

Supernova neutrino search and underground physics at LVD

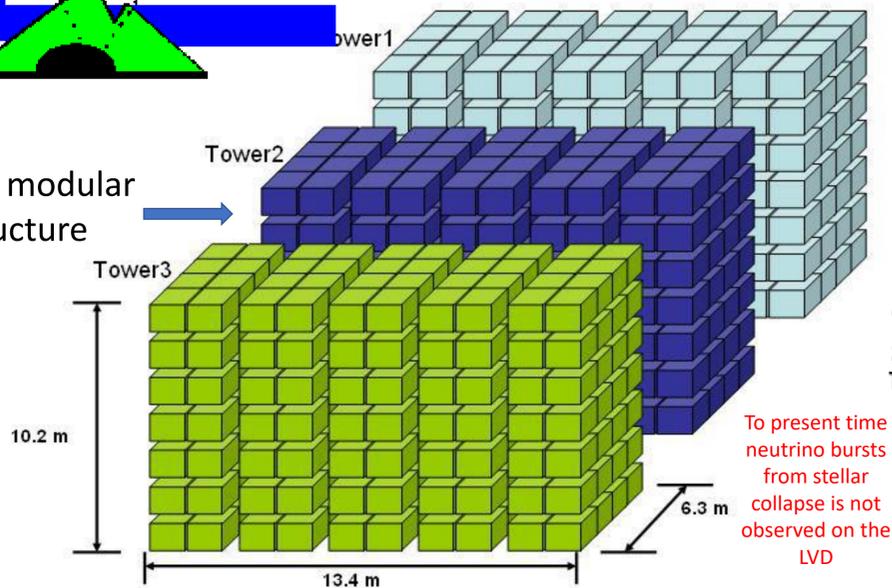


physics at LVD

Natalia Agafonova, Vsevolod Ashikhmin on behalf of LVD collaboration



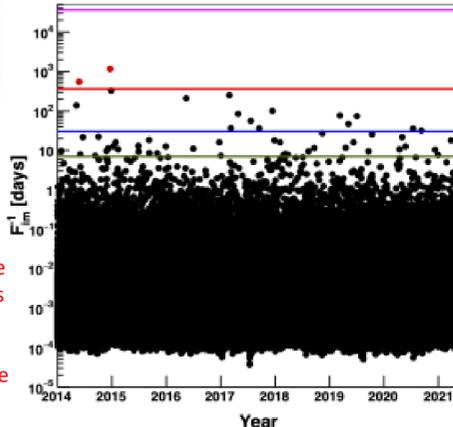
LVD modular structure



To present time neutrino bursts from stellar collapse is not observed on the LVD

$$F_{im_i} = f_{bk_i}^2 \Delta t_{max} \sum_{k \geq m_i - 2} P(k, f_{bk_i} \Delta t_i)$$

$P(k, f_{bk_i}, \Delta t_i)$ is the Poisson probability to have k events in the time window Δt_i if f_{bk_i} is the background frequency



Distributions of detected clusters versus time for the present dataset. Red dots represent clusters with imitation frequency less than $F_{im} = 1 \text{ year}^{-1}$. Green, blue, red and purple lines corresponds to $F^{th} = 1 \text{ week}^{-1}$, $F^{th} = 1 \text{ month}^{-1}$, $F^{th} = 1 \text{ year}^{-1}$, $F^{th} = 1/100 \text{ year}^{-1}$, respectively. Cluster multiplicity greater than or equal than 2 and cluster duration less than or equal 100 seconds

ν interaction channel	E_ν threshold	%
1 $\bar{\nu}_e + p \rightarrow e^+ + n$	(1.8 MeV)	(88%)
2 $\nu_e + {}^{12}\text{C} \rightarrow {}^{12}\text{N} + e^-$	(17.3 MeV)	(1.5%)
3 $\bar{\nu}_e + {}^{12}\text{C} \rightarrow {}^{12}\text{B} + e^+$	(14.4 MeV)	(1.0%)
4 $\nu_i + {}^{12}\text{C} \rightarrow \nu_i + {}^{12}\text{C}^* + \gamma$	(15.1 MeV)	(2.0%)
5 $\nu_i + e^- \rightarrow \nu_i + e^-$	(-)	(3.0%)
6 $\nu_e + {}^{56}\text{Fe} \rightarrow {}^{56}\text{Co}^* + e^-$	(10. MeV)	(3.0%)
7 $\bar{\nu}_e + {}^{56}\text{Fe} \rightarrow {}^{56}\text{Mn} + e^+$	(12.5 MeV)	(0.5%)
8 $\nu_i + {}^{56}\text{Fe} \rightarrow \nu_i + {}^{56}\text{Fe}^* + \gamma$	(15. MeV)	(2.0%)

Reactions of neutrino (antineutrino) interaction with LVD

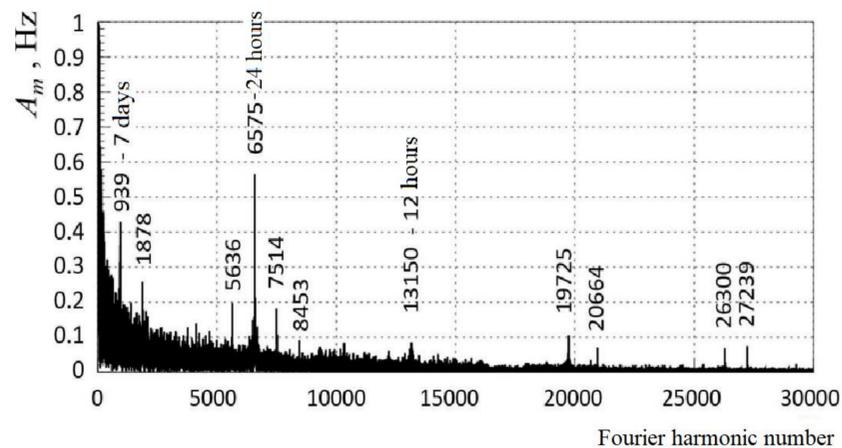
The Large Volume Detector (LVD) is located underground at a depth of 1400 m under rock (minimal depth 3000 m w.e.), in the INFN Gran Sasso National Laboratory (Italy). The experiment consists of an array of 840 scintillator counters, 1.5 m³ each, viewed from the top by three photo multipliers (PMTs) and arranged in a modular geometry[9]. This modularity allows LVD to achieve a very high duty cycle, that is essential in the search of unpredictable sporadic events. Failures involving one or more counters do not affect, in general, other counters. The detector maintenance can be done during data acquisition by stopping only the part that needs to be maintained, even a single counter. This peculiarity allows a dynamic active mass M_{act} and a high duty cycle. The experiment has been in operation since 1992, June 9th after a short commissioning phase, its mass increasing from 300 t to its final one, 1000 t, at time of building phase completion in January 2001.

The main goal of the experiment is detection of neutrinos from the gravitation collapse of stellar cores due to IBD reaction ($\bar{\nu}_e + p \rightarrow e^+ + n$). LVD has been observing our Galaxy searching for neutrino bursts since 9 June 1992. The detector is one of the founding members of the **SuperNovae Early Warning System** project.

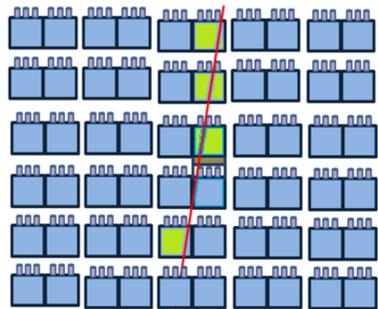
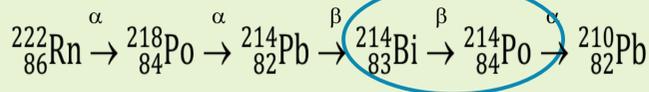
Also LVD be able to register natural radioactivity products, and in particular, from radon decay.

Length× Width× Height	22.7×13.2×10 m
Iron mass	1020 t
Scintillator mass	1008 t
Amount of counters	840
Amount of PMTs	2520
Average depth (min)	3620 m w.e. 3000 m w.e.
Average muon energy	280 GeV
E_μ at see lev. (min.)	1.3 TeV
Muon rate (per 1 tower)	~ 120 h ⁻¹
ϵ_{th} threshold -inner -external	4 MeV, 7 MeV

Thanks to a lot of time of radon counting rate (during to 18 years of observation) measuring we have got its Fourier spectrum



Monitoring of radon concentrations is possible owing to the detection of gammas from decays of daughter nuclei of the radon isotope ²²²Rn.



$$n + p \rightarrow d + \gamma, E_\gamma = 2.2 \text{ MeV}$$

$${}^{56}\text{Fe} \rightarrow {}^{57}\text{Fe} + \Sigma \gamma, \langle E_\gamma \rangle \approx 7 \text{ MeV}$$

250 ns

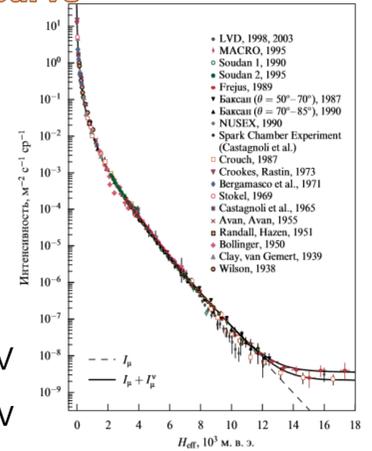
Muon

τ ~ 150 μs

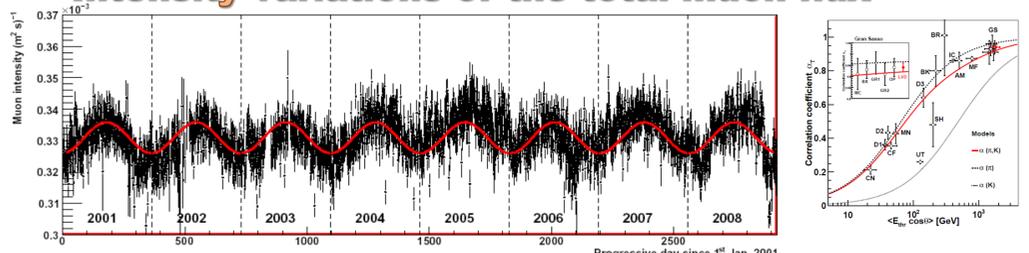
Neutron(s)

capture

depth-intensity curve

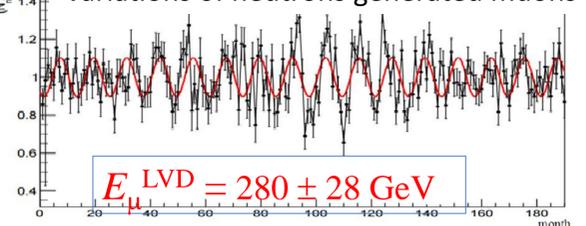


Intensity variations of the total muon flux



$$I_\mu = I_0^\mu + \delta I^\mu \cos\left(\frac{2\pi}{T}(t - t_0)\right)$$

Variations of neutrons generated muons



$$Y_n(E_\mu) \propto E_\mu^{0.78}$$

$$E_\mu^{LVD} = 280 \pm 28 \text{ GeV}$$

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