** is the best choice due to **low dark rate** level and **high signal** for Cherenkov's light.

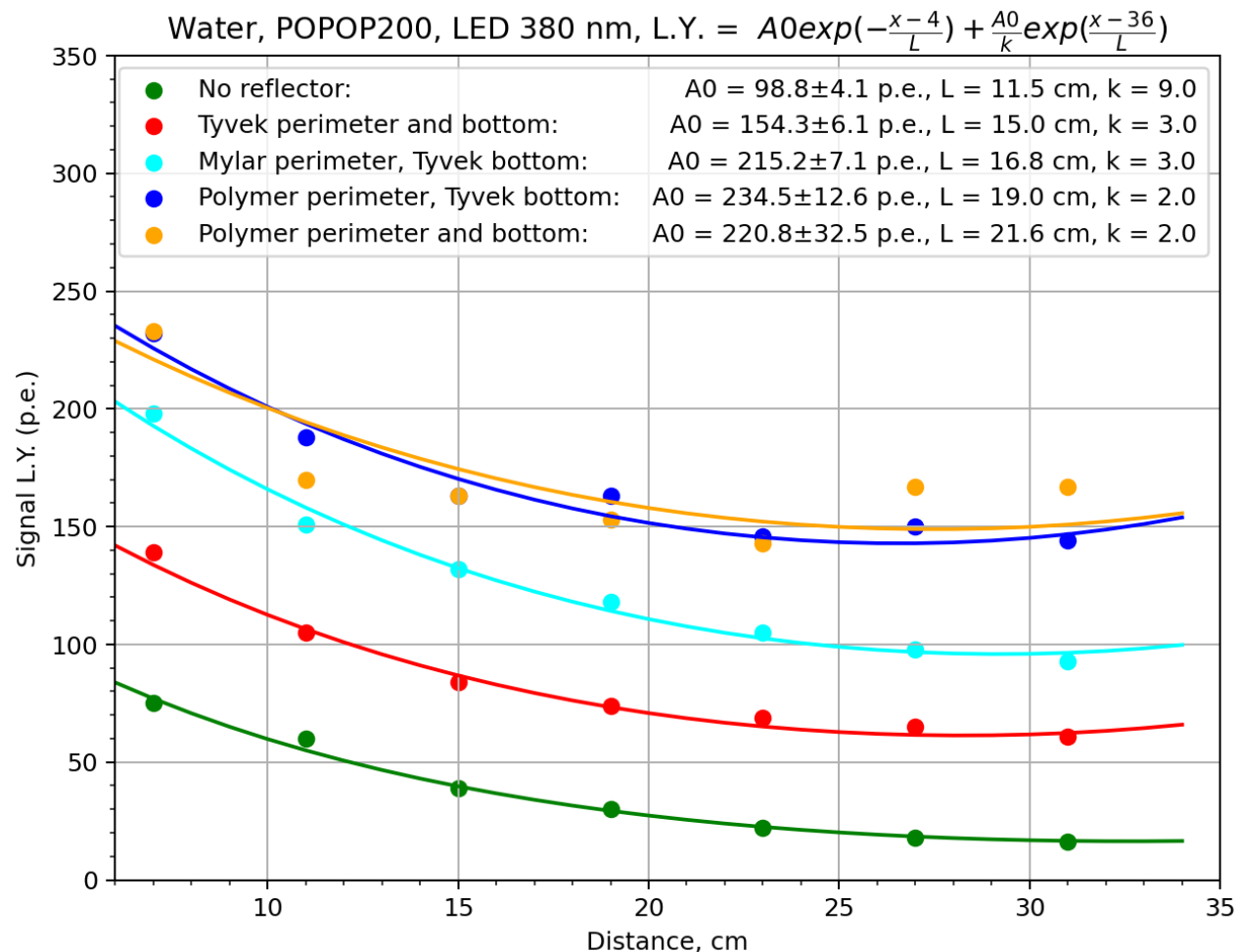
# REFLECTORS STUDY

Usage of a **reflector** is a good idea to **increase light collection** of WLS plate. Multireflected light has higher possibility to transfer into sensitivity spectrum of PMT.

There are three types of reflectors to choose:

- Tyvek (polymer, diffusion reflector)
- Mylar (aluminized film, mirror reflector)
- 3D DF2000MA (polymer, mirror reflector)

The most effective combination of reflectors is Polymer Perimeter 3D DF2000MA and Tyvek bottom (**increase of L.Y. by 2 times compared to plate without reflectors**)



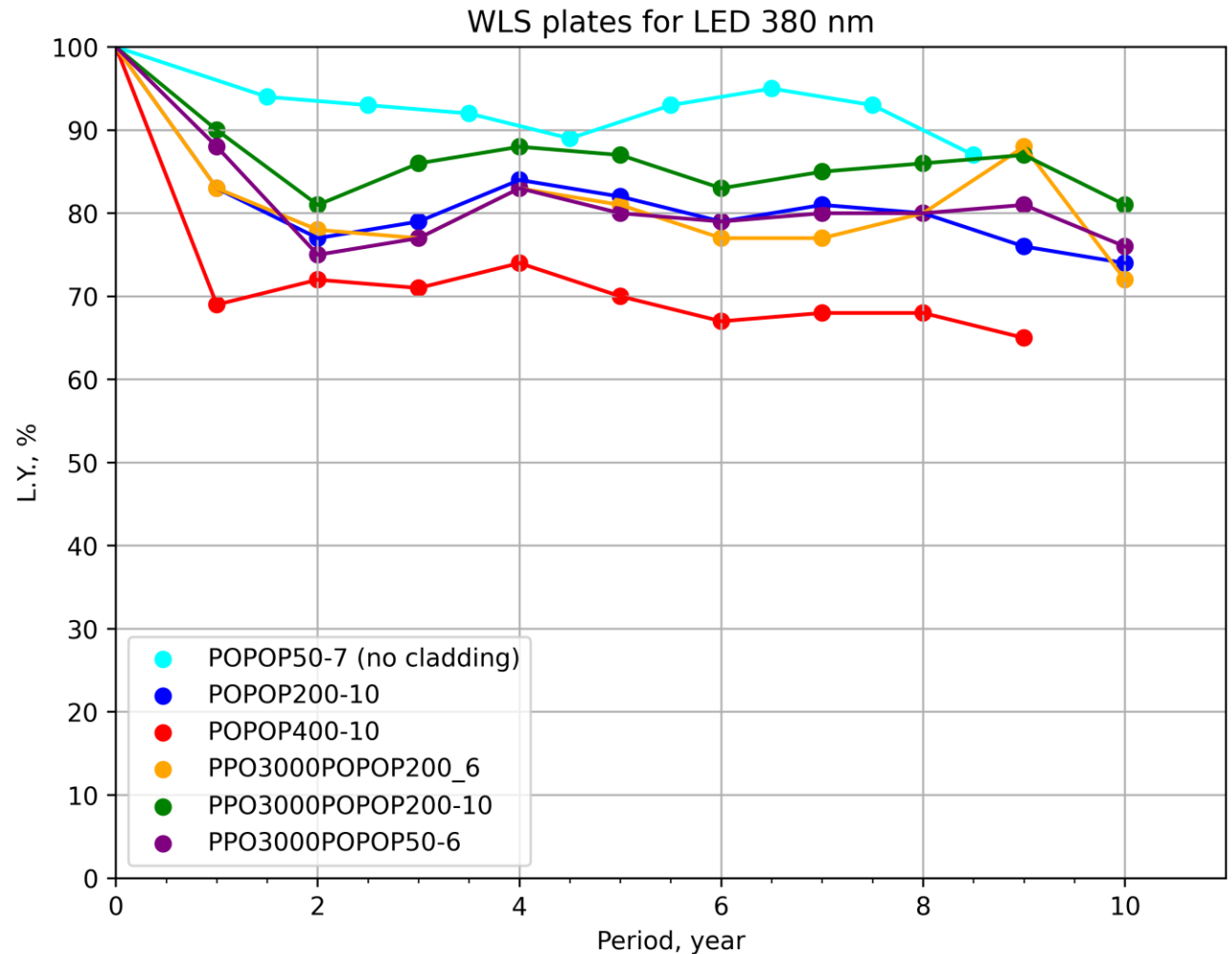
Pic.7. Comparison of reflectors in water.

# WLS PLATES AGING

Aging of WLS Plates is an important aspect to measure. All plates were tested in special thermal camera with constantly temperature **59 °C**. **2 weeks** in camera is equal **one year** at **13 °C** in **OD**.

Cladding - 3D DF2000MA perimeter. All signals was calibrated on plate signal before aging test.

All plates are practically **stable after first year of heating**. The possible reason of first year aging is **glue** in cladding. Excluding this fact the **aging** is less than **< 2%/year**.



Pic.8. Aging of WLS plates.



# CONCLUSION

The results of PMT/WLS plates optical modules for Cherenkov detectors study:

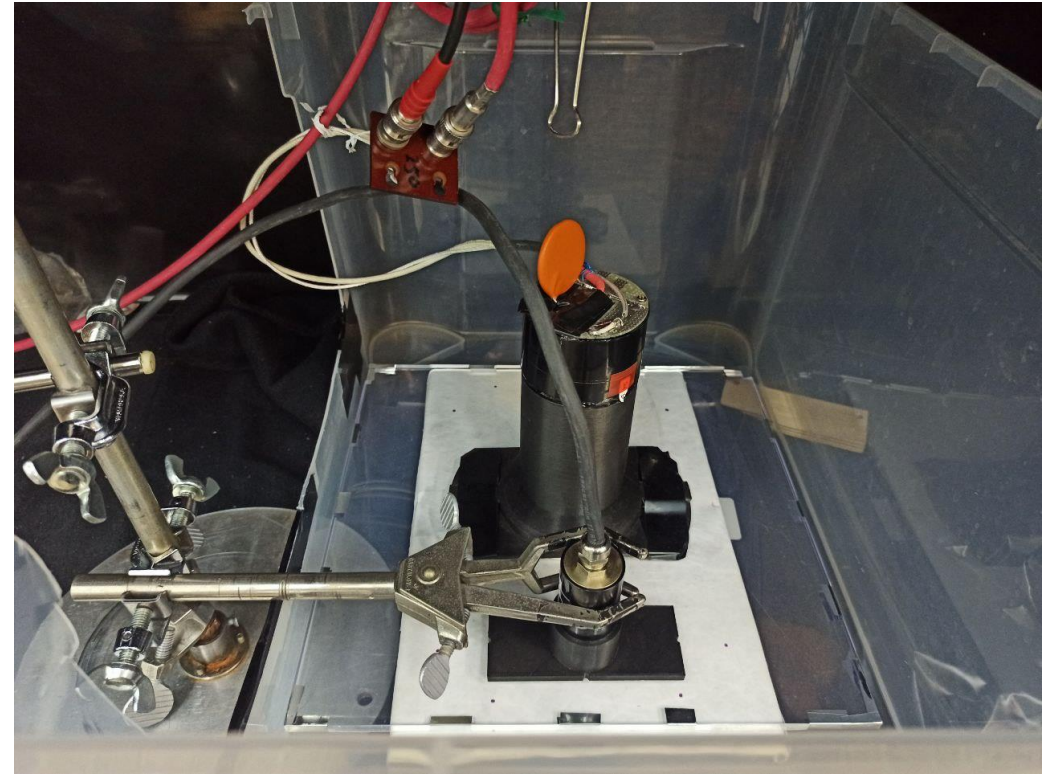
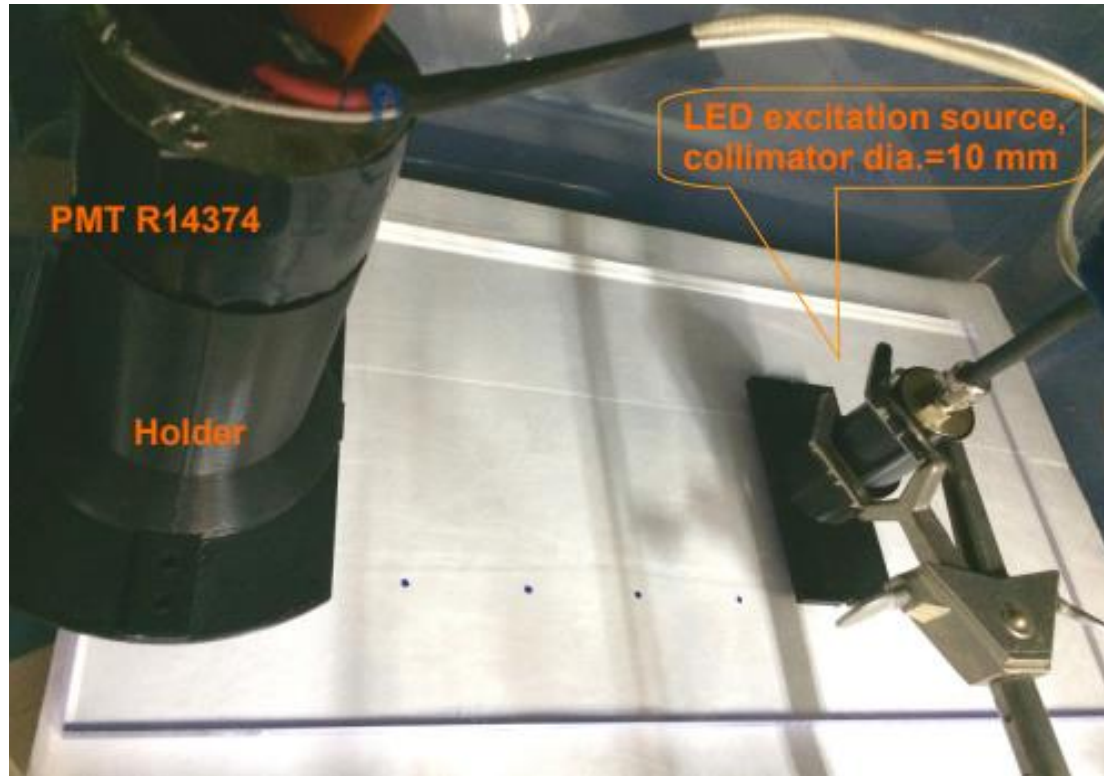
1. **Hamamatsu R14374** has **better cathode surface** than **NNVT 2031**. **Non-sensitivity** area **starts** from **34 mm** cathode radius. The **signal decrease level** in **non-sensitivity** area is nearby **60%** for both PMTs.
2. **Dark rate** for most Hamamatsu **R14374** with threshold **0.3 p.e** and temperature **20 0C** is **less** than **~320 Hz**. Some PMTs have bigger dark rate.
3. **POPOP 50** is the best choice due to **low dark rate** level and **high light signal** for Cherenkov's light.
4. **Reflectors** can **improve** WLS plates **light collection** efficiency **by 2 times**. The **best choice** is **3D DF2000MA** film on **perimeter** and **Tyvek** reflector on **bottom side** of WLS plate.
5. **WLS plates aging** is less than **< 2%/year** excluding glue factor.



**Thank you for your attention!**

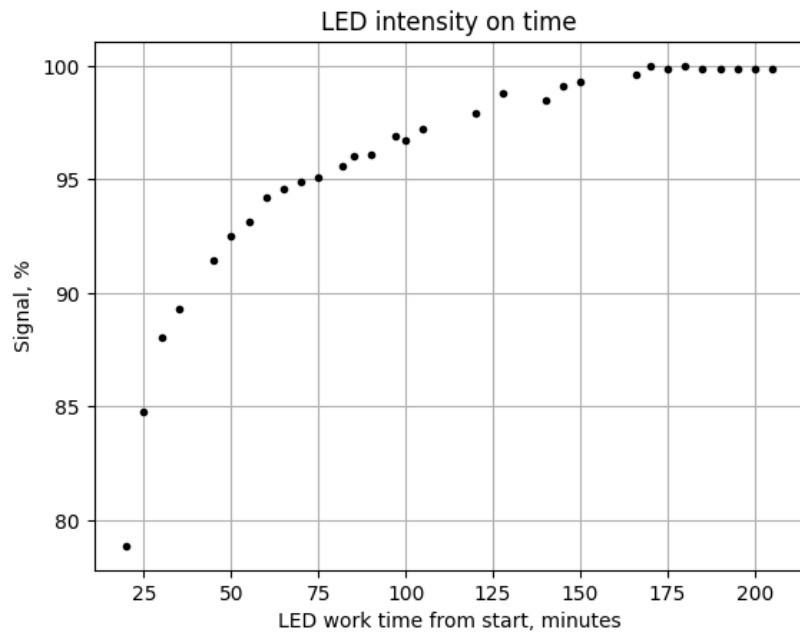


# BACK UP SLIDES

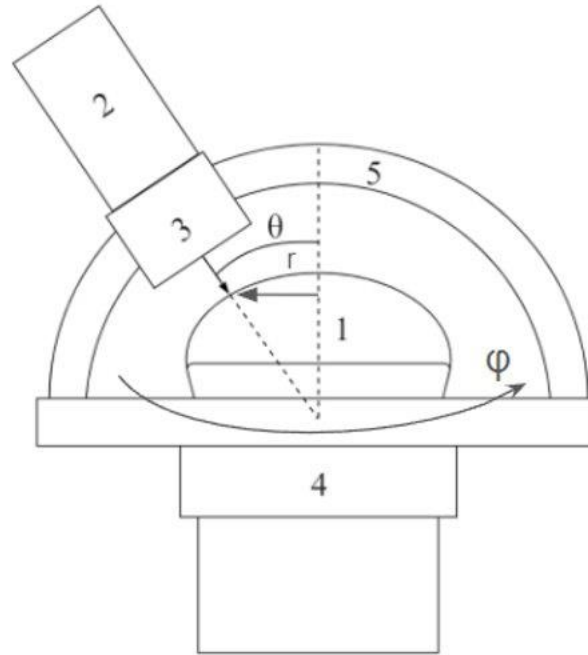


Pic.9. Setup for LED tests of WLS plates. Measuring of L.Y. on distance of LED position.

# PMT STUDY



Pic.10. LED stability on time.



1. PMT
2. LED
3. Collimator 1  $\varnothing$ mm
4. Graduated azimuthal ring and rotation-platform for azimuthal angle  $\varphi$  variation
5. Arcs for polar angle  $\theta$  variation

Pic.11. Schematic illustration of device for cathode sensitivity study.



# HAMAMATSU R14374 SPECIFICATION

## ●R14374

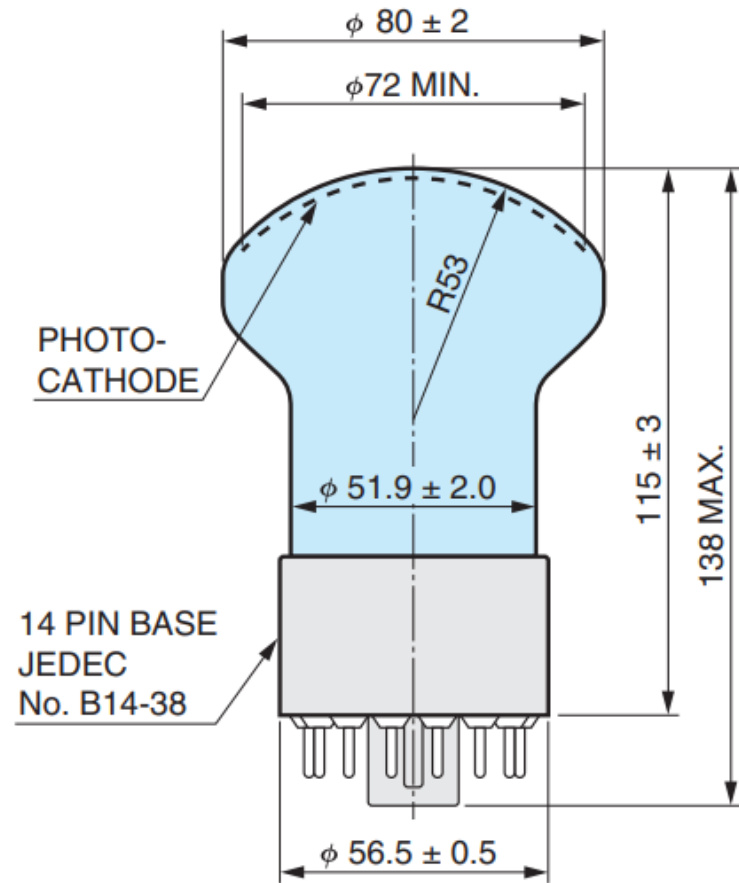


Figure 1: Typical spectral response

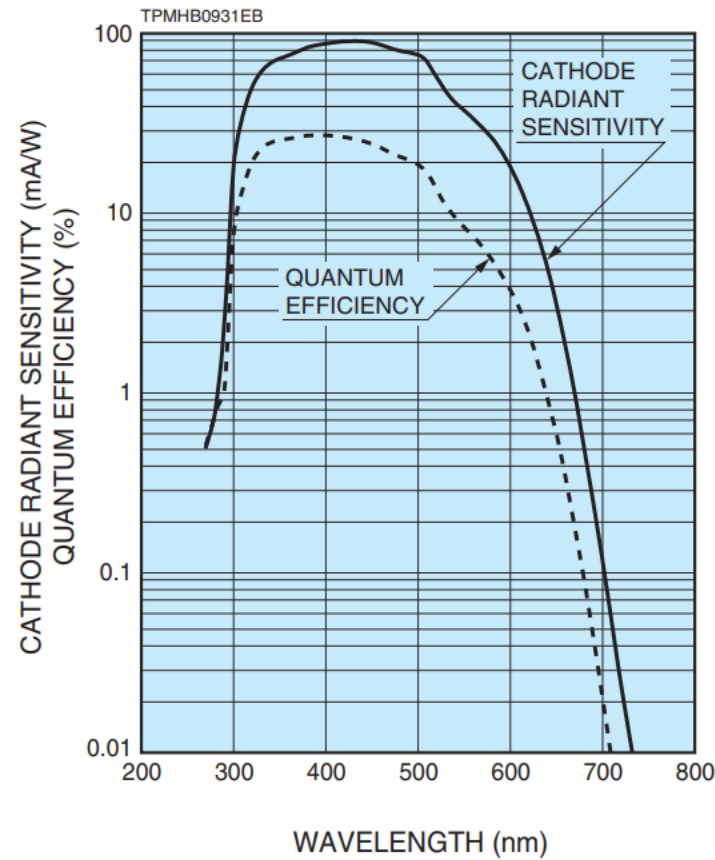
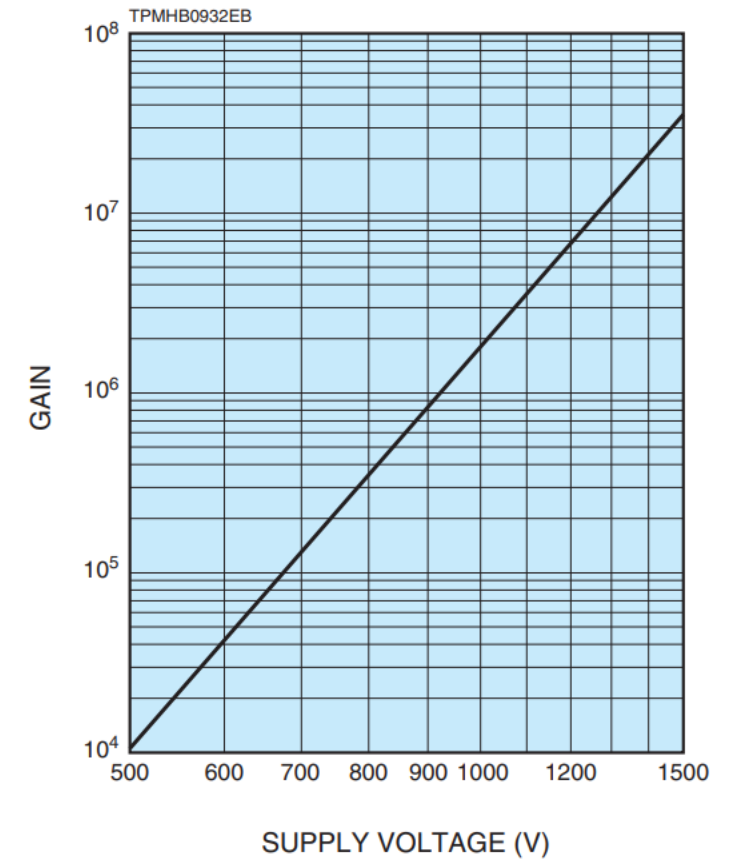
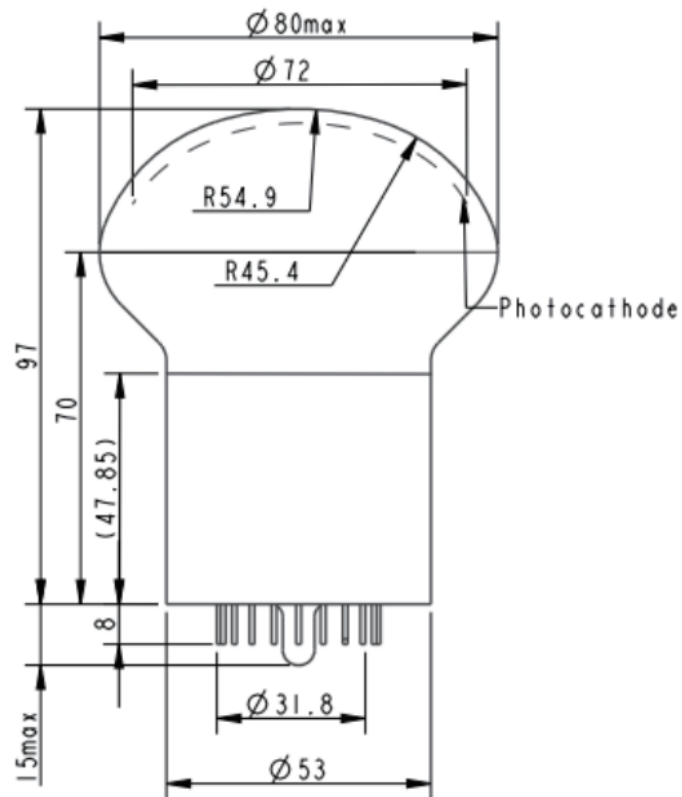


Figure 2: Typical gain

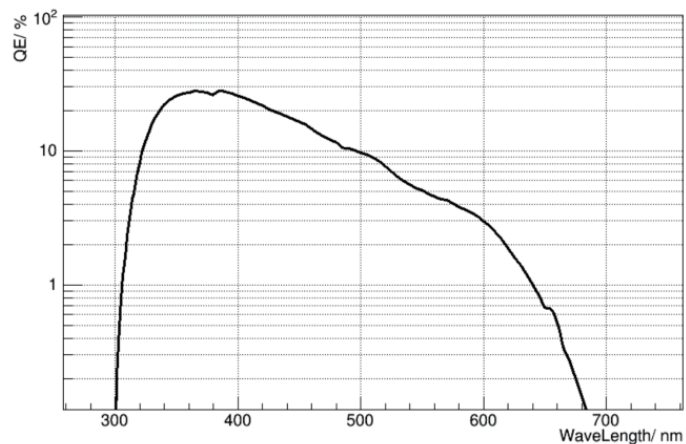


Pic.12. Hamamatsu R14374 specification

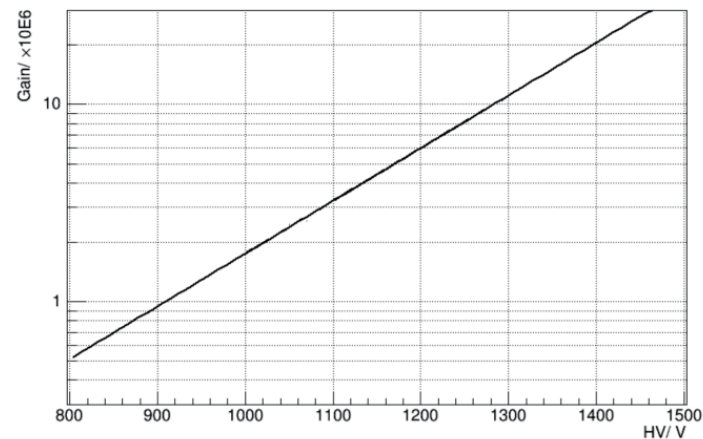
# NNVT 2031 SPECIFICATION



**N2031光电倍增管结构图**  
**N2031 PMT structure**



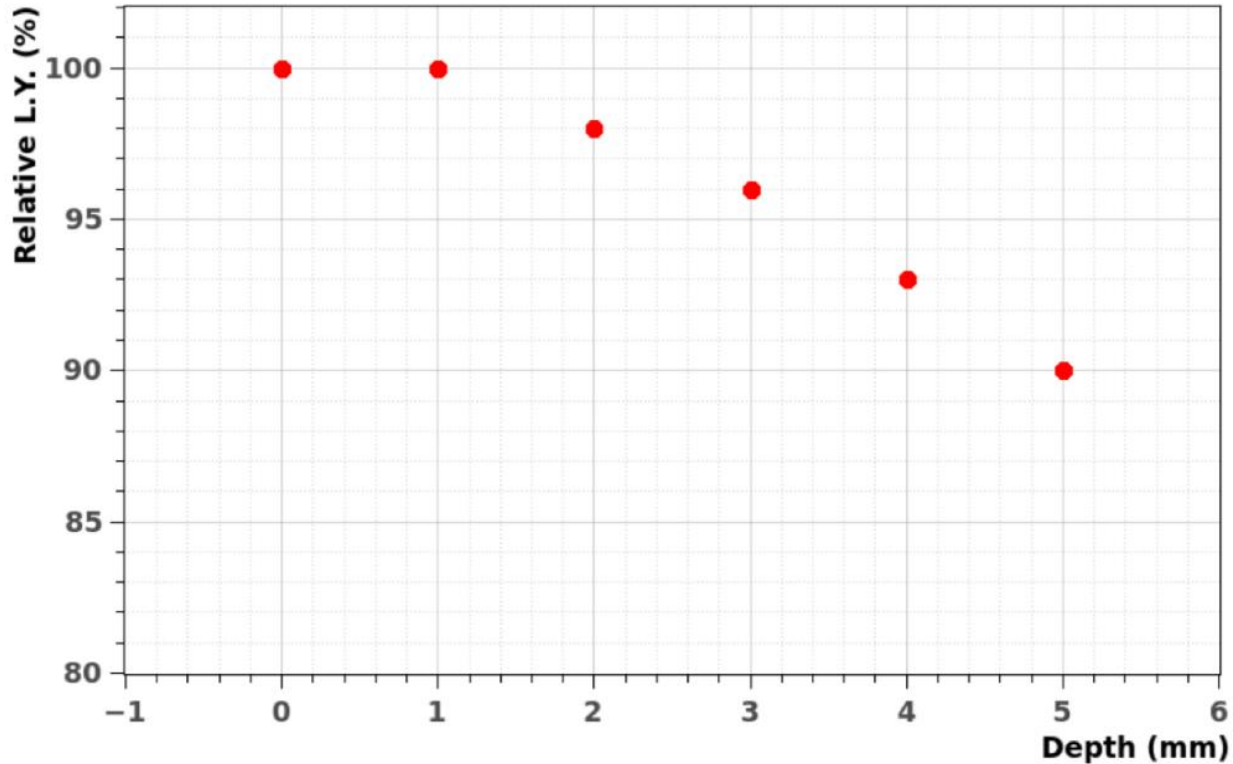
**典型光谱响应曲线**  
**Typical spectral response curve**



**典型增益曲线**  
**Typical gain curve**

Pic.13. NNVT2031 specification.

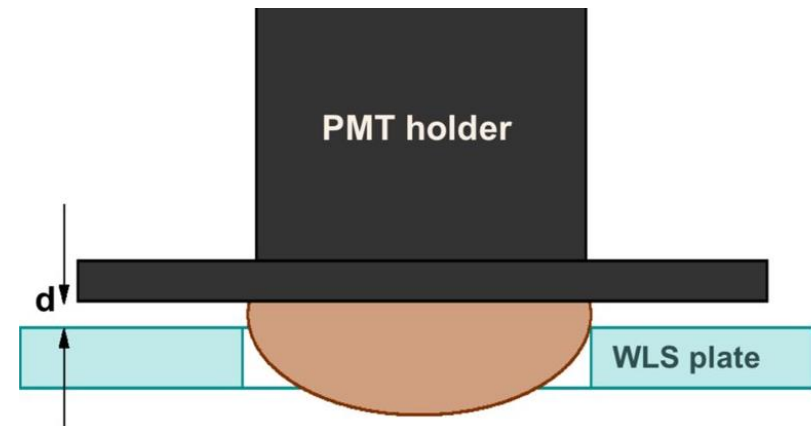
# SEAT DEPTH



Pic.14. Relative L.Y. on PMT's seat depth against WLS plate.

Seat depth is important parameter for complex usage with PMT. WLS plate covers cathode surface of PMT and influences on direct light. The right distance-parameter can provide the most PDE.

All signals is normalized on 0-depth L.Y. As shown on plot, the most optimal depth is 1mm. Cause PMT bulb doesnt contact the WLS plate.



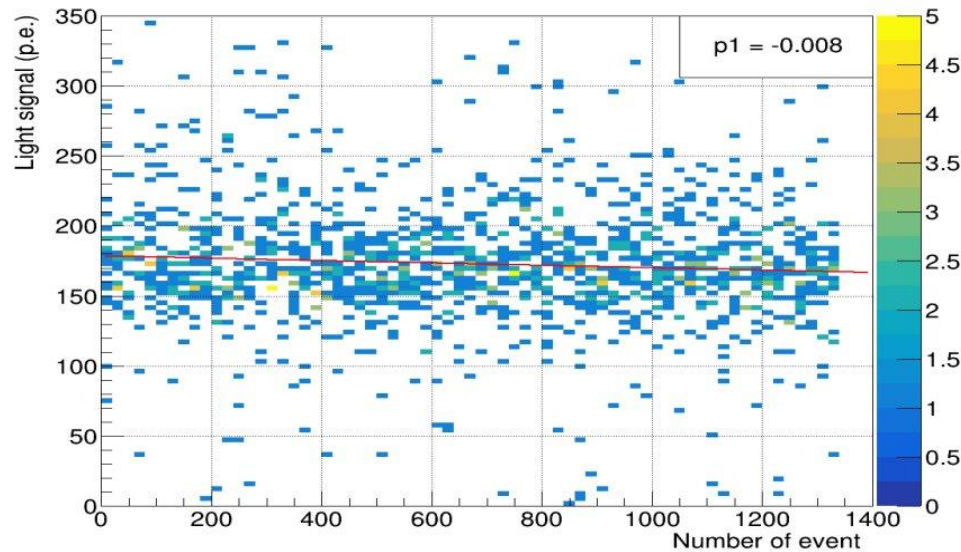
Pic.15. Schematic illustration of PMT's position against WLS plate.

# INFANT-K SETUP



InfantK setup is a small water Cherenkov detector ( $68 \times 33 \times 58 \text{ cm}^3$ ).  
Volume of water = 100 L.

Light signal at one WLS plate vs sequential number of event during 2 days



A few calibrations were made to see the rate of degradation. 4 different plates (both single and double-fluor ones) were re-measured with time interval of  $\sim 40$  days.

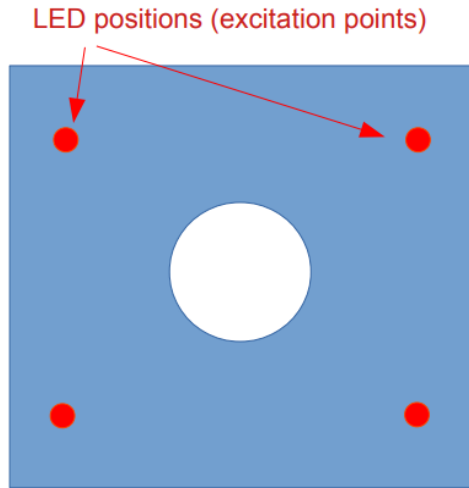
The average light signal degradation was calculated as:

**0.86% per day** for double fluor plates  
**0.77% per day** for single fluor plates

We have measured the signal degradation over time for only PMT (PMT is optically isolated from a WLS plate). PMT has shown the rate: **0.27% per day.**

*No good explanation for difference in degradation rates for PMT and WLS.*

# SPECTRAL RESPONSE CHECK



Pic.16. LED positions for spectral response checking.

Looks like the WLS plate with double fluor does not see short UV due to InfantK tests. We have to check the spectral response of the WLS plates to UV range

Measurement procedure for each LED wavelength: LED irradiates a WLS plate in 4 fixed points. The measurement of each plate was done twice. The average of all results for the plate of the same type was calculated. POPOP50PPO3000 is a calibration plate.

| LED wavelength      | 265 nm | 315 nm | 380 nm | 405 nm |
|---------------------|--------|--------|--------|--------|
| POPOP 50 — PPO 3000 | 100%   | 100%   | 100%   | 100%   |
| POPOP 50            | 13%    | 133%   | 124%   | 118%   |
| POPOP 100           | 33%    | 130%   | 105%   | 142%   |
| POPOP 200           | 46%    | 124%   | 100%   | 160%   |
| POPOP 400           | 84%    | 154%   | 117%   | 230%   |
| POPOP 800           | 100%   | 153%   | 124%   | 261%   |

Tested plates:

- POPOP50PPO3000 - 3 pcs
- POPOP50 - 3 pcs
- POPOP800 - 2 pcs
- All others - 1 piece

# REFLECTORS

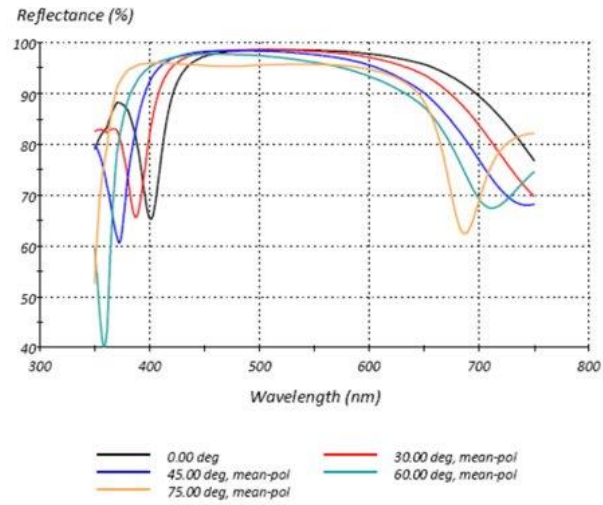
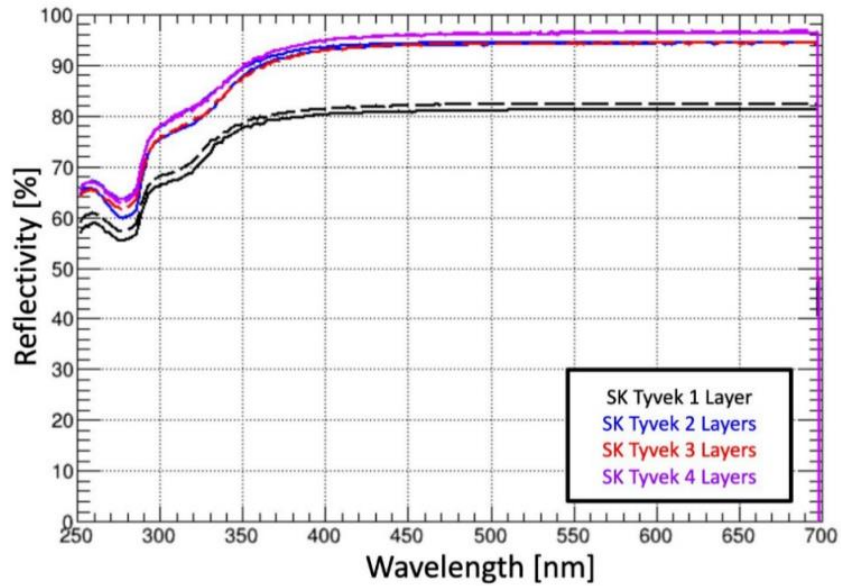
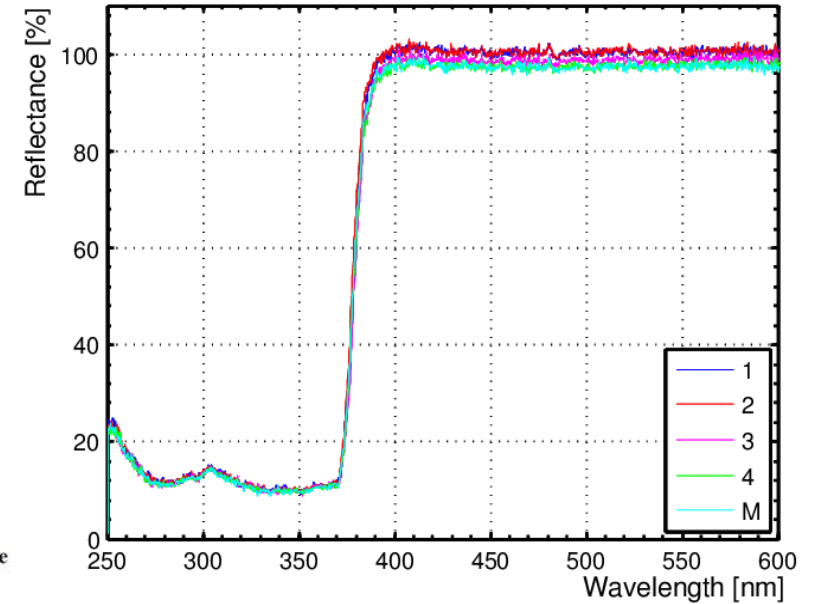
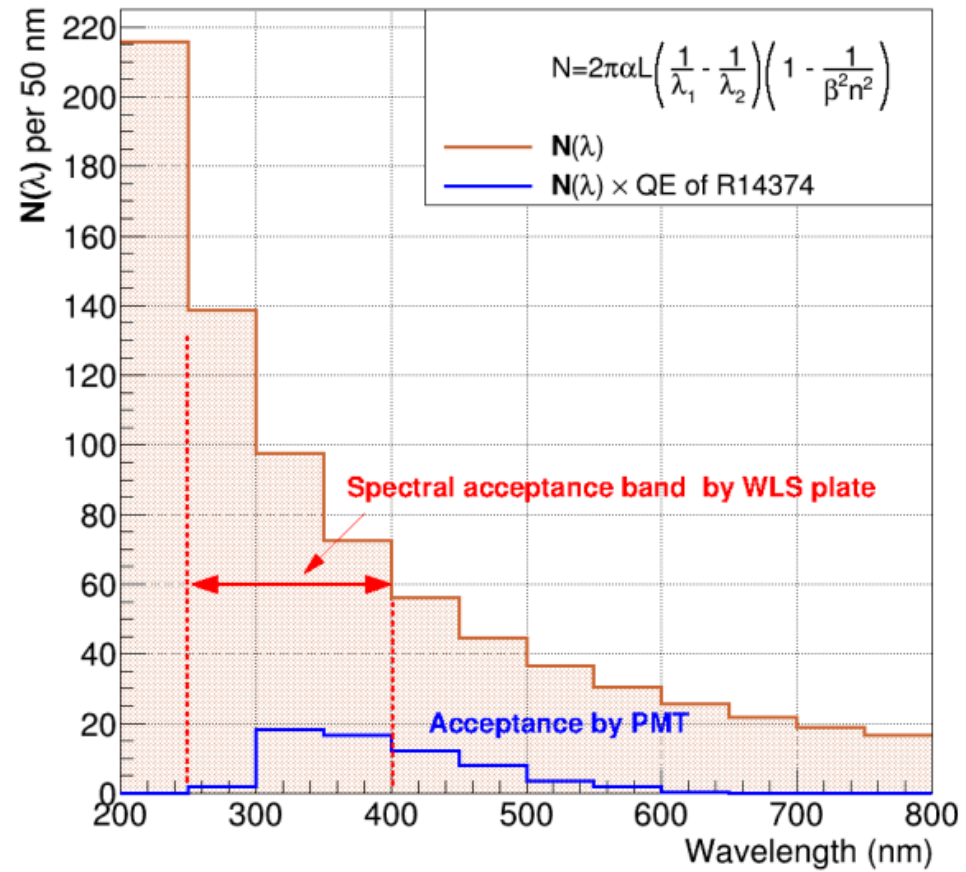


Fig. 6: Enhanced aluminum spectral reflectance as function of angle of incidence

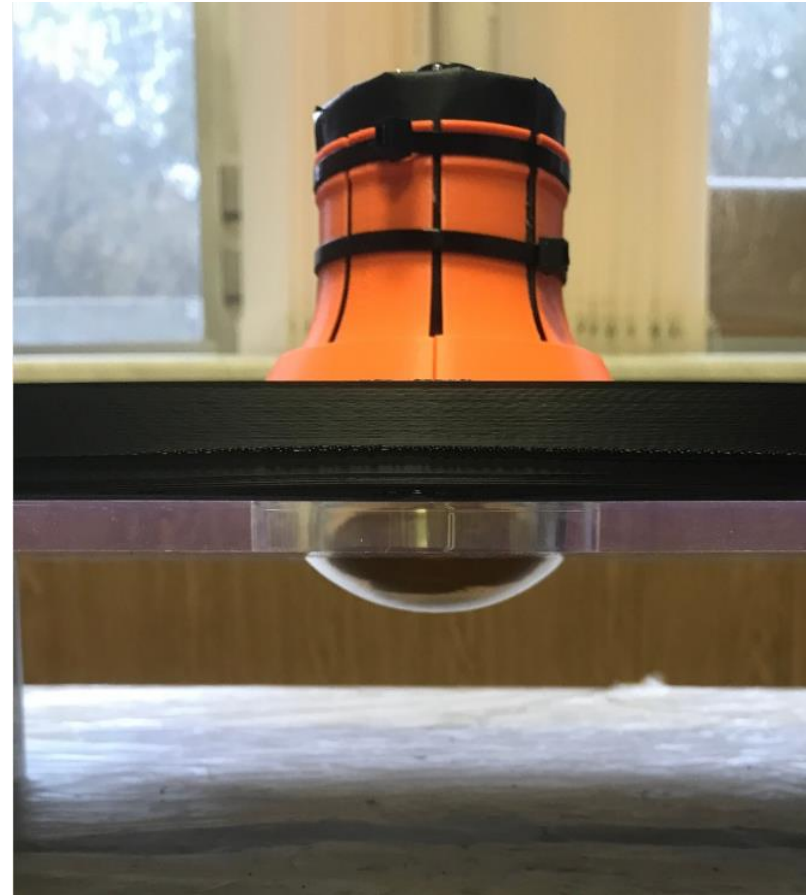


Pic.17. Reflectance for different materials: Tyvek, Mylar, 3D DF2000MA (from left to right).

# ADDITIONAL



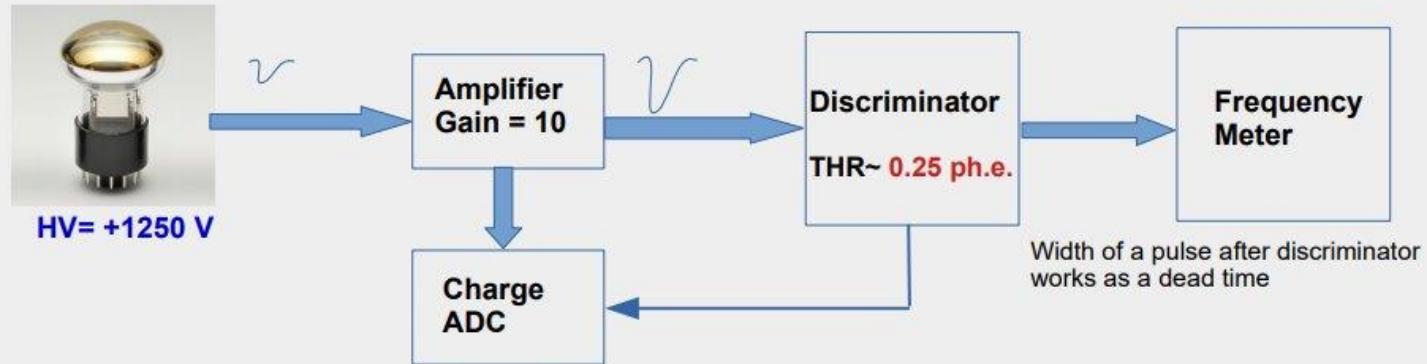
Pic.18. Spectral acceptance for optical module's components.



Pic.19. Mount of PMT and WLS plate.

# ADDITIONAL

## Dark rate of R14374 Hamamatsu PMT at **positive** HV



The measurements were made inside a **temperature controlled volume**.

Tested setup: **PMT + WLS plate with a single fluor (POPOP50)** wrapped by 3M DF2000MA cladding reflector.

Discriminated pulse width was set to be **1.5  $\mu$ s** to suppress:

- the generation of multiple transition pulses with low threshold at falling pulse tail;
- multiple correlated pulses at PMT discharges (pulse bursts)
- afterpulses