



DarkSHINE – Search for Light Dark Matter at the SHINE facility in Shanghai

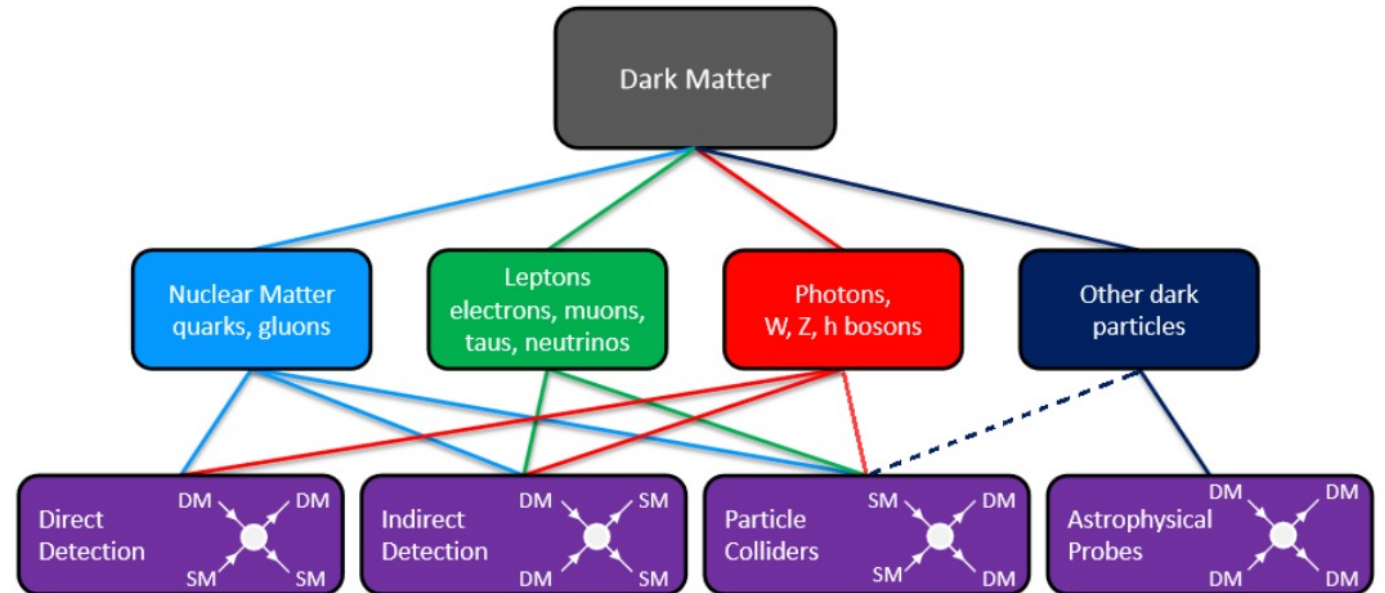
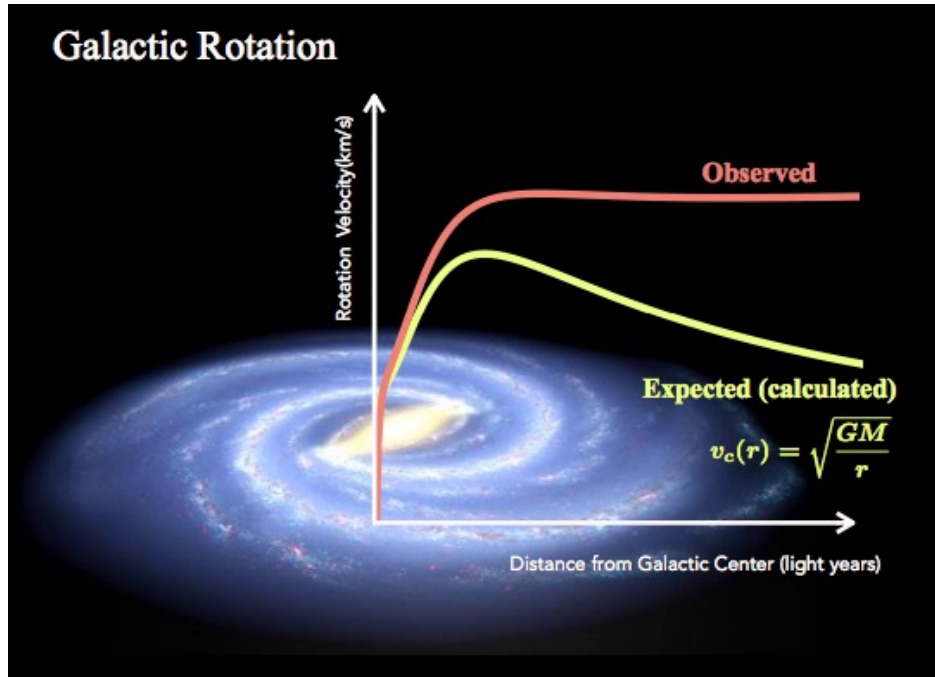
Danning Liu, on behalf of the DarkSHINE R&D Term

October, 2024, Moscow, Russia



李政道研究所
TSUNG-DAO LEE INSTITUTE

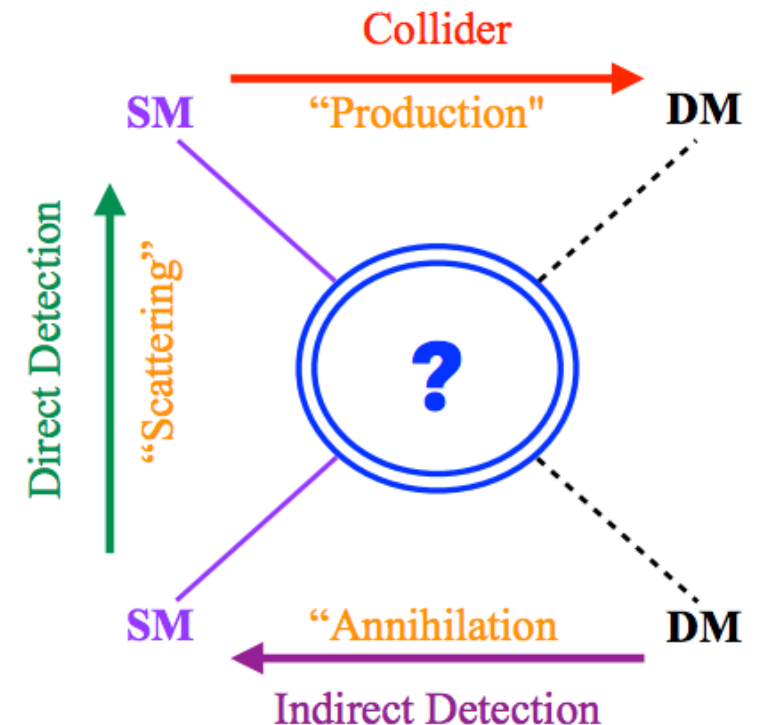
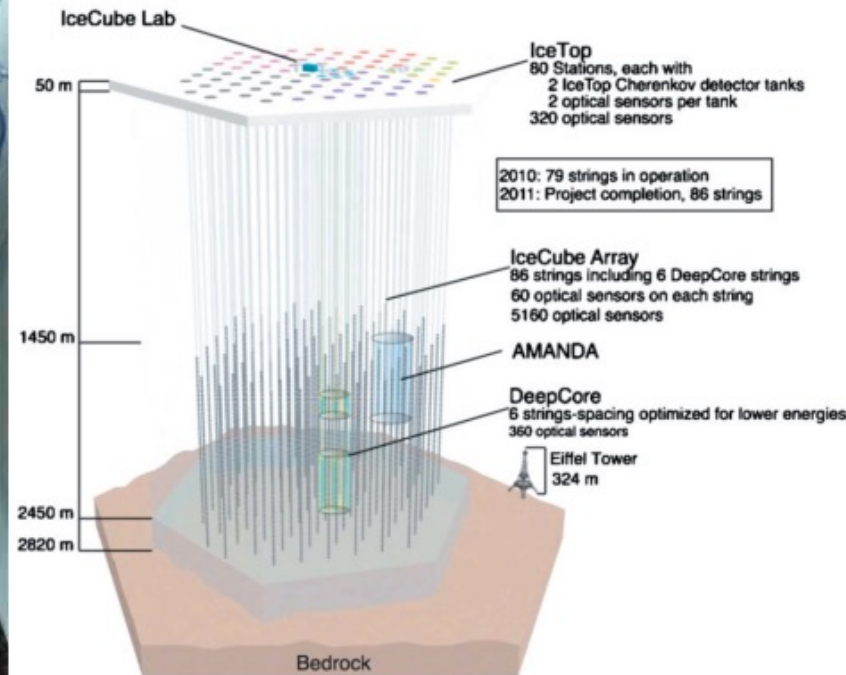
Dark Matter Evidence



- Dark Matter (DM) evidence, from astronomical observations and gravitational effects
 - Galactic rotation curves, Gravitational lensing, Cosmic Microwave Background ...

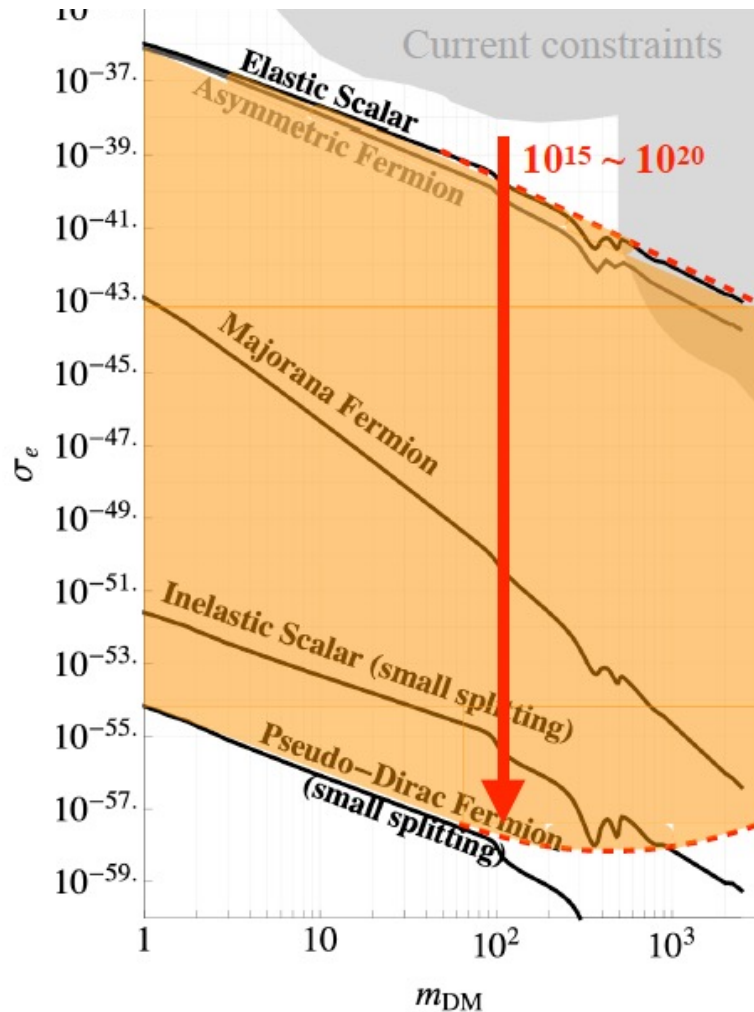
Dark Matter Detection

- **Direct Detection:** nuclear recoils from DM-nuclei scattering (PandaX, XENONnT, LZ, CDEX ...)
- **Indirect Detection:** products from DM annihilation (DAMPE, IceCube ...)
- **Colliders:** DM production in high-energy collisions, especially focusing on the productions of a SM particle with large missing E_T

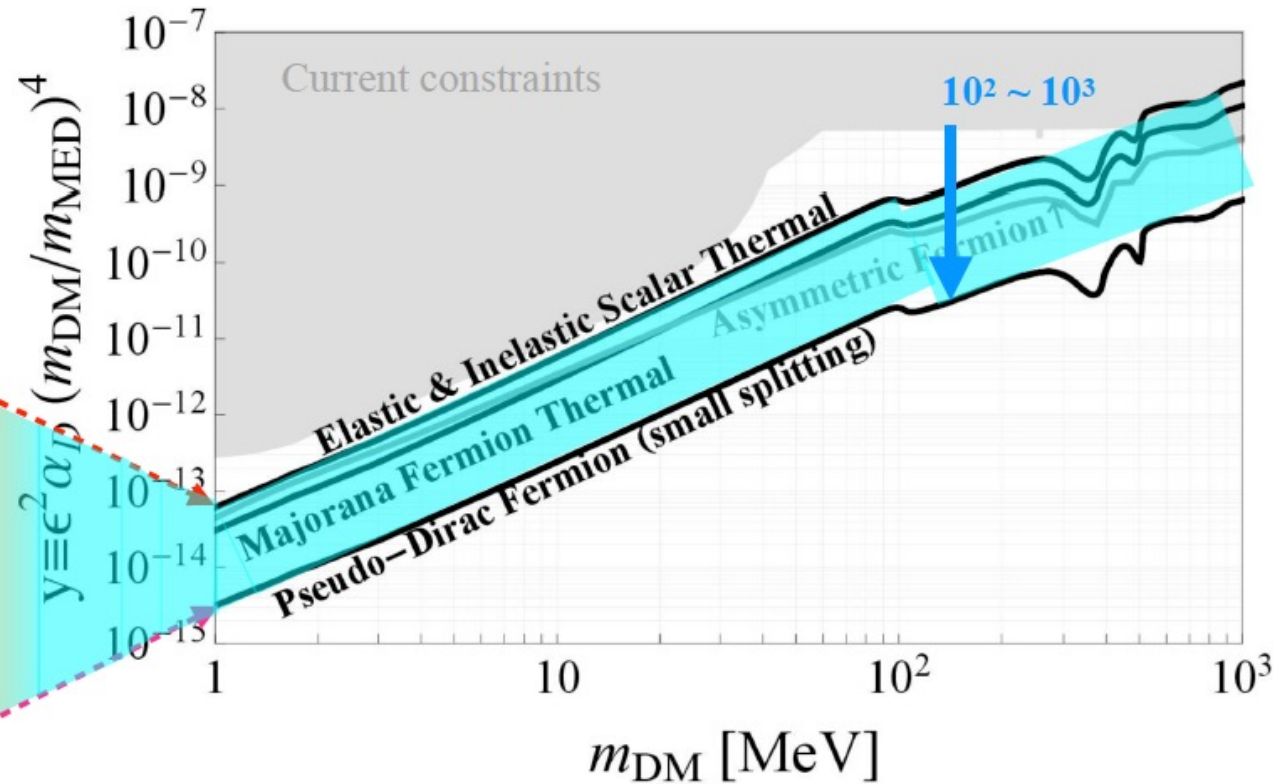


Dark Matter Search at Accelerator Experiments

Direct Detection



Accelerator based experiments



- In accelerator-based experiments, this difference can be reduced to $10^2 \sim 10^3$ orders of magnitudes, due to the fact of insensitive to DM's mass and spin in its production.

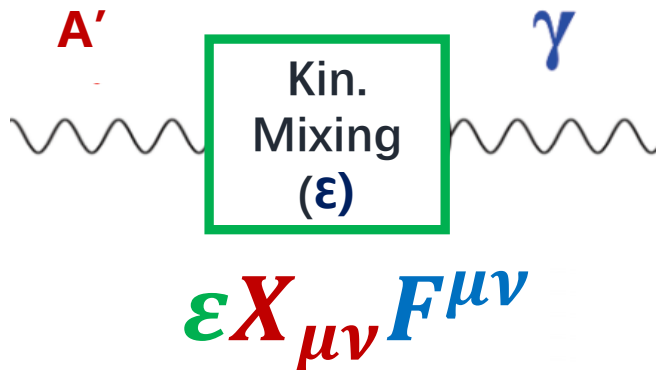
Dark Photon Theory

- Introduce extra $U(1)_X$ symmetry \rightarrow New Gauge Field $X \rightarrow$ Dark Photon Mediator A'
- $U(1)_{em} \rightarrow U(1)_{em} \times U(1)_X$

$$\mathcal{L} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + A_{\mu} j_{em}^{\mu}}_{\text{SM Photon } \gamma} \underbrace{-\frac{1}{4} X_{\mu\nu} X^{\mu\nu} + X_{\mu} j_X^{\mu}}_{\text{Dark Photon } A'}$$

SM Photon γ

Dark Photon A'



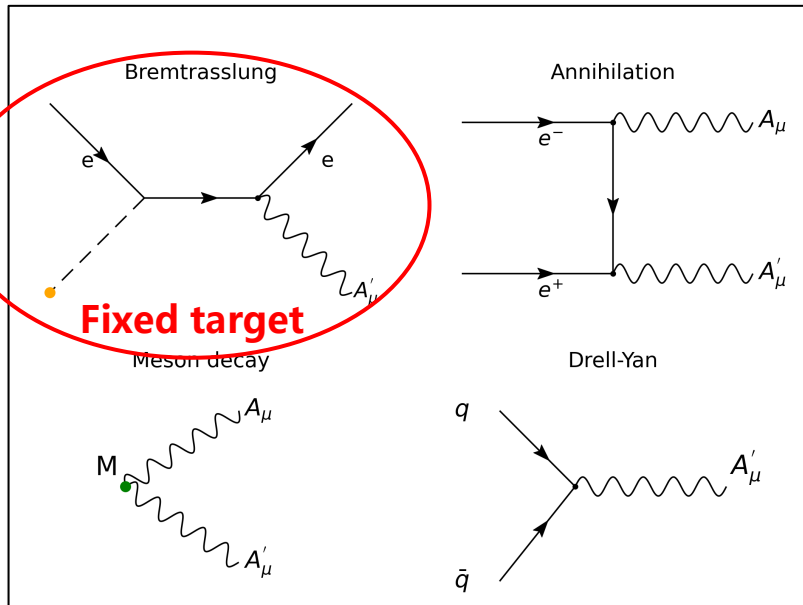
- A' & γ kin. mixing
- Renormalizable and Gauge Invariant
- Straightforward for experimental search
 - Free param, kin. mixing (ϵ)、 mass ($m_{A'}$)

B. Holdom, Phys. Lett. B 166, 196 (1986)

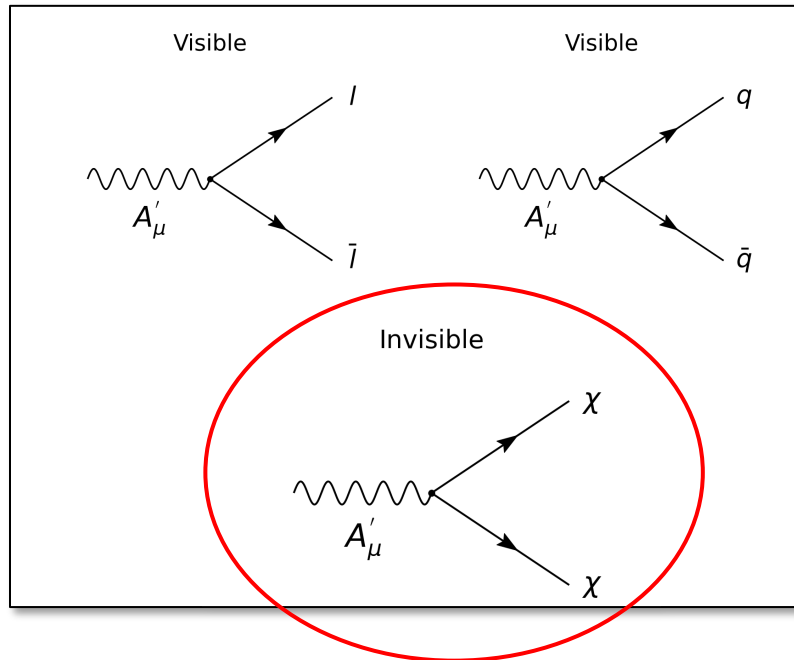
R. Foot & X.-G. He, Phys. Lett. B 267, 509 (1991)

Physics Processes and Anticipated Signatures

Processes to search for **dark photon A'** : Bremsstrahlung, Annihilation, Meson decay and Drell-Yan process



(Dark photon production)

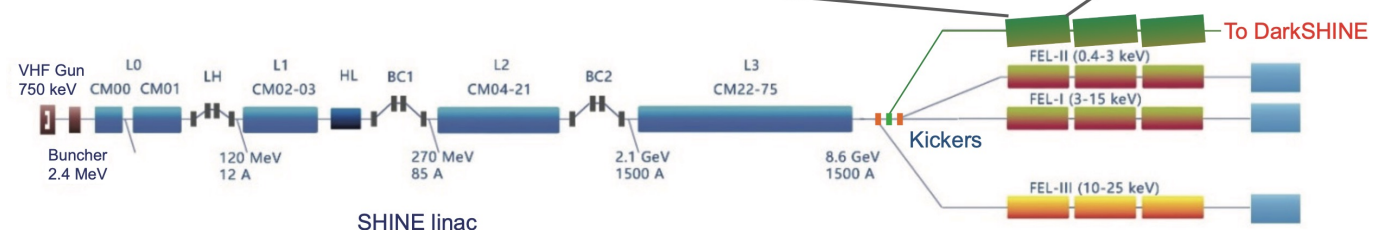
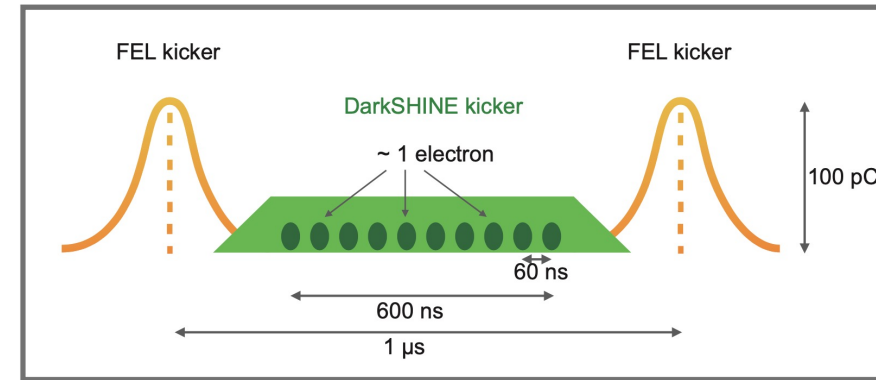
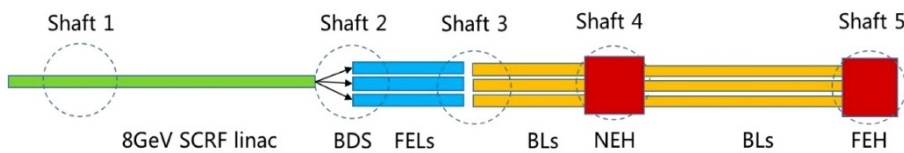


(dark photon decay)

- **Goal:** put constraints on the kinetic mixing parameter ϵ .
- **Challenge:** small production rate \rightarrow suppress bkg. from SM processes.
- **Experimental signatures:** missing energy, missing momentum.

The SHINE Facility

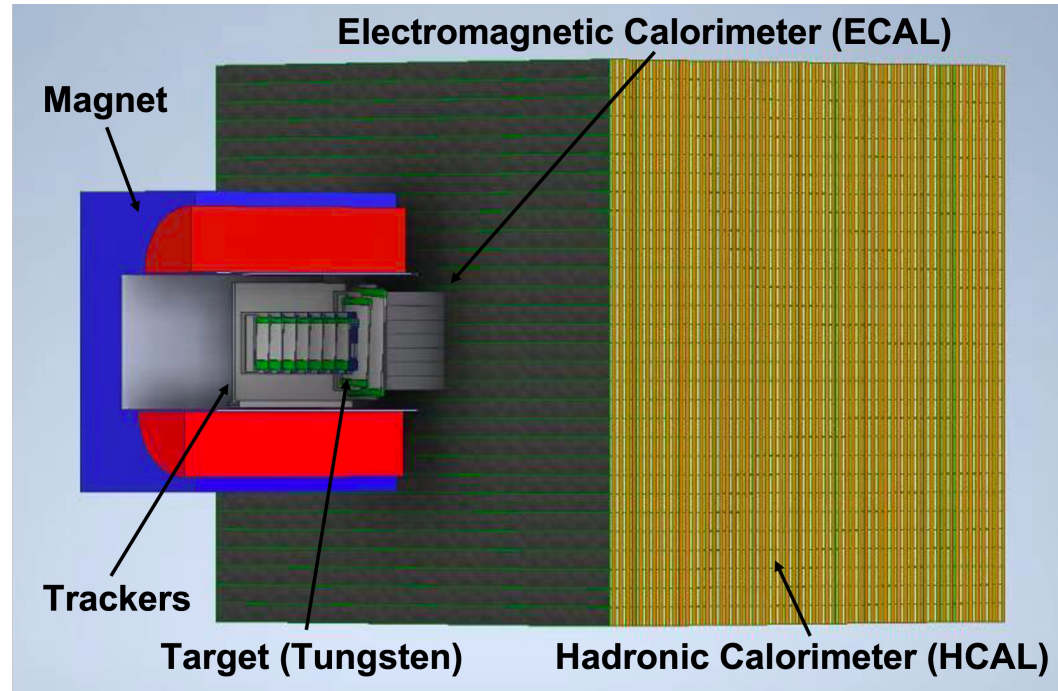
- Shanghai High Repetition-Rate XFEL and Extreme Light Facility (SHINE) can provide high repetition rate single electron beams → with dedicated kicker to be designed and deployed
 - Electron energy: 8 GeV, Rep. Rate: 1 MHz
 - Beam intensity: 100 pC (6.25E8 electrons/bunch)
 - $\sim 3 \times 10^{14}$ electrons-on-target (EOT) per year



DarkSHINE Experiment Conceptual Design

The Dark SHINE detector hardware technical R&D is carried out in parallel to the full detector system simulation and prospective study/optimization

Tracking system
Measure the tracks of the incident and recoil electrons.



Electromagnetic calorimeter
Measure the deposited energy: electron and photon.

Hadronic calorimeter
Measure the deposited energy: veto muon and hadron backgrounds.

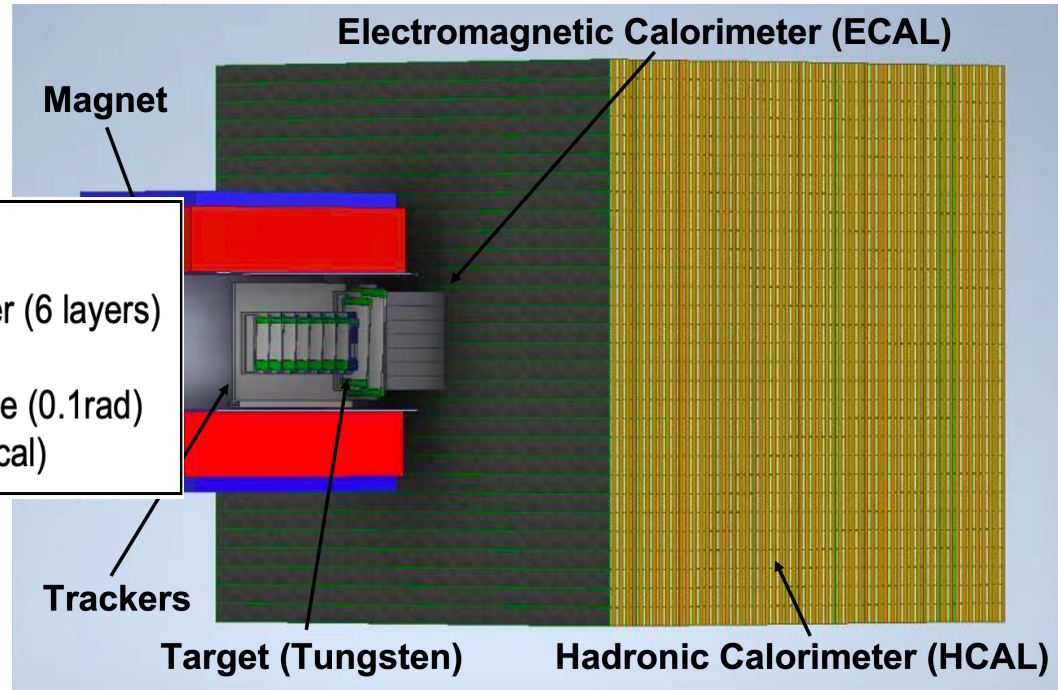
Dark SHINE detector sketch

Additional system:
Readout electronics, trigger system, TDAQ, magnetic system (1.5 T), etc.

DarkSHINE Experiment Conceptual Design

The Dark SHINE detector hardware technical R&D is carried out in parallel to the full detector system simulation and prospective study/optimization

- **Tracker:**
 - Tagging tracker (7 layers) + recoil tracker (6 layers)
 - Incident and recoil electron tracks
 - Two silicon strip sensors w/ a small angle (0.1rad)
 - Resolution: $6\mu\text{m}$ (horizontal), $60\mu\text{m}$ (vertical)



Dark SHINE detector sketch

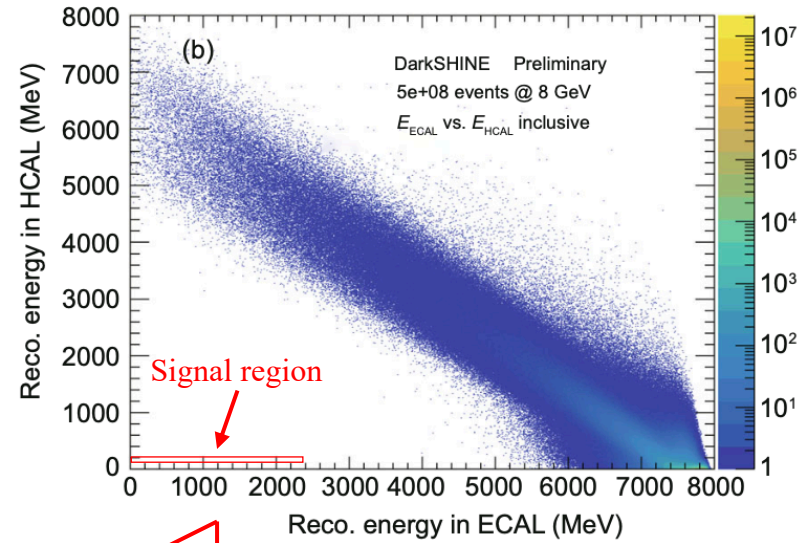
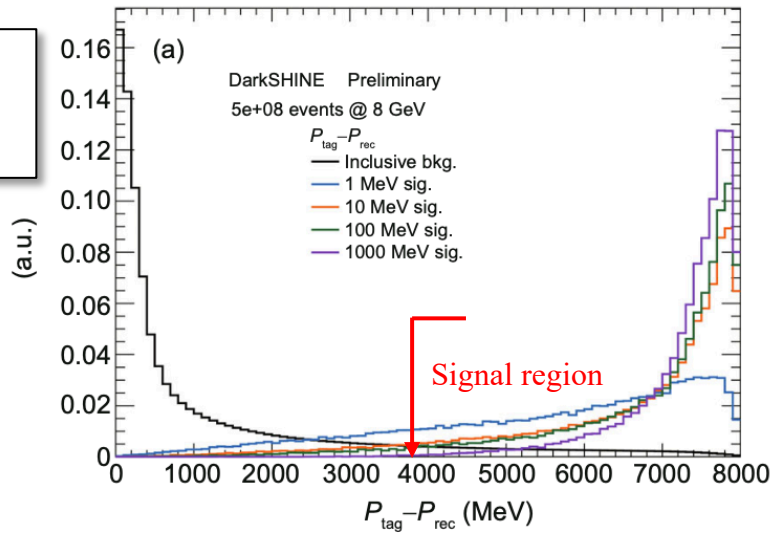
- **Additional system:**
 - Magnet: 1.5T magnetic field
 - Readout electronics

- **ECAL:**
 - Electron & photon
 - Scintillator: LYSO(Ce)
 - high light yield (30000 p.e./MeV), fast decay time (40 ns), low electronic noise
 - $20\times 20\times 11$ crystals
 - $2.5\times 2.5\times 4\text{cm}^3$
 - Energy resolution of LYSO: 5%

- **HCAL:**
 - Veto hadronic background
 - Scintillator w/ steel absorber
 - $4\times 4\times 1$ modules

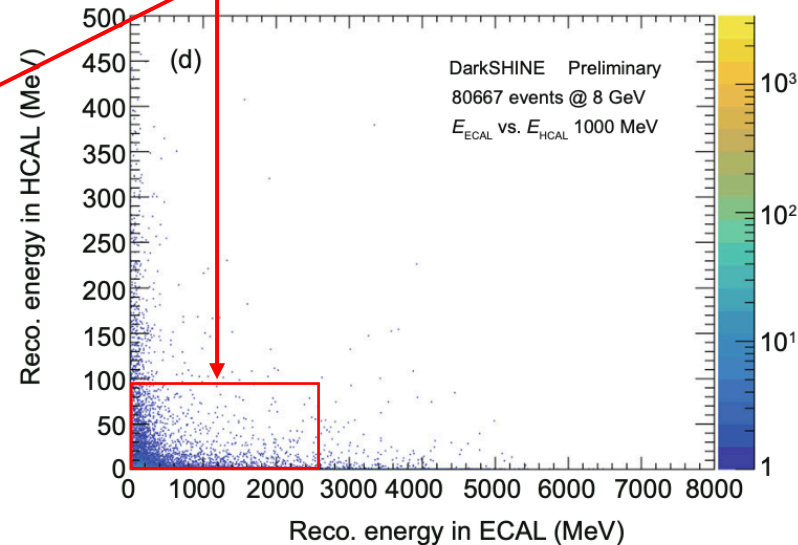
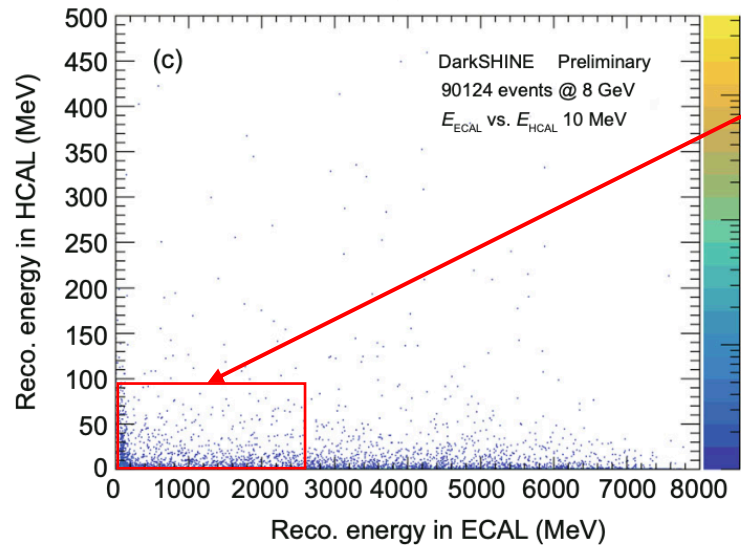
Determination of Signal Selections

Inclusive bkg
& signal

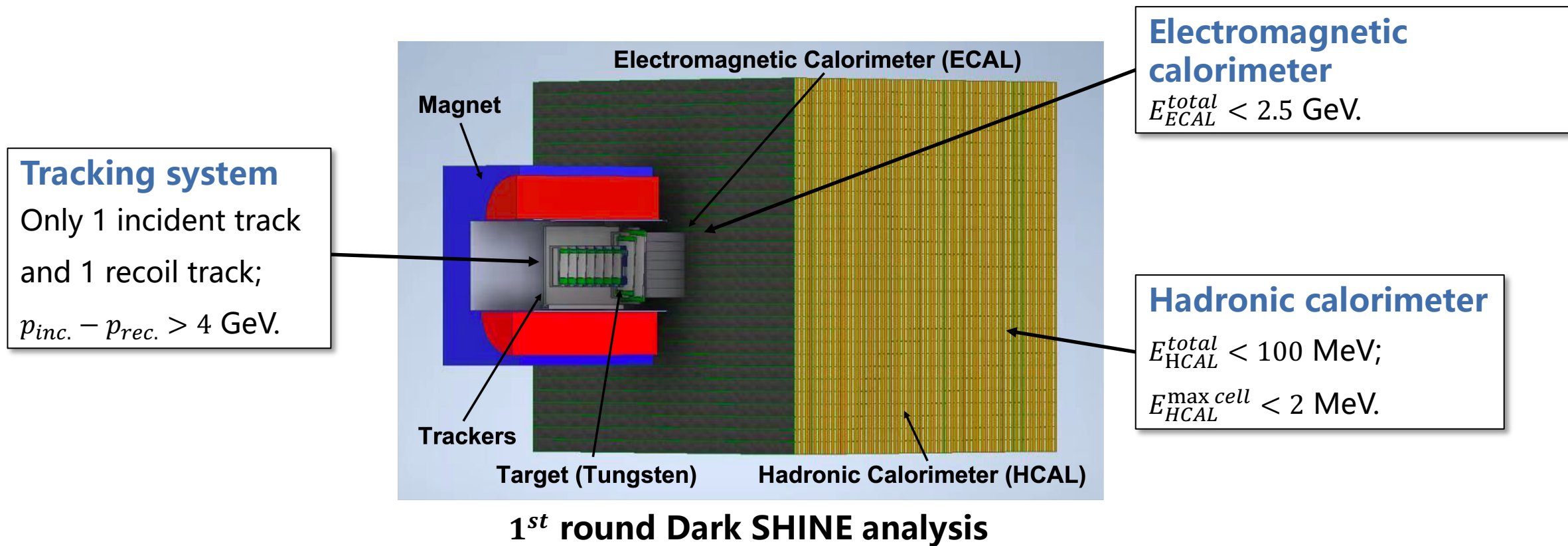


Inclusive bkg

10 MeV A'

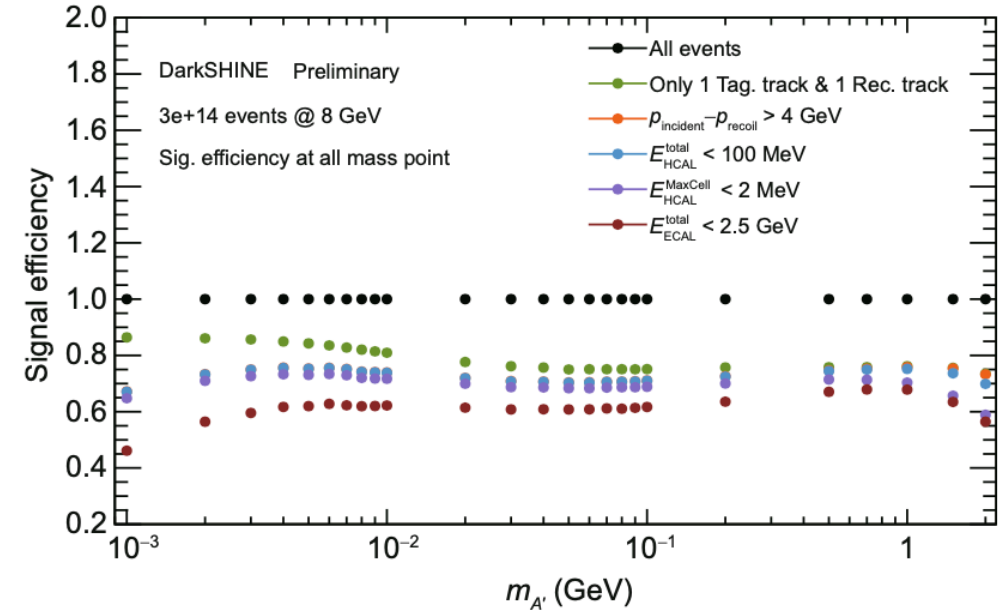


1 GeV A'



Analysis Details

- Cut Efficiencies:
 - Signal Cut efficiencies as function of dark photon mass
 - 67% efficiency arrived and applied to extract significance
- Event Cutflow for each background samples on event numbers



	EN_ECAL	PN_ECAL	GMM_ECAL	EN_target	PN_target	GMM_target	Hard_brem	Inclusive
Total events	2.48×10^7	1.66×10^8	1.74×10^7	1.09×10^8	1.05×10^7	1.05×10^7	1.02×10^7	2.50×10^9
Only 1 track	1.46×10^7	1.17×10^8	1.52×10^7	6.38×10^6	6.17×10^5	77	8.03×10^6	2.11×10^9
$p_{\text{tag}} - p_{\text{rec}} > 4 \text{ GeV}$	1091	5531	707	6.08×10^6	5.73×10^5	1	7.19×10^6	1.20×10^8
$E_{\text{HCAL}}^{\text{total}} < 100 \text{ MeV}$	135	1348	0	322135	75501	0	1.19×10^8	2.89×10^7
$E_{\text{HCAL}}^{\text{MaxCell}} < 10 \text{ MeV}$	56	676	0	141808	27949	0	1.12×10^8	2.72×10^7
$E_{\text{HCAL}}^{\text{MaxCell}} < 2 \text{ MeV}$	30	363	0	63644	9999	0	1.01×10^8	2.46×10^7
$E_{\text{ECAL}}^{\text{total}} < 2.5 \text{ GeV}$	0	0	0	0	0	0	0	0



Prospective Study of Light Dark Matter Search with a Newly Proposed DarkSHINE Experiment

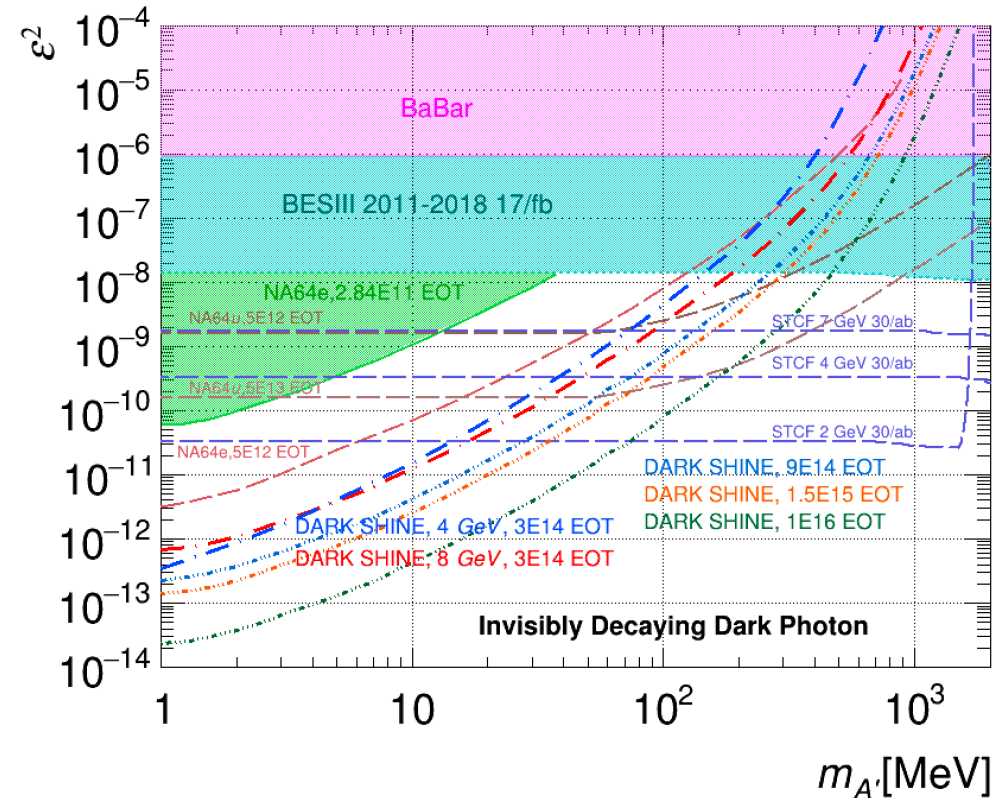
Jing Chen^{†b,c,a}, Ji-Yuan Chen^{b,c}, Jun-Feng Chen^h, Xiang Chen^{b,c}, Chang-Bo Fu^{ij}, Jun Guo^{b,c}, Le He^f, Zheng-Ting He^{a,n}, Kim Siang Khaw^{a,b,c}, Jia-Lin Li^{b,c}, Liang Li^{b,c}, Shu Li^{a,b,c,d,e*}, Meng Lv^g, Dan-Ning Liu^{a,b,c}, Han-Qing Liu^{b,c}, Kun Liu^{a,b,c*}, Qi-Bin Liu^{a,b,c}, Yang Liu^{a,b,c}, Ze-Jia Lu^{b,c}, Cen Mo^{b,c}, Si-Yuan Song^{b,c}, Xiao-Long Wang^{ij}, Yu-Feng Wang^{†a,b,c}, Zhen Wang^{a,b,c}, Zi-Rui Wang^m, Wei-Hao Wu^{b,c}, Dao Xiang^{k,a,l}, Hai-Jun Yang^{b,c,a*}, Jun-Hua Zhang^{a,b,c}, Yu-Lei Zhang^{†b,c}, Zhi-Yu Zhao^{a,b,c}, Xu-Liang Zhu^{a,b,c}, Chun-Xiang Zhu^{b,c}, and Yi-Fan Zhu^{b,c}

Highlight remarks

Sci. China-Phys. Mech. Astron. 66(1): 211063 (2023)

“Editor’s Focus”

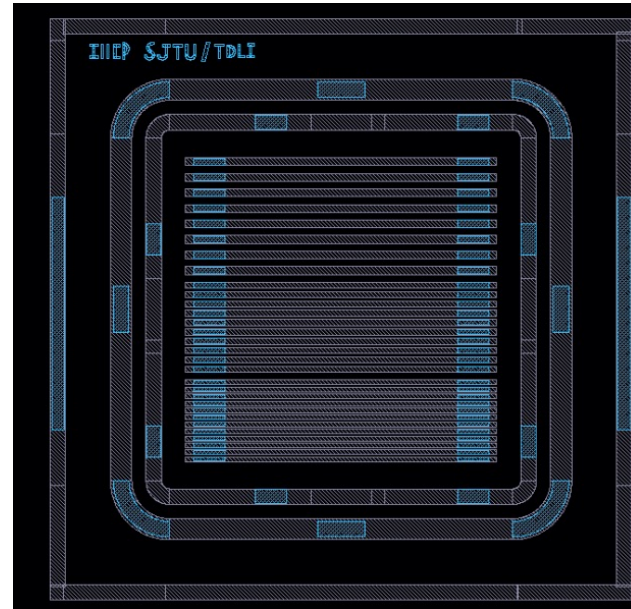
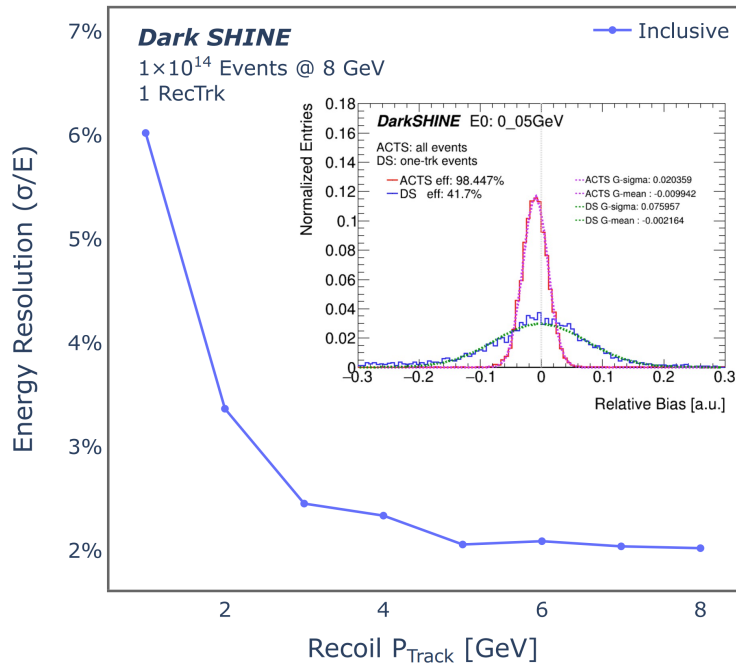
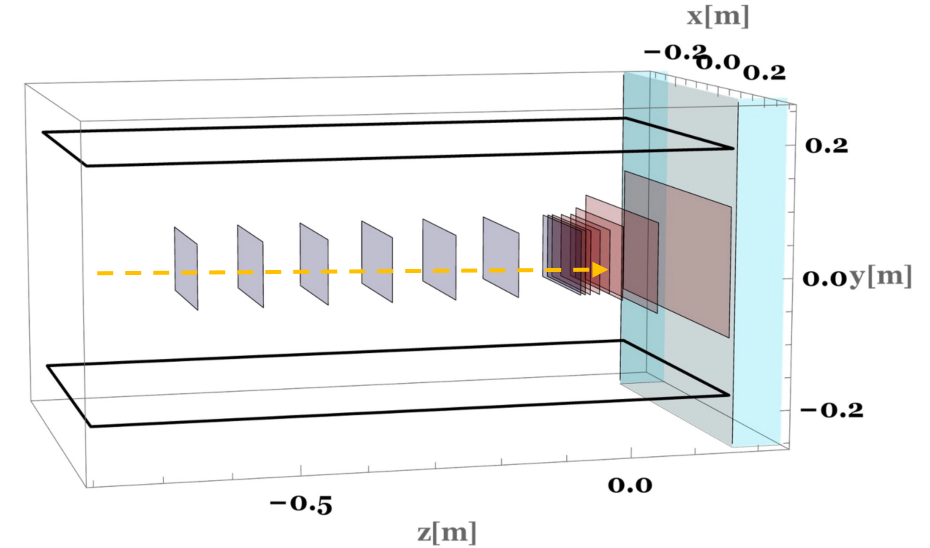
Sci. China-Phys. Mech. Astron. 66(1): 211061 (2023)



- Anticipated to have the sensitivity improved by one order of magnitude compared to other experiment (e.g. NA64)
- Aim to deliver 10^{16} EOT stat. and cover most of the sensitivity regions of interests

Tracker

- Incident and recoil electron tracks.
- Tagging tracker (7 layers) + recoil tracker (6 layers).
- Two silicon strip sensors w/ a small angle (0.1rad).
- Resolution: $10\mu\text{m}$ (horizontal), $60\mu\text{m}$ (vertical).



AC-LGAD silicon strip sensor $1 \times 1 \text{ mm}^2$ designed, in collaboration with Prof. Zhijun Liang and Prof. Mei Zhao from IHEP.

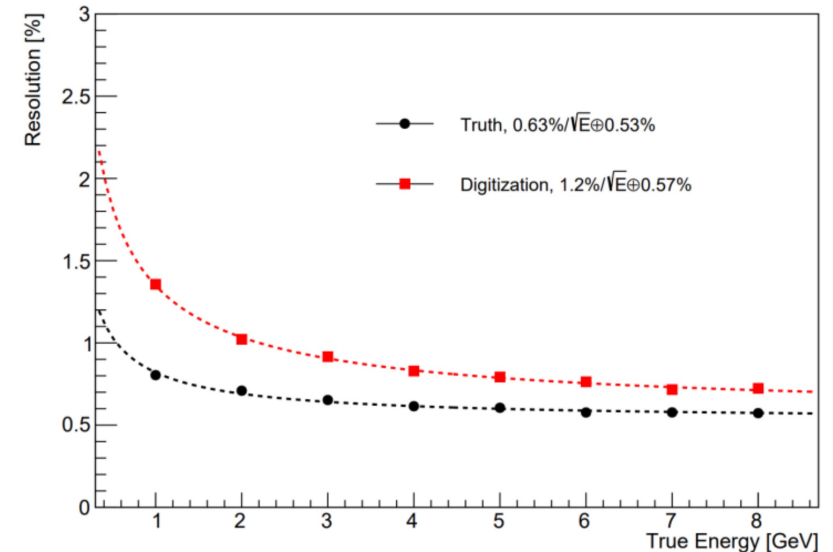
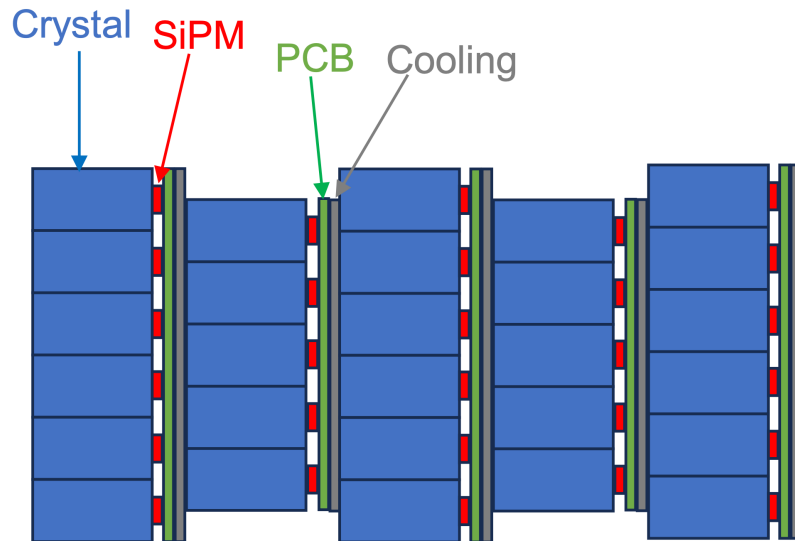
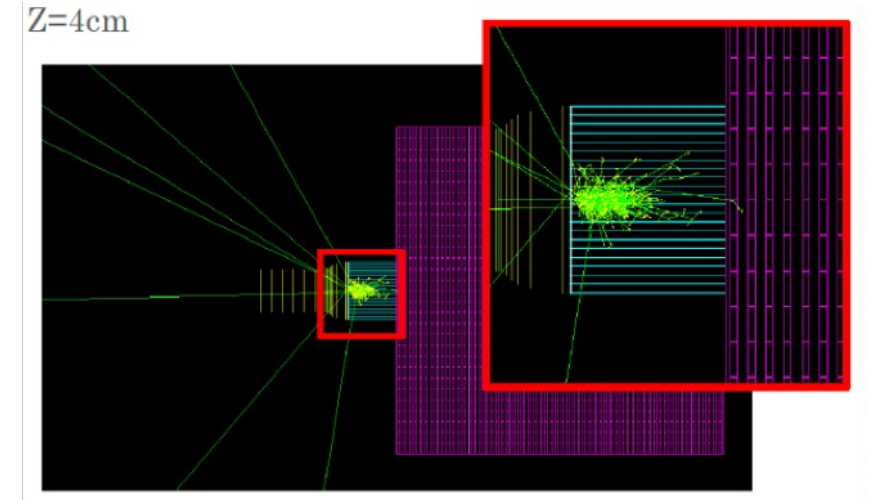
Hardware R&D: Crystal ECAL

arXiv: 2407.17800
submitted to NST



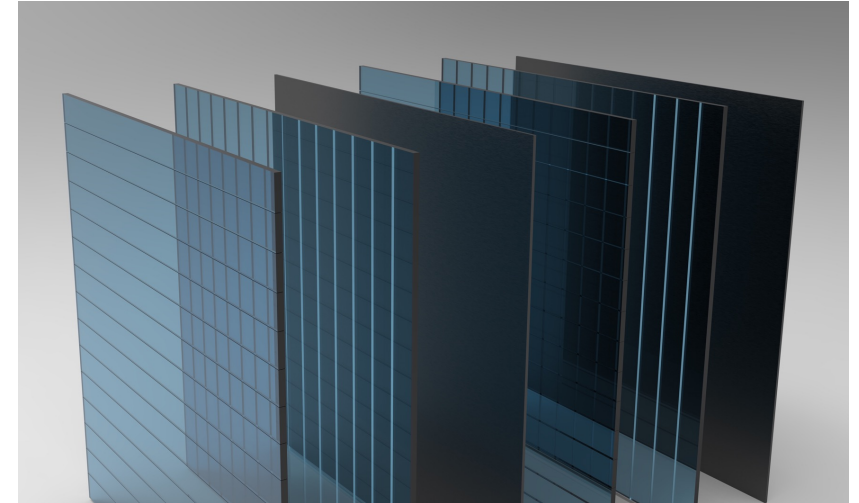
Electromagnetic Calorimeter:

- Designed resolution: better energy resolution than 5%.
- LYSO crystal ($Lu_{(1-x-y)}Y_{2y}Ce_{2x}SiO_5$):
 - High light yield (30000 p.e/MeV) with good linearity
 - Short decay time (40 ns)
- **21×21×11 crystals, 2.5cm×2.5cm×4cm**
- Readout with SiPM and waveform sampling
- More intrinsic radiation and radioactive source tests.



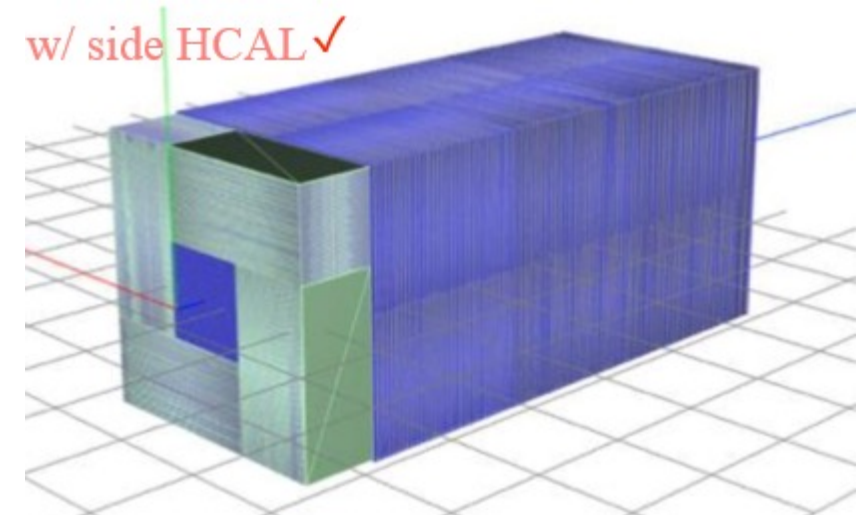
Hadronic Calorimeter:

- Veto backgrounds ($1.5 \times 1.5 \times 2.5 \text{ m}^3$, $\sim 11 \lambda$)
- Plastic scintillator + iron absorber
 - Plastic scintillator: 90 degree rotation between two adjacent layers
 - Wavelength shifter fiber + SiPM
- Side-HCAL: encircling the ECAL
- Design has been optimized



Veto inefficiency

Energy[MeV] \ Particle	n	k^0	π^0	p
100	1.17E-03	3.16E-02	7.30E-06	3.07E-02
500	1.84E-05	3.30E-06	1.00E-07	8.04E-06
1000	3.70E-06	4.30E-06	1.00E-07	1.00E-07
2000	2.70E-06	1.15E-05	1.00E-07	1.00E-07



Collaboration and Highlights at DarkSHINE



Journal Paper Publications:

- **Sci. Chin.-Phys. Mech. Astron.**, 66(1) : 211062 (2023)
- **Nucl. Sci. Tech.** 35, 148 (2024)
- **Nucl. Sci. Tech.** 35, 201 (2024)
- arXiv:2407.17800 (submitted to NST)
- **arXiv:2407.20723 (submitted to JINST)**

Conference Talks :

- **ICHEP** 2024
- **Lepton Photon** 2023
- Dark Matter 2023
- HKUST-IAS-HEP 2024、2023
- AEI 2023、2022
- IPAC 2023
- ...

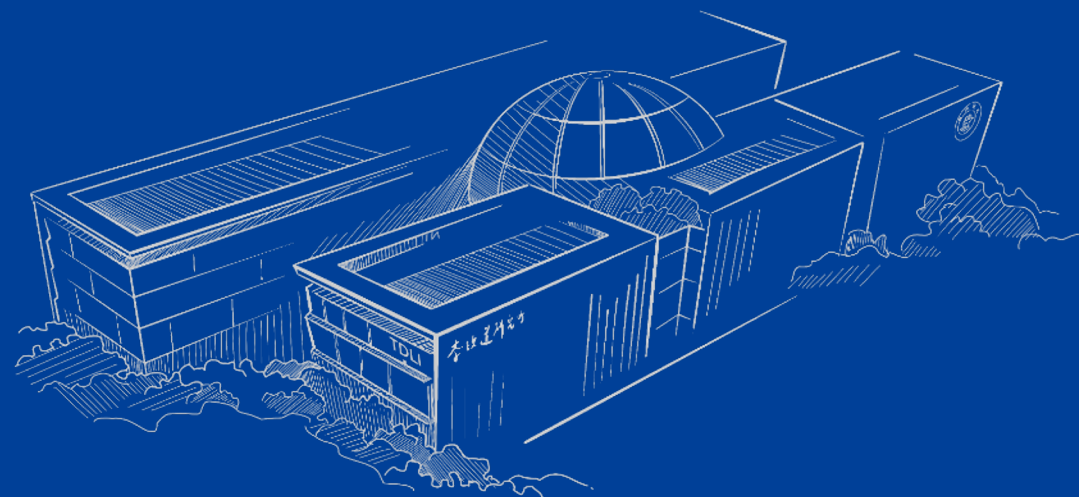
Summary

- The DarkSHINE: a fixed-target experiment searching for dark matter.
- **The DarkSHINE will be almost background free experiment**
 - Expected 0.02 background in 3×10^{14} electron-on-target (w.r.t 1 year running)
 - Above 50% dark photon signal acceptance efficiency
- **The DarkSHINE has competitive sensitivity** ([Sci. China-Pay. Mech. Astron., 66\(1\):211062 \(2023\)](#))
 - Sensitive to most of phase space predicted by models with 3 years running
- **Detector key technology R&D studies are going on** ([arXiv:2407.20723](#) , [Nucl. Sci. Tech. 35, 148 \(2024\)](#) , [Nucl Sci Tech 35, 201 \(2024\)](#))
- With more physics opportunities ahead, please stay tuned!





谢 谢



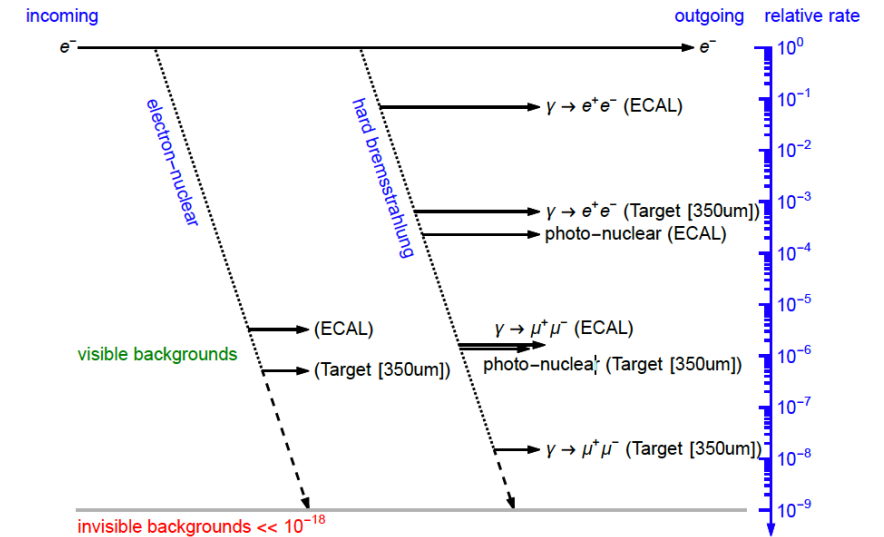
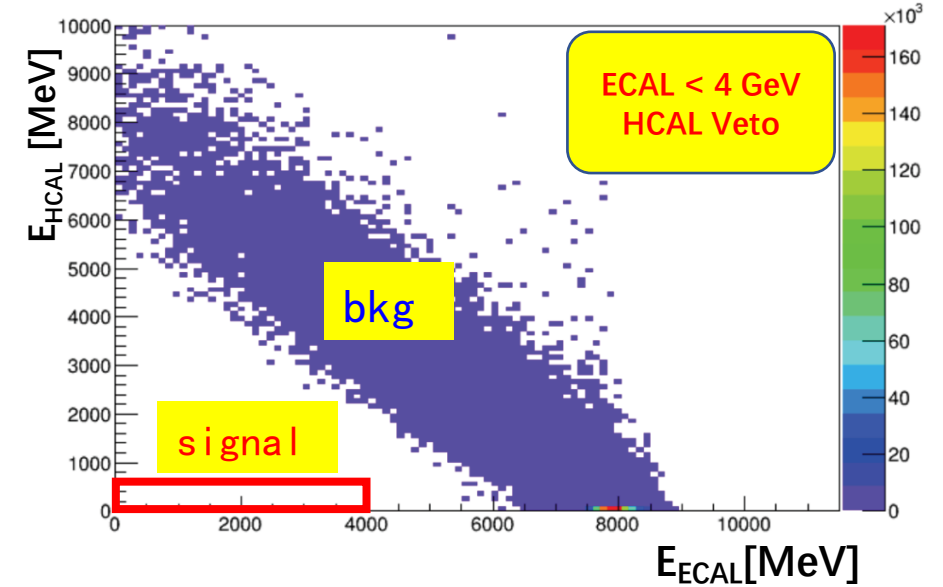
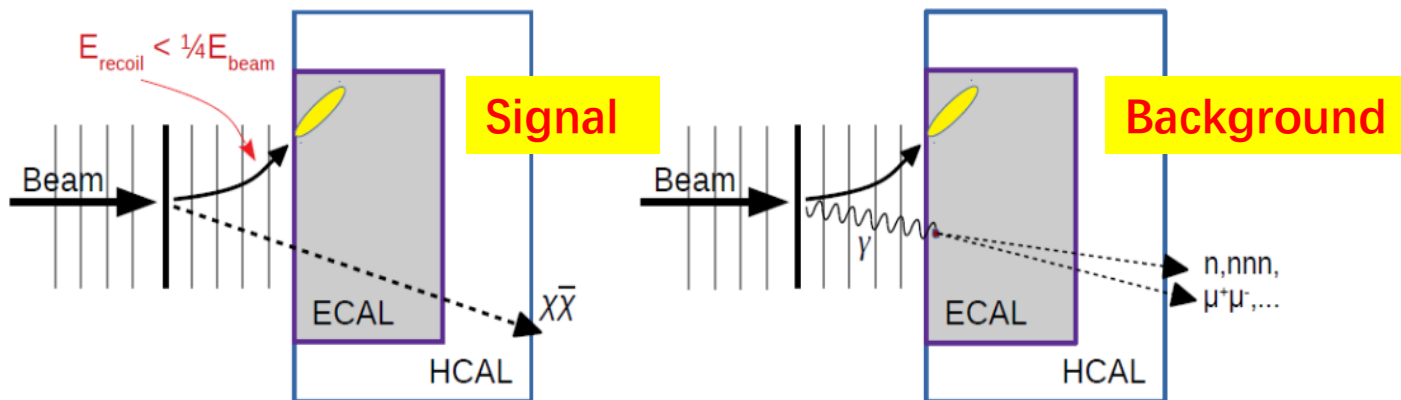
Experimental Approaches

- **High repetition rate single electron beam**

- More striking record of single electron-on-target event energy loss after recoiling
- Requirements: fast detector response and readout electronics, radiation hardness to allow high e-on-target statistics

- **Energy + Momentum loss detection**

- Synergy of high precision tracking and calorimetry



Simulated background statistics:

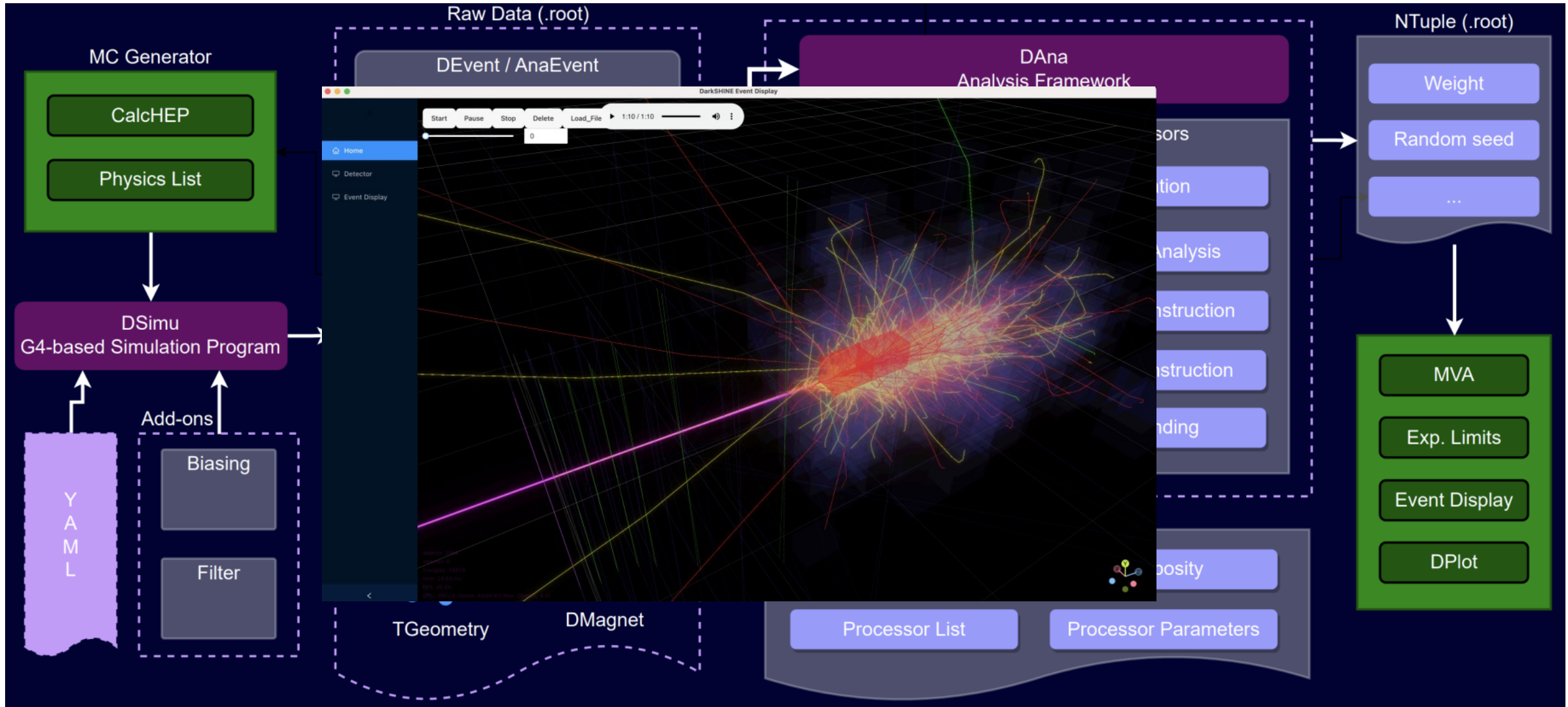
Process	Generate events	Branching ratio	EOTs
Inclusive	2.5×10^9	1.0	2.5×10^9
Bremsstrahlung	1×10^7	6.70×10^{-2}	1.5×10^8
GMM_target	1×10^7	$1.5(\pm 0.5) \times 10^{-8}$	4.3×10^{14}
GMM_ECAL	1×10^7	$1.63(\pm 0.06) \times 10^{-6}$	6.0×10^{12}
PN_target	1×10^7	$1.37(\pm 0.05) \times 10^{-6}$	4.0×10^{12}
PN_ECAL	1×10^8	$2.31(\pm 0.01) \times 10^{-4}$	4.4×10^{11}
EN_target	1×10^8	$5.1(\pm 0.3) \times 10^{-7}$	1.6×10^{12}
EN_ECAL	1×10^7	$3.25(\pm 0.08) \times 10^{-6}$	1.8×10^{12}

Event cut-flow of each background process:

Table 4 Event cut flow for each background sample in Table 2. The selection efficiencies of each cut are listed in the table (%)

	EN_ECAL	PN_ECAL	GMM_ECAL	EN_target	PN_target	GMM_target	Hard_brem	Inclusive
Total events	100	100	100	100	100	100	100	100
Only 1 track	58.87	70.48	87.36	5.85	5.88	$< 10^{-3}$	78.73	84.40
$p_{\text{tag}} - p_{\text{rec}} > 4 \text{ GeV}$	0.0044	0.0033	0.0041	5.58	5.46	$< 10^{-5}$	70.49	4.80
$E_{\text{HCAL}}^{\text{total}} < 100 \text{ MeV}$	$< 10^{-3}$	$< 10^{-3}$	0	0.30	0.72	0	69.61	4.76
$E_{\text{HCAL}}^{\text{MaxCell}} < 10 \text{ MeV}$	$< 10^{-3}$	$< 10^{-3}$	0	0.13	0.27	0	65.00	4.48
$E_{\text{HCAL}}^{\text{MaxCell}} < 2 \text{ MeV}$	$< 10^{-3}$	$< 10^{-3}$	0	0.058	0.095	0	58.14	4.04
$E_{\text{ECAL}}^{\text{total}} < 2.5 \text{ GeV}$	0	0	0	0	0	0	0	0

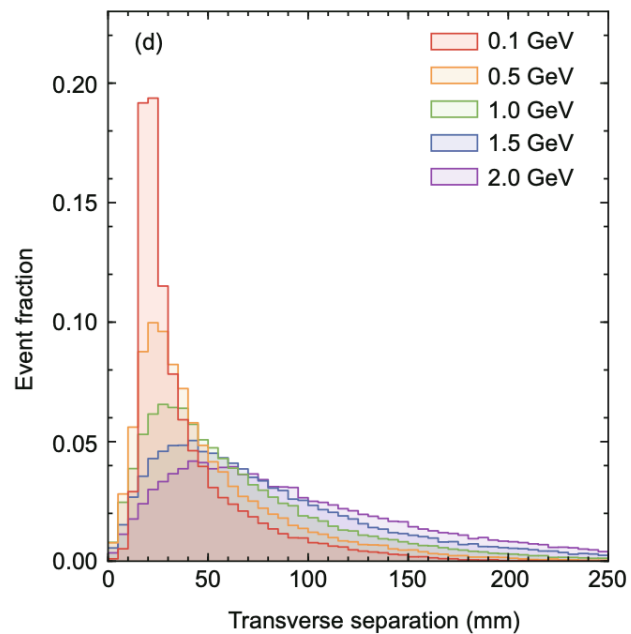
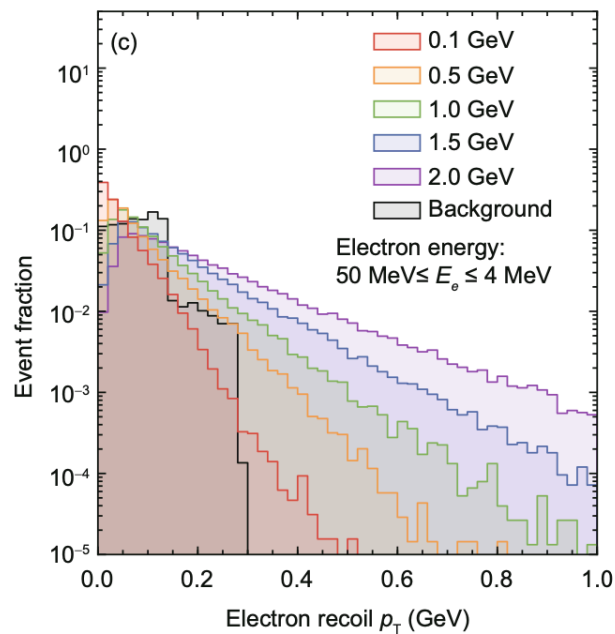
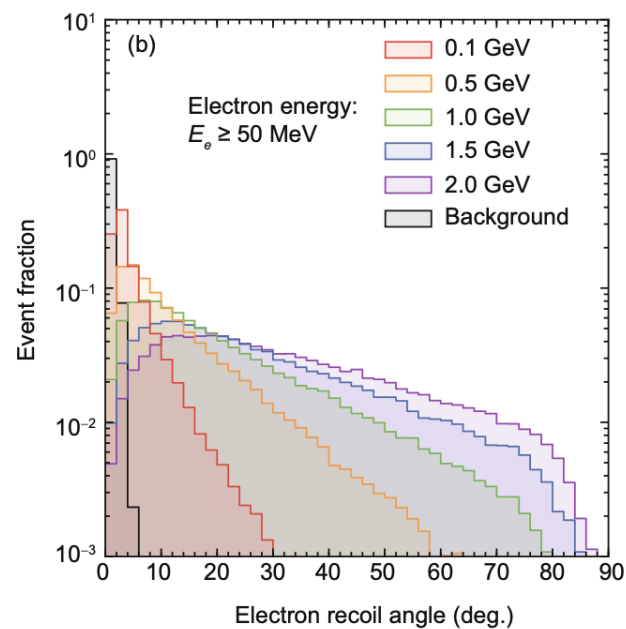
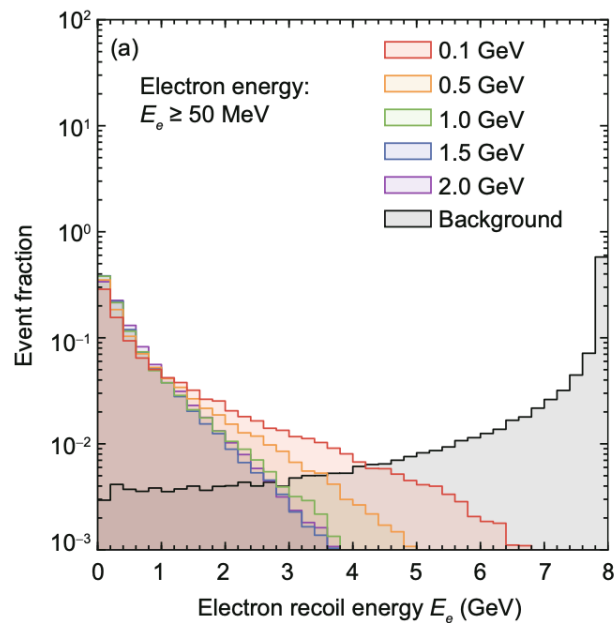
Software Framework Optimization



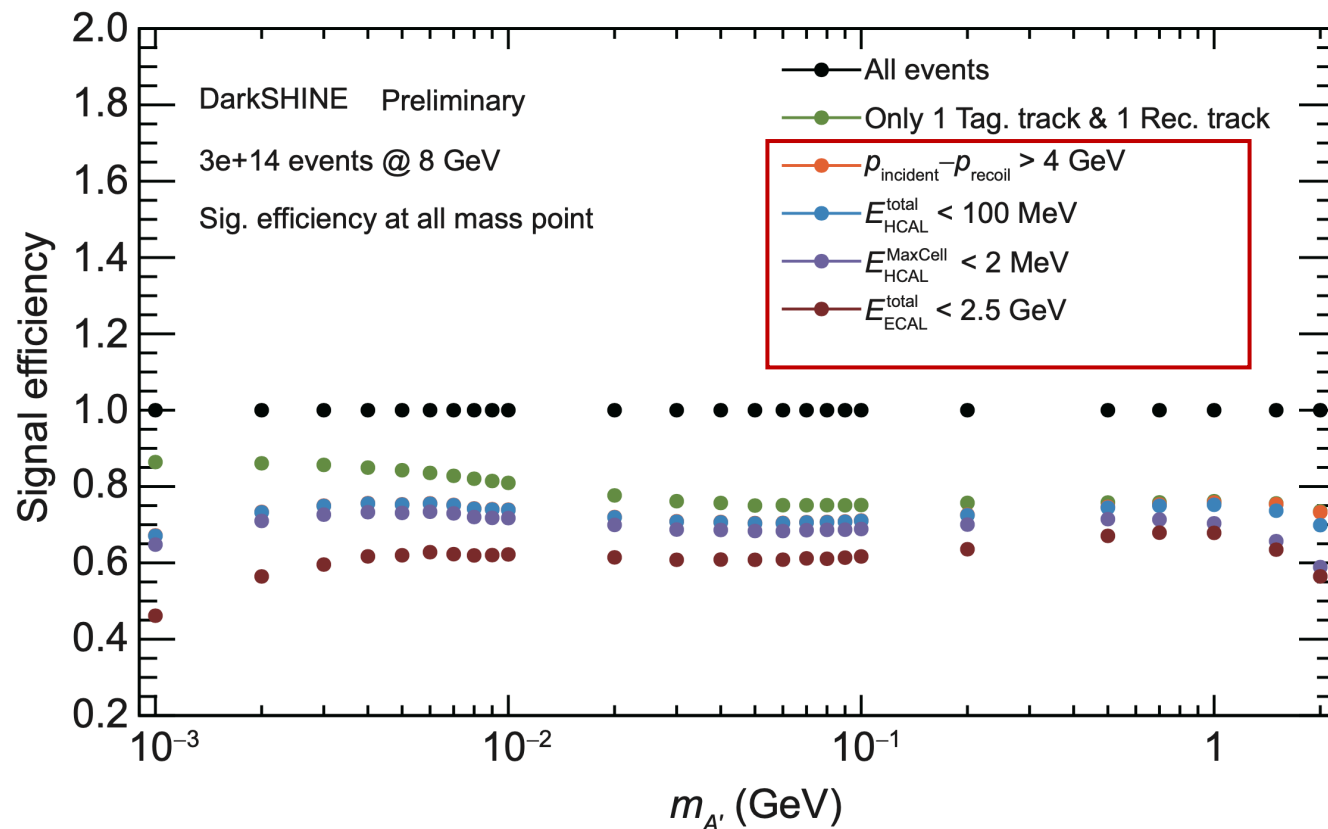
How DM Shining at the DarkSHINE Experiment?



Other Kinematic Distributions



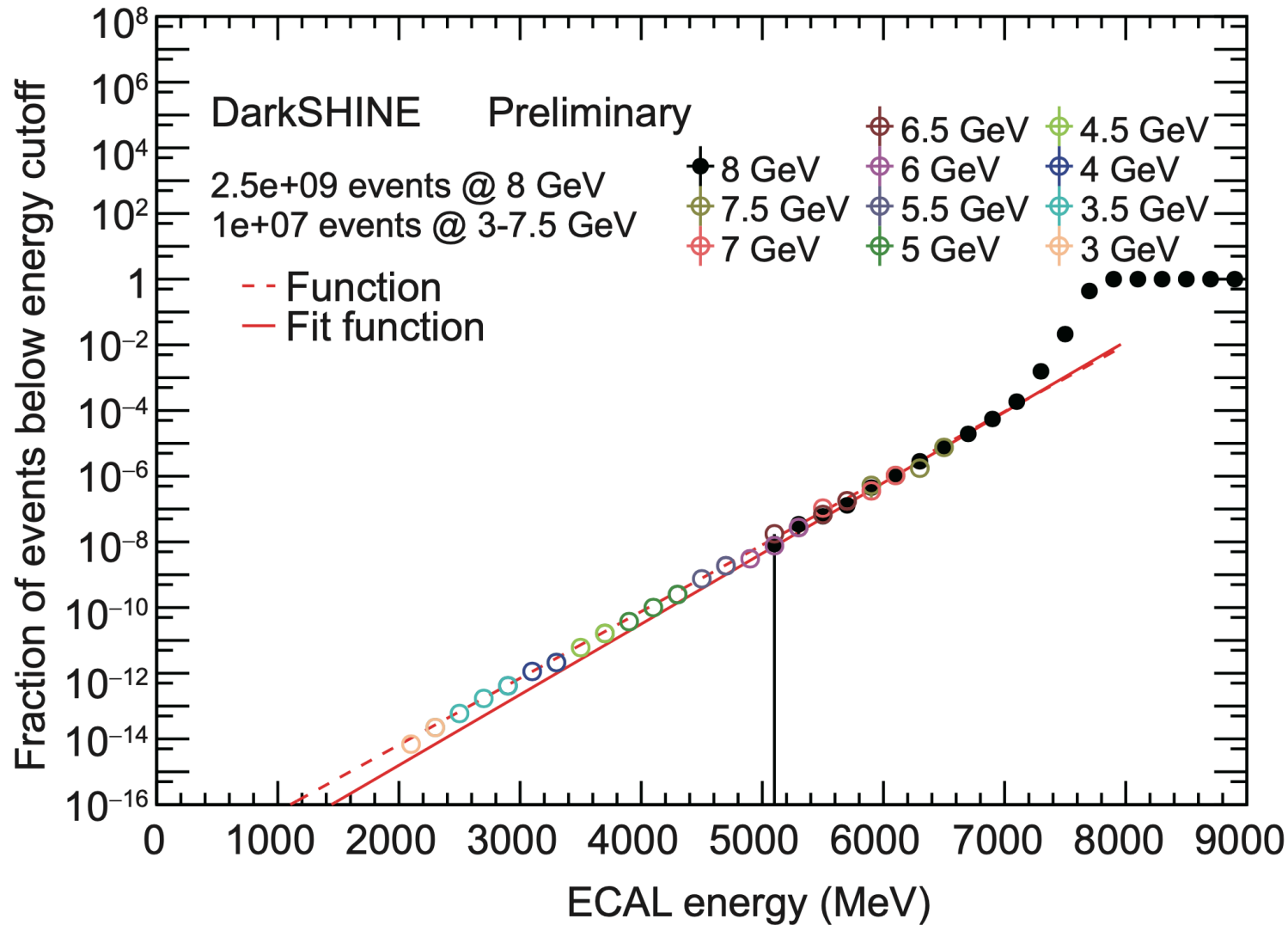
Acceptance in the Signal Region



Signal Region selections

- 60% signal events survive the cut-flow, no background survive (2.5e9)
- Acceptance efficiency drops in:
 - **Low-mass** region of a few MeV: tight energy cuts.
 - **High-mass** region above 1 GeV: particles with large incident/recoil angle go into the HCAL directly.

Background Estimation

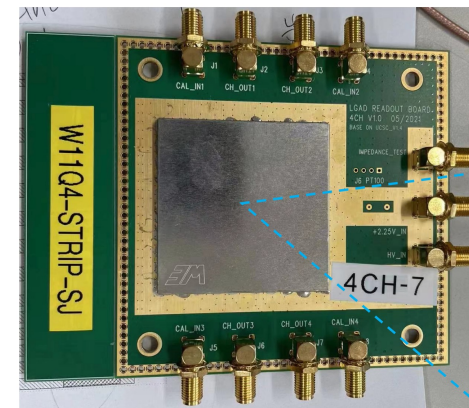
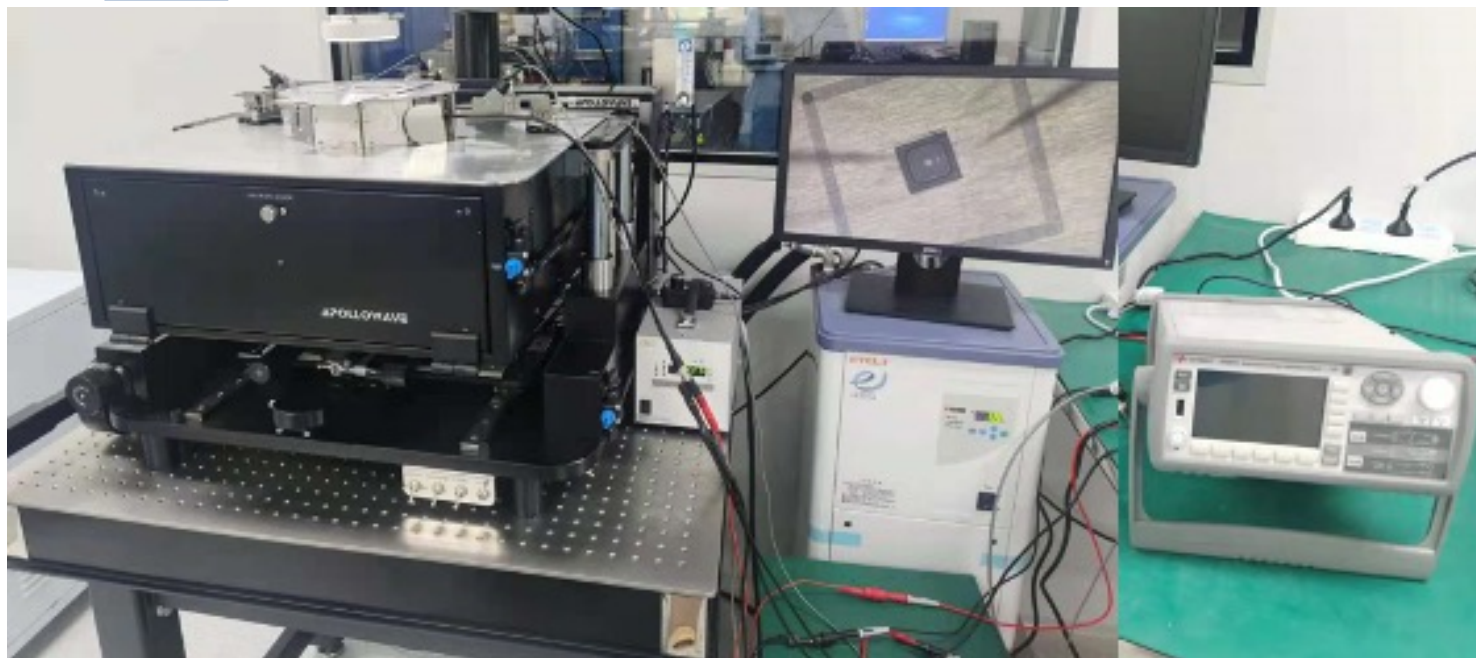


Expected background yields go down quickly at lower ECAL energy.

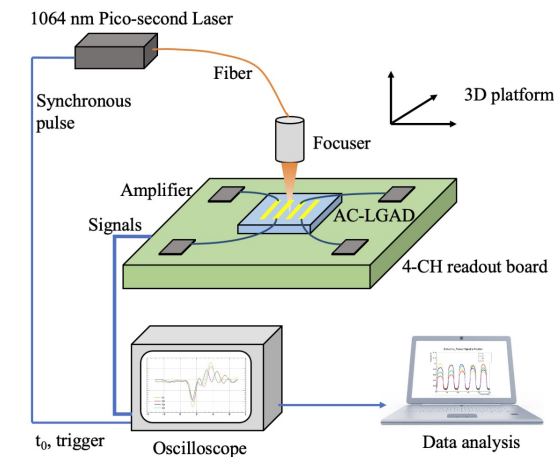
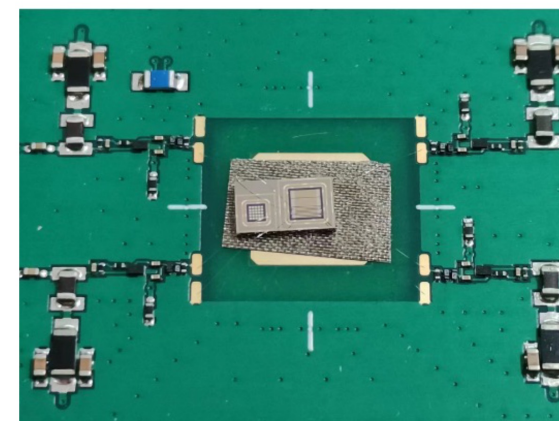
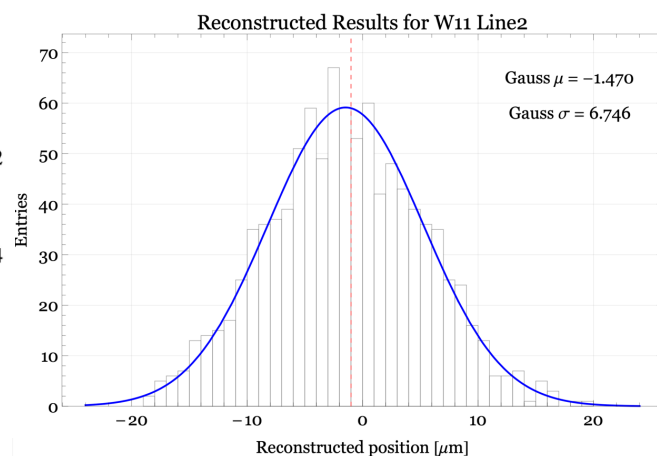
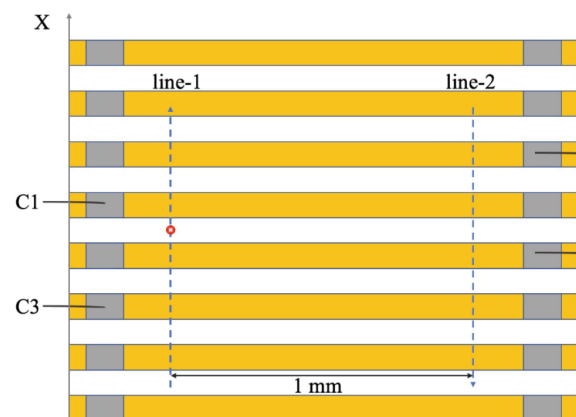
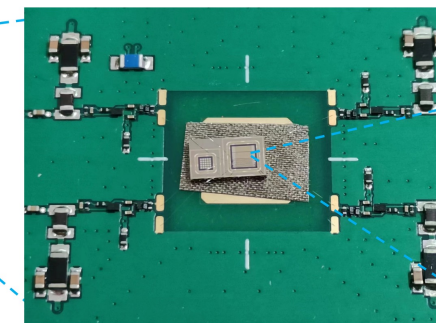
In order to estimate background yields in 10^{14} EOT, extrapolation method is used

- ▶ fit from inclusive background process
- ▶ extrapolation from low energy samples

Hardware R&D: Tracking



Working point
W11: 350V
W12: 150V



Hardware R&D: ECAL

- Experiments based on LYSO and SiPM

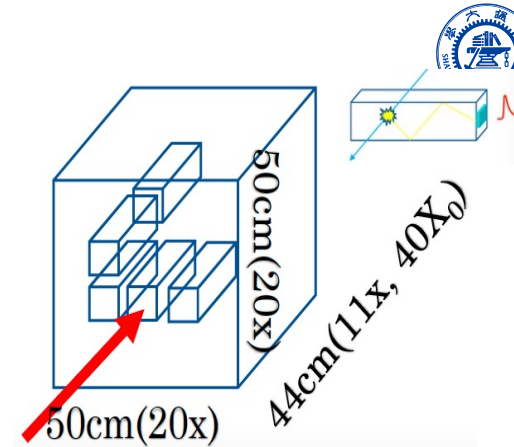
a. LYSO intrinsic radiation from $^{176}_{71}\text{Lu} \rightarrow ^{176}_{72}\text{Hf}$

c. Cosmic ray test

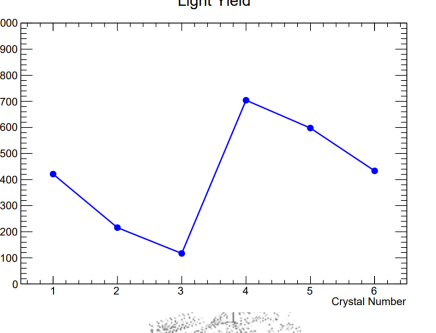
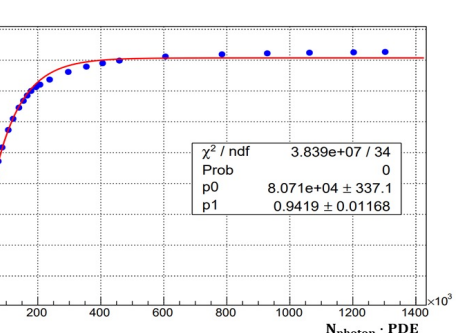
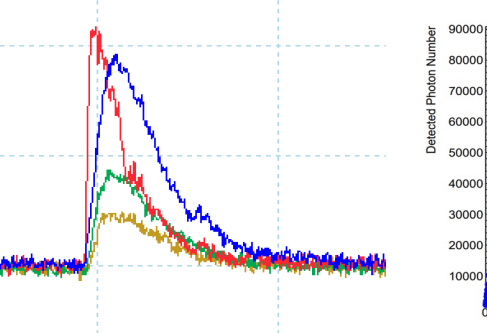
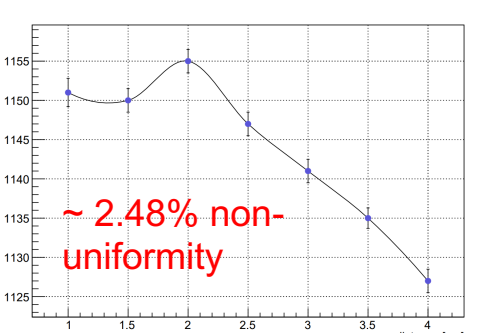
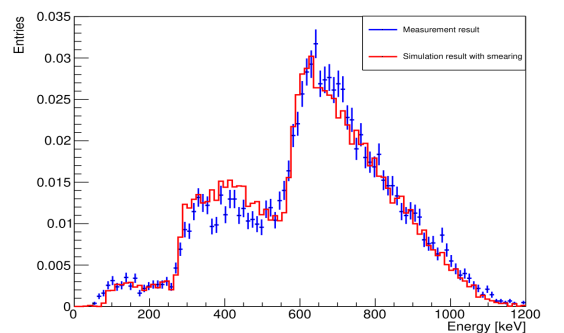
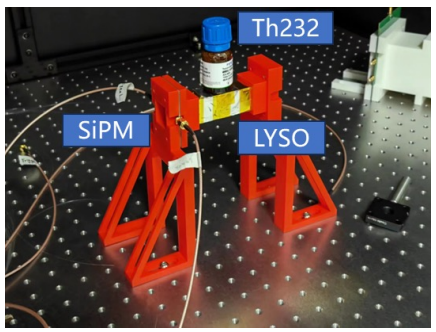
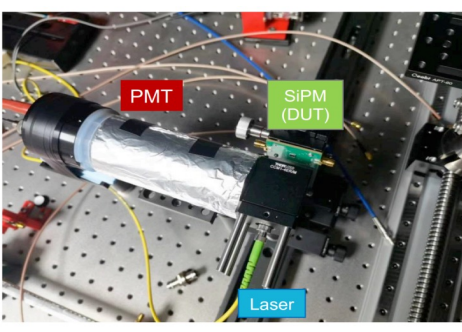
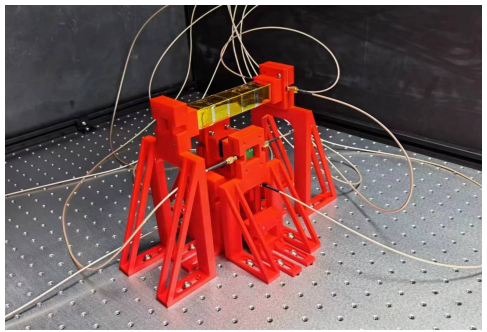
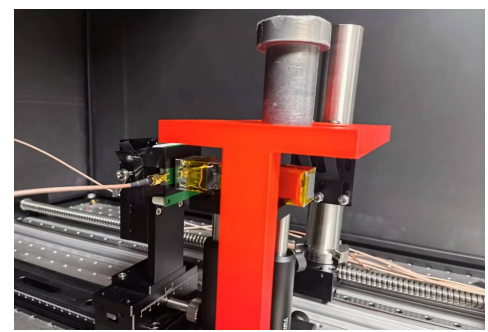
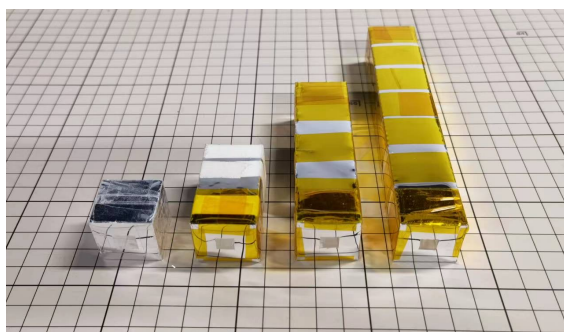
e. Light yield changed with crystal size

b. Uniformity test with $^{60}_{27}\text{Co}$ source

d. SiPM dynamic range test



Baseline design of each crystal: X,Y = 2.5 cm, Z = 4 cm (radiation length: 1.14 cm)



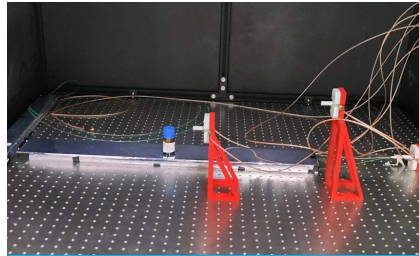
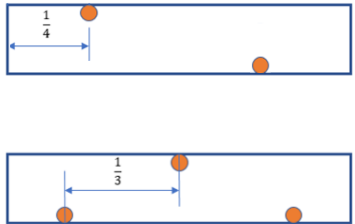
Hardware R&D: HCAL

- **Scintillator test in TDLI lab**

- SiPMs performance are studied first, both size, gain and noise are considered, and picked one type (Hamamatsu S13360-3050, gain $1.7e6$) for the rest tests

- **Radioactive source test for uniformity**

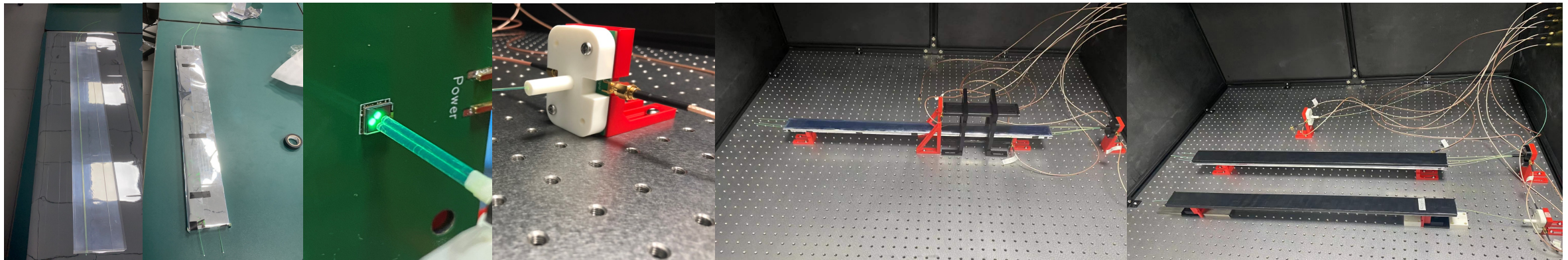
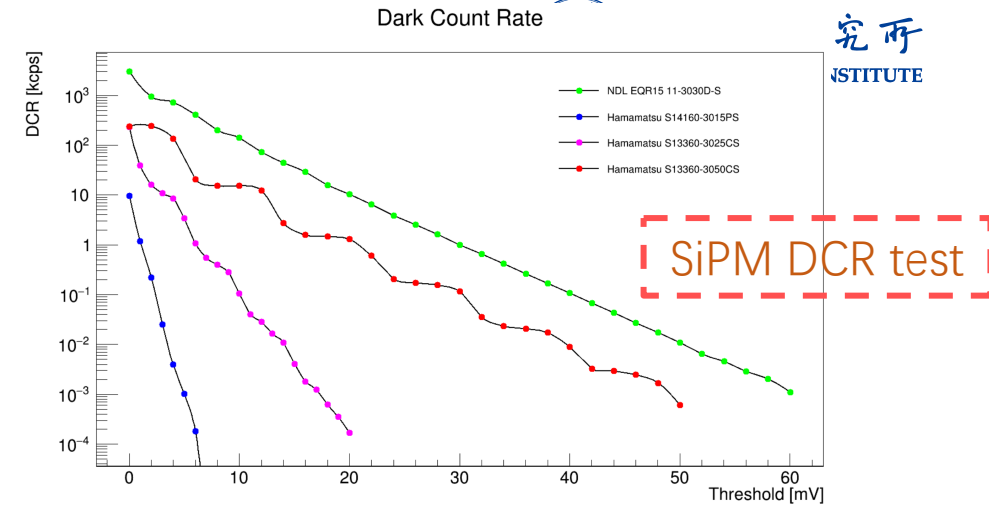
- Very good uniformity of the scintillator along the 75 cm side



Radioactive source test

- **Cosmic ray test for photon yields**

- Various types of scintillator are tested : sizes (75cm×5cm/10cm×1cm/2cm), number of fiber grooves/used, manufacturer/composition are tested
- Noise test: to decide the minimal cut we could use in analysis



Scintillator (w/ & w/o wrap)

WLS & SiPM

Fiber collimation

Cosmic ray test: photon yields test (left) and noise test (right)

More Physical Motivations

- **Dark SHINE could explore a vast array of sub-GeV physics with unique sensitivity**
 - New force carriers coupling to electrons, decaying visibly or invisibly
 - Quasi-thermal DM, e.g. asymmetric DM or ELDER DM
 - New long-lived resonances produced in the dark sector (SIMP)
 - Freeze-in models with heavy mediators
 - Axion-like particles (ALP)
 - Milli-charged dark sector particles
 - Probe neutrino-nuclear interactions (νN) via electro-nuclear (eN) measurements and constrain nuclear models
 - ...